**KAON2019 - XI International Conference on Kaon Physics** 

# Dark Sector studies at Belle II: First results and prospects

University of Perugia, Perugia, September 10-13<sup>th</sup>, 2019

Luigi Corona ~ INFN and University of Pisa

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• on behalf of the Belle II collaboration



Outline

SuperKEKB and the Belle
 II experiment

# Dark Sector with Belle II

Conclusions

# SuperKEKB collider

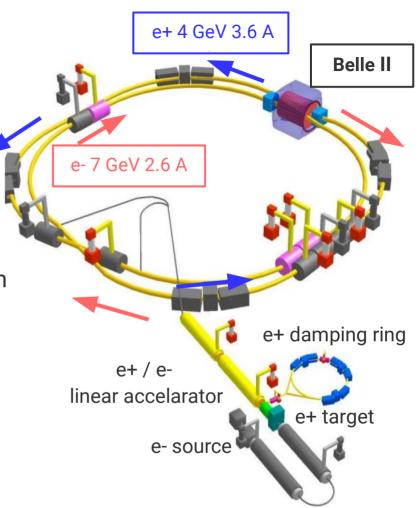
**KEKB** 

- Second generation of B-factory, successor of KEKB (Tsukuba, Japan)
- $e^+/e^-$  asymmetric collider E\* ~10.58 GeV ( =  $M_{\gamma(4S)}$ )
- highest world luminosity → Nano-beam scheme
  - Squeeze beams → increase probability of interaction
  - Higher currents and smaller beams than KEKB

SuperKEKB

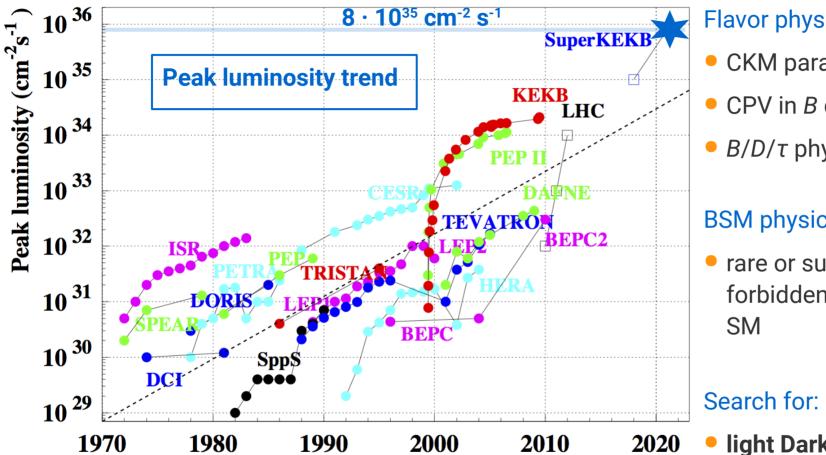


I(A): ~ 1.6/1.2 I(A): ~ 3.6/2.6 ( $\uparrow$  x 2) β\*<sub>y</sub>(mm)~ 5.9/5.9 β\*<sub>y</sub>(mm)~ 0.27/0.3 ( $\downarrow$  x 1/20)



x40 higher luminosity than KEKB

# Physics program



#### Flavor physics and SM Test:

- **CKM** parameters
- CPV in B decays
- $B/D/\tau$  physics

#### **BSM physics**:

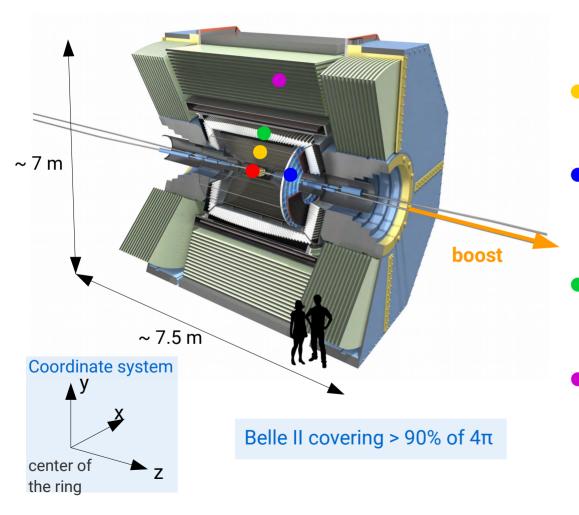
rare or suppressed or forbidden processes in the

light Dark Sector

#### x40 higher luminosity than KEKB

#### **Belle II detector**

► See Pavel Krokovny talk: 13/09, 9:00!



#### VerteX Detector (VXD):

2 layer DEPFET pixel detectors (PXD)\*\*\* 4 layer DSSD silicon vertex detectors (SVD)

**Central Drift Chamber (CDC):** He $(50\%)C_2H_6(50\%)$ 

Particle Identification:
 Time-of-Propagation (TOP) (barrel)
 Aerogel Ring Cherenkov (ARICH) (FWD)

E.M. calorimeter (ECL):
 Csl(Tl) crystals

#### $K_L$ and $\mu$ detector (KLM):

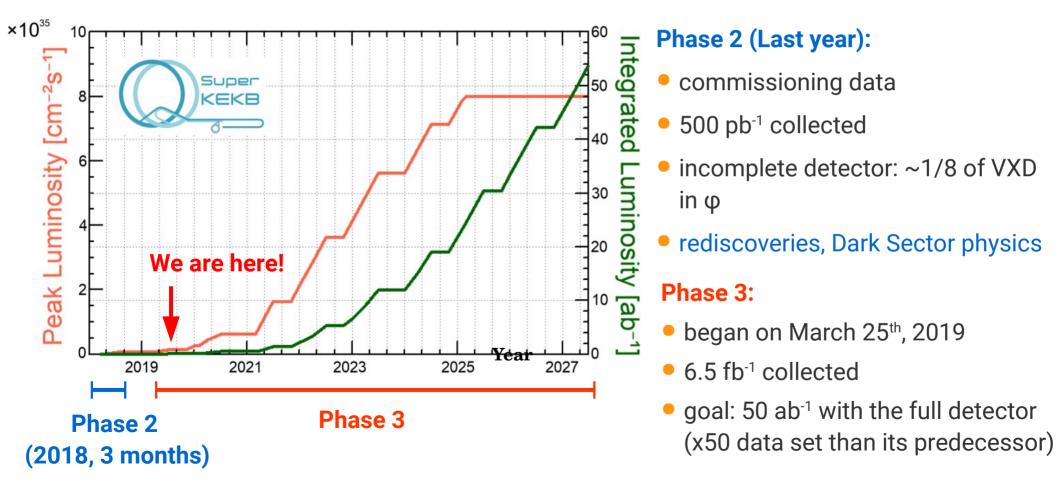
Resistive Plate Chambers (RPC) (outer barrel) Scintillators + SiPM (endcaps, inner barrel)

\*\*\*second layer not complete

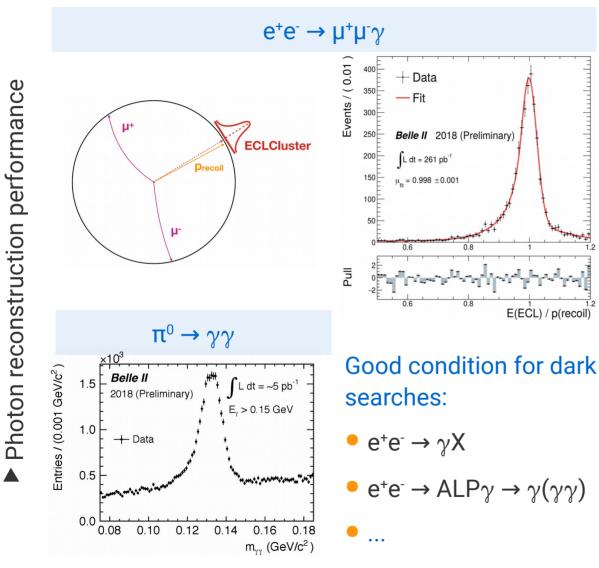
# SuperKEKB schedule

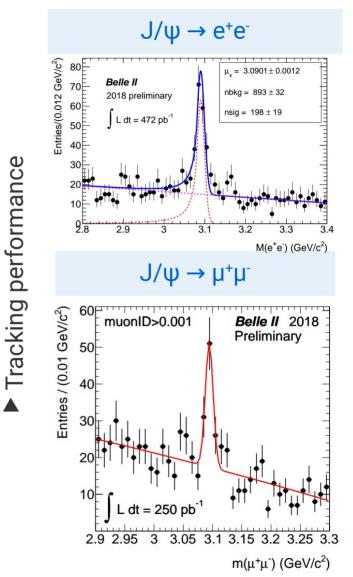
#### Phase 1 (2016):

dedicated to machine commissioning, w/o detector



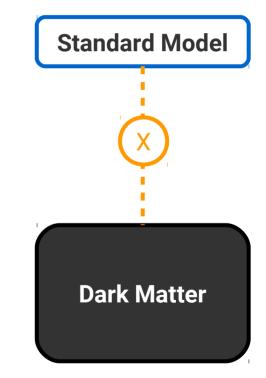
# Highlights from Phase 2





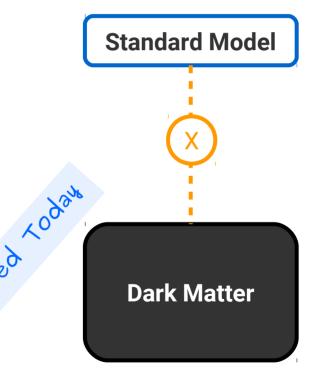
#### **Dark Sector Physics**

- Many astrophysical observations provide evidence for the existence of some kind of matter that interacts mostly gravitationally with the Standard Model (SM) particles: Dark Matter.
- Possible GeV and sub-GeV theoretical scenarios: Light-DM weakly coupled with SM through a light Dark Sector mediator X
- Different possible portals between Dark Matter and SM depending on the mediator X:
  - Vector Portal  $\rightarrow$  Dark Photon A', Dark Z'
  - Pseudo-scalar Portal → Axion Like Particles (ALPs, axion)
  - Scalar Portal → Dark Higgs / Dark Scalar
  - ▶ **Neutrino Portal** → Sterile Neutrinos



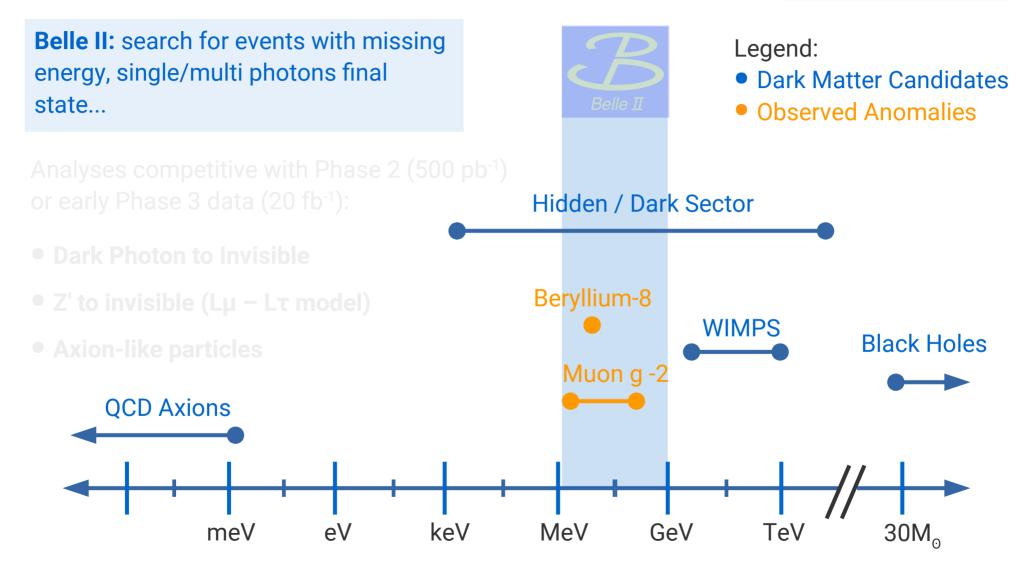
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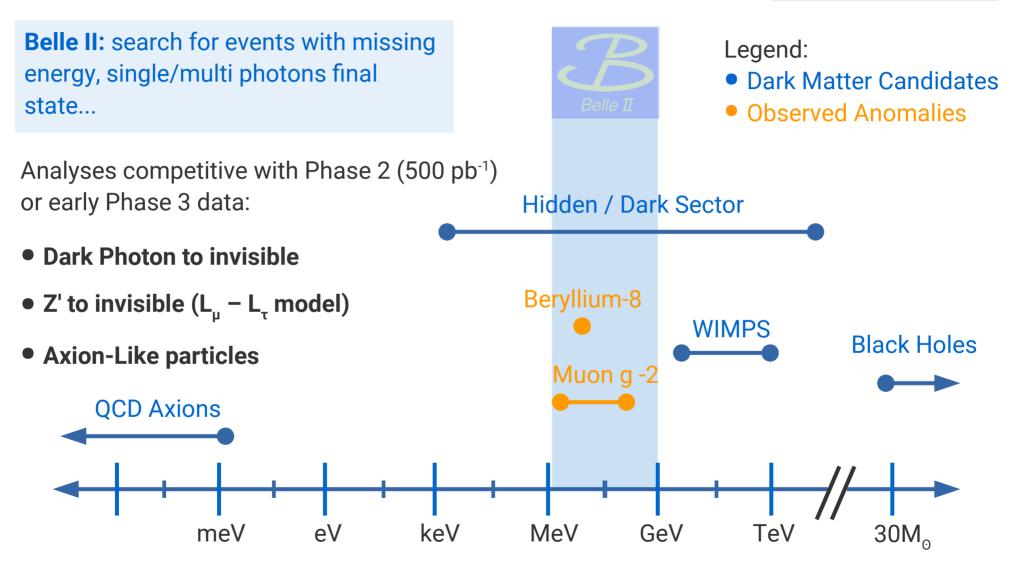


#### Search for Dark Matter

Ref. arXiv:1707.04591



#### Search for Dark Matter



# Dark photons A'

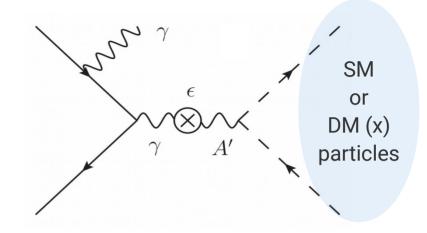
- Extension of the SM:  $\rightarrow$  additional U(1)' symmetry
- Massive dark photon that mixes with the photon with strenght ε

$$e\varepsilon J^{\mu}_{SM}A'_{\mu}$$

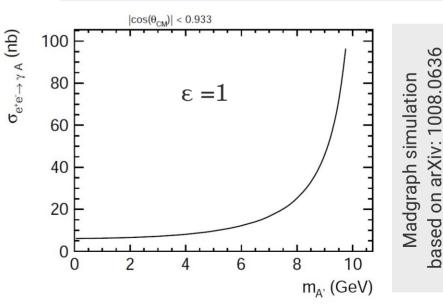
- $m_x > 0.5 m_{A'} \rightarrow A'$  visible decays to SM particles
- $m_x < 0.5 m_{A'} \rightarrow A'$  invisible decays to Light-DM particles

References: P. Fayet, Phys. Lett. B 95, 285 (1980), P. Fayet Nucl. Phys. B 187, 184 (1981) B. Holdom, Phys. Lett. B 166, 196 (1986)

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#### Production cross section

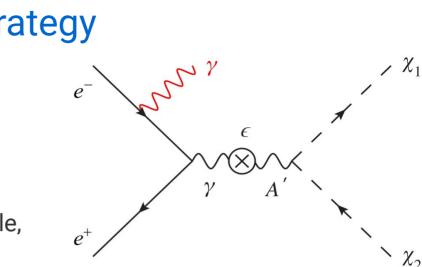


#### Invisible dark photon searching strategy

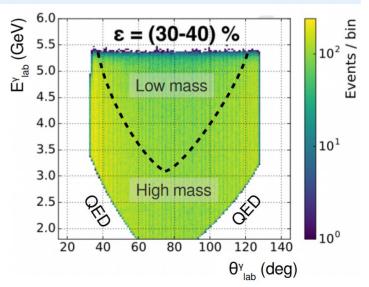
- Signal signature:
  - Single high-energetic ISR photon final state
  - Needs a single photon trigger (not available in Belle,
     ≈ 10% of data in BaBar)
  - Look for a bump in the variable:

$$E_{\gamma} = \frac{s - M_{A'}^2}{2\sqrt{s}}$$

• Selection criteria  $\rightarrow E_{\gamma}^{LAB}$  vs  $\theta_{\gamma}^{LAB}$ 

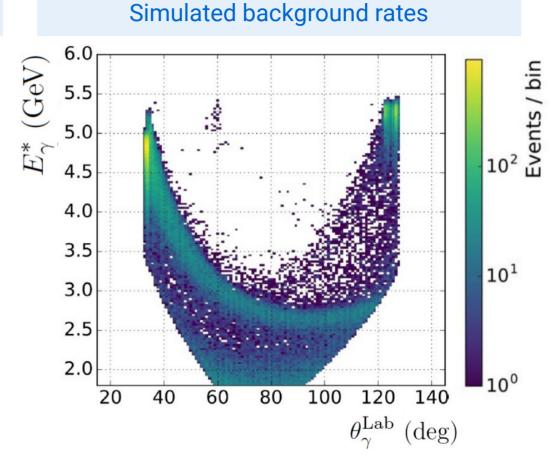


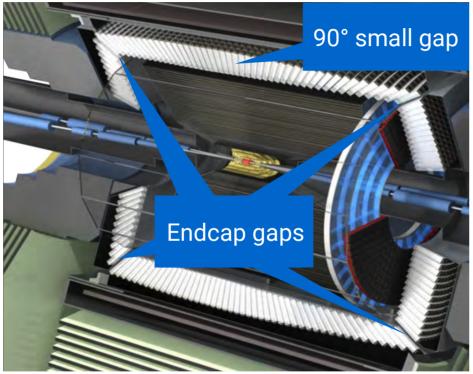
#### Simulated signal efficiency



# Invisible dark photon backgrounds

#### Sources of ECL inefficiency





Main Backgrounds events:

- ►  $e^+e^- \rightarrow e^+e^-\gamma$
- ►  $e^+e^- \rightarrow \gamma\gamma(\gamma)$

# Invisible dark photon backgrounds

#### Sources of ECL inefficiency



900

6.0

5.5

5.0

4.5

4.0

3.5

3.0

2.5

2.0

20

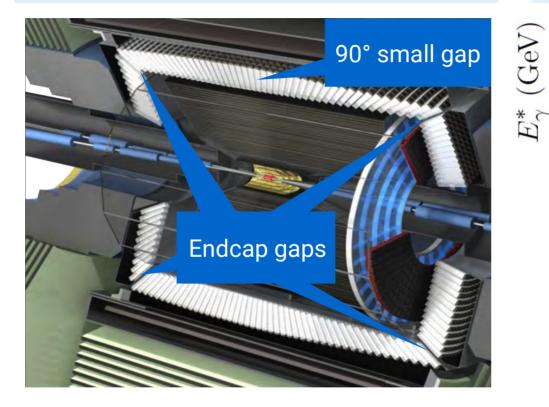
40

 $e+e- \rightarrow e+e-\gamma$ 

60

leptons out of acceptance

endo



• Main Backgrounds events:

- ►  $e^+e^- \rightarrow e^+e^-\gamma$
- ►  $e^+e^- \rightarrow \gamma\gamma(\gamma)$

Peaking  $e^+e^- \rightarrow \gamma \gamma(\gamma)$  dominates the analysis

80

100

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0 c Events / bin

10<sup>1</sup>

100

endcap

gap

140

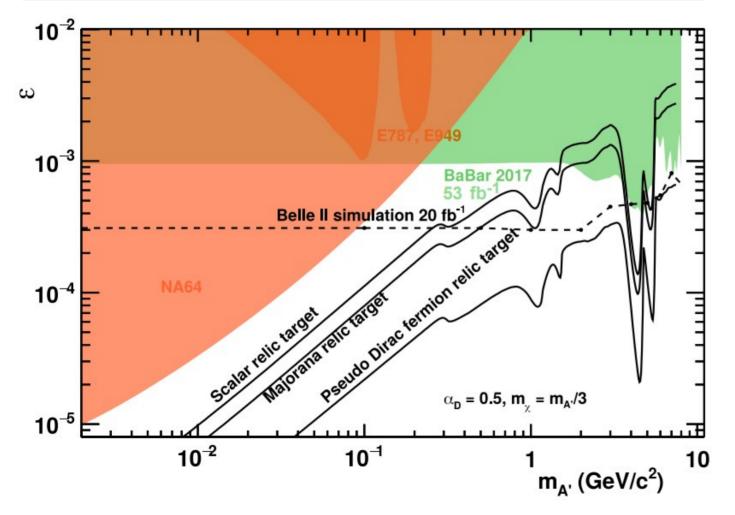
(deg)

120

 $\theta^{\text{Lab}}$ 

#### Projected upper limits on $\varepsilon$ , invisible dark photon

 $e^+e^- \rightarrow \gamma A', A' \rightarrow invis.$ 



BaBar 2017, 53 fb<sup>-1</sup> arXiv:1702.03327

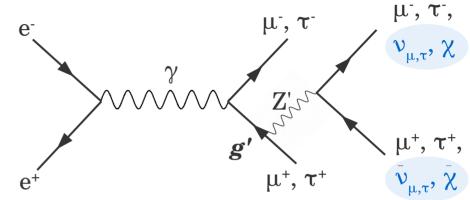
Belle II projection at
20 fb<sup>-1</sup> arXiv:1808.10567

- Significantly better than BaBar (53 fb<sup>-1</sup>) due to:
  - better hermeticity of the ECL
  - better efficiency of the KLM

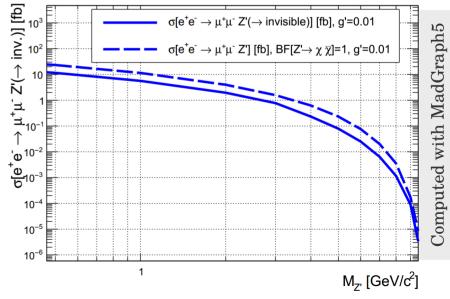
# Z' to invisible: theory

- Extension of the SM: additional U(1)<sub> $L\mu L\tau$ </sub> symmetry
  - $\mathcal{L} = \sum_{\ell} \theta g' \bar{\ell} \gamma^{\mu} Z'_{\mu} \ell$
- Introduces a light vector boson Z' with coupling g' only with the 2<sup>nd</sup> and 3<sup>rd</sup> generation of leptons
- If kinematically accessible, Z' could decay to DM
- May explain:  $(g-2)_{\mu}$ , abundance of DM in the Universe
- Invisible decay channel to be explored for the first time: e<sup>+</sup>e<sup>-</sup> → µ<sup>+</sup>µ<sup>-</sup> + missing energy

Shuve et al. (2014), arXiv:1403.2727 Altmannshofer et al. (2017) arXiv:1609.04026



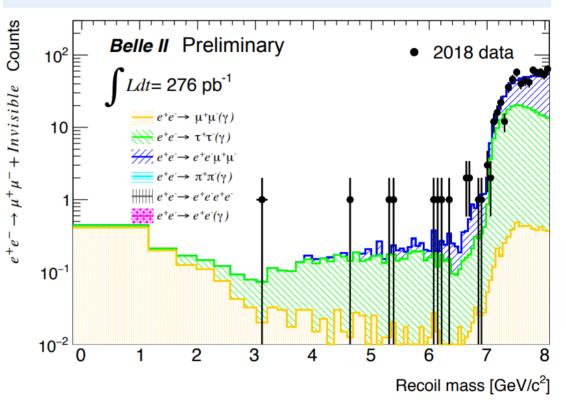
#### Production cross section



# Z' to invisible search

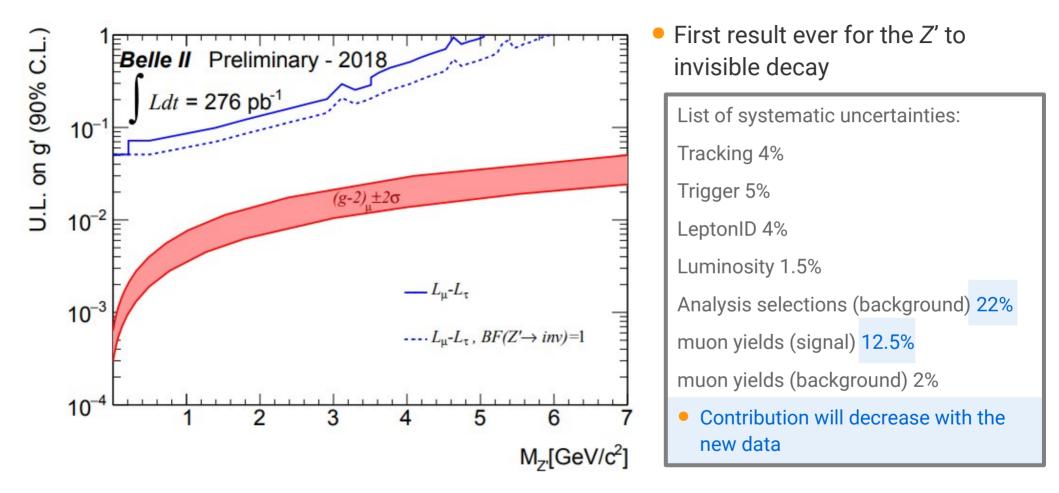
- Signal signature:
  - look for bumps in the recoil mass against a µ<sup>+</sup>µ<sup>-</sup> pair
  - nothing in the rest of event
- Main background events from QED processes:
  - ▶  $e^+e^- \rightarrow \mu^+\mu^-(\gamma)$
  - $e^+e^- \rightarrow \tau^+\tau^-(\gamma)$ ,  $(\tau \rightarrow \mu vv)$
  - ►  $e^+e^- \rightarrow e^+e^-\mu^+\mu^-$
- Only 276 pb<sup>-1</sup> used due to trigger conditions for two tracks events

#### **Recoil Mass distribution**



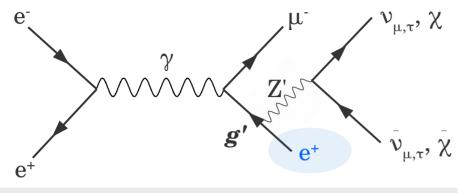
\*No sensitivity to the parameter space region for  $\rm M_{Z'}$  > 8 GeV/c^2

# g' upper limits



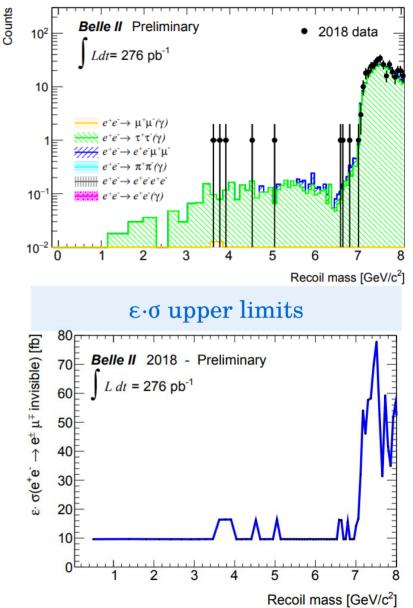
\*If Dark Matter is kinematically accessible, it can be assumed  $BR(Z' \rightarrow inv.) = 1$ 

#### Lepton Flavor Violating Z'



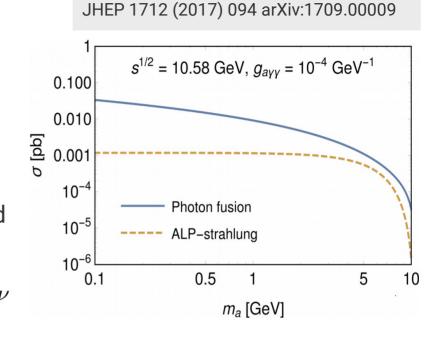
I. Galon et al.: arXiv:1610.08060, arXiv:1701.08767

- Searching for a lepton flavor violating Z' that couples to eµ: e<sup>+</sup>e<sup>-</sup> → e<sup>+</sup>µ<sup>-</sup> Z' (Z' → inv.)
- Expected low background from SM processes
- No working model to test → Model independent approach
- Same selection criteria from flavor-conserving Z' to invisible search used



# Axion-like particles (ALPs)

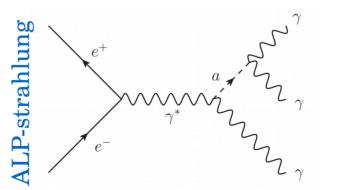
- Pseudo-scalar particles which couple to SM bosons
- No relation between mass and coupling  $\rightarrow$  different from QCD axions
- $m_a < MeV \rightarrow excellent DM candidates$
- $m_a \sim \text{GeV} \rightarrow \text{mediator of interaction between SM and}$ vet undiscovered DM particle
- Focus on coupling to photons  $\rightarrow -\frac{g_{a\gamma\gamma}}{4}aF_{\mu\nu}\tilde{F}^{\mu\nu}$

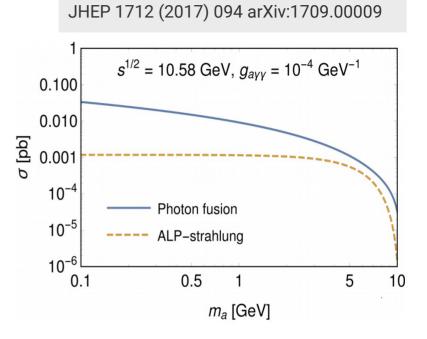


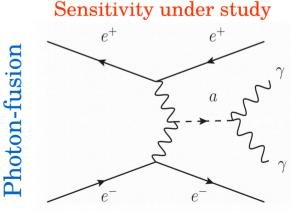
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- No relation between mass and coupling → different from QCD axions
- $m_a < MeV \rightarrow excellent DM candidates$
- m<sub>a</sub> ~ GeV → mediator of interaction between SM and yet undiscovered DM particle
- Focus on coupling to photons  $\rightarrow -\frac{g_{a\gamma\gamma}}{4}aF_{\mu\nu}\tilde{F}^{\mu\nu}$
- Two production processes possible:
  - Focus on ALP-strahlung
    - $e^+e^- \rightarrow \gamma + inv$ .
    - $e^+e^- \rightarrow 3\gamma$









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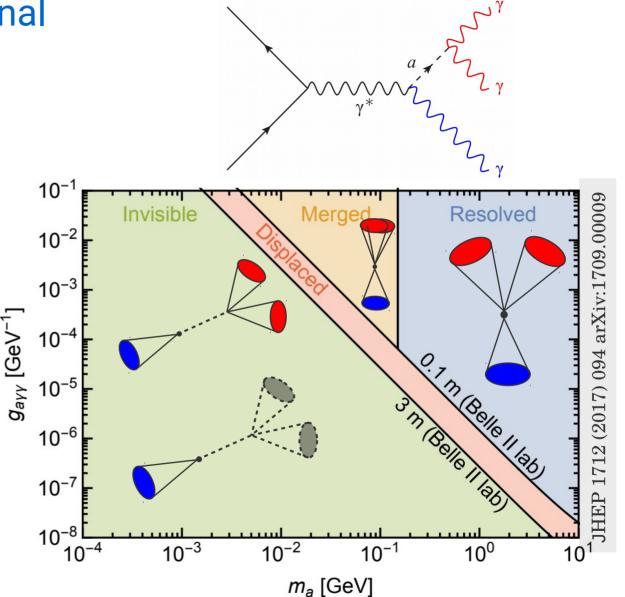
# Axion-like particles: signal

Parameters  $(m_a, g_{a\gamma\gamma})$  determine the displacement and the  $\theta$  angle between the  $2\gamma$ 

• ALP lifetime:  $\tau \sim 1/m_a^2 g_{a\nu\nu}^2$ 

Four Signatures:

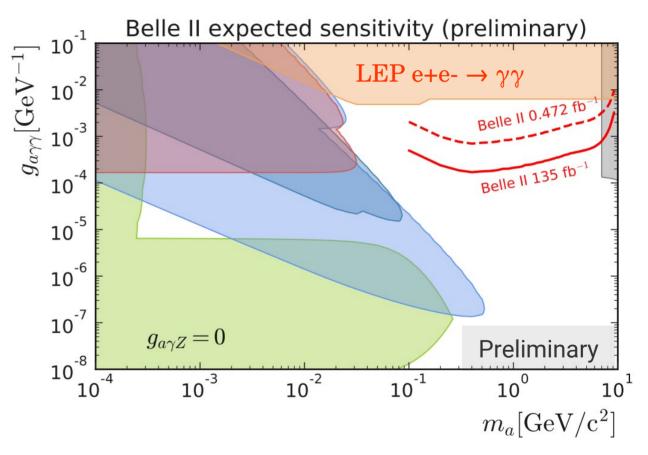
- Resolved: prompt decay, large  $\theta$
- Merged: prompt decay, small  $\theta$
- Invisible: a decays outside the detector or a decays to invisible particles, as DM particles
- Displaced: veto this region (indistinguishable from  $e^+e^- \rightarrow \gamma \gamma$ )



18

# Axion-like particles: sensitivity

- Focus on:  $e^+e^- \rightarrow \gamma a$ ,  $a \rightarrow \gamma \gamma$
- Observable: Invariant mass of the two photons
- Main SM background:
  - ►  $e^+e^- \rightarrow \gamma\gamma(\gamma)$
  - ►  $e^+e^- \rightarrow e^+e^-(γ)$
  - $e^+e^- \rightarrow P\gamma(\gamma), P = (\pi^0, \eta, \eta')$
- Belle II expected limits
  - No systematics included
  - beam background assumed to be negligible



#### Conclusions

- The Belle II experiment, designed mainly for B-physics, has a broad and active program to explore the Dark Sector physics
- 2018: successful SuperKEKB commissioning and collected ~0.5 fb<sup>-1</sup> of data → b and charm physics rediscoveries, but also search for Dark Sector
- 2019: phase 3 started this year on March  $25^{th} \rightarrow up$  to now 6.5 fb<sup>-1</sup> collected
- Many searches are ongoing, A', Z', ALPs, and there will be the possibility to explore many more Dark Sector models
  - ► **A'** → *inv*., expected sensitivity:  $\varepsilon \sim 2 \cdot 10^{-4}$  with L<sub>int</sub> = 20 fb<sup>-1</sup>, better than the current limits set by BaBar
  - ►  $Z' \rightarrow inv.$ , expected sensitivity: g' ~ 10<sup>-2</sup> 10<sup>-1</sup> with Phase 2 data; with Phase 3: possibility to exclude the parameter region that explain (g-2)<sub>u</sub>
  - ► ALPs,  $a \rightarrow \gamma \gamma$ , expected sensitivity:  $g_{a\gamma\gamma} \sim 10^{-3} 10^{-2}$  with Phase 2 data, better than current limits

# Thank you

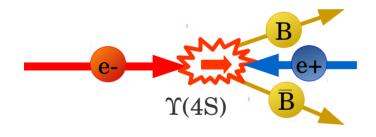
#### Luigi Corona ~ INFN and University of Pisa

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#### **B-Factories**

**B-factory**: asymmetric e+/e- collider (SuperKEKB: E(e-) = 7 GeV, E(e+) = 4 GeV ) optimized for the production of B mesons (but also charm physics, tau physics...)

 $e^+e^- \rightarrow \Upsilon(4S)[bb] (10.58 \text{ GeV/c}^2)$ B.R.( $\Upsilon(4S) \rightarrow BB$ ) > 96%,



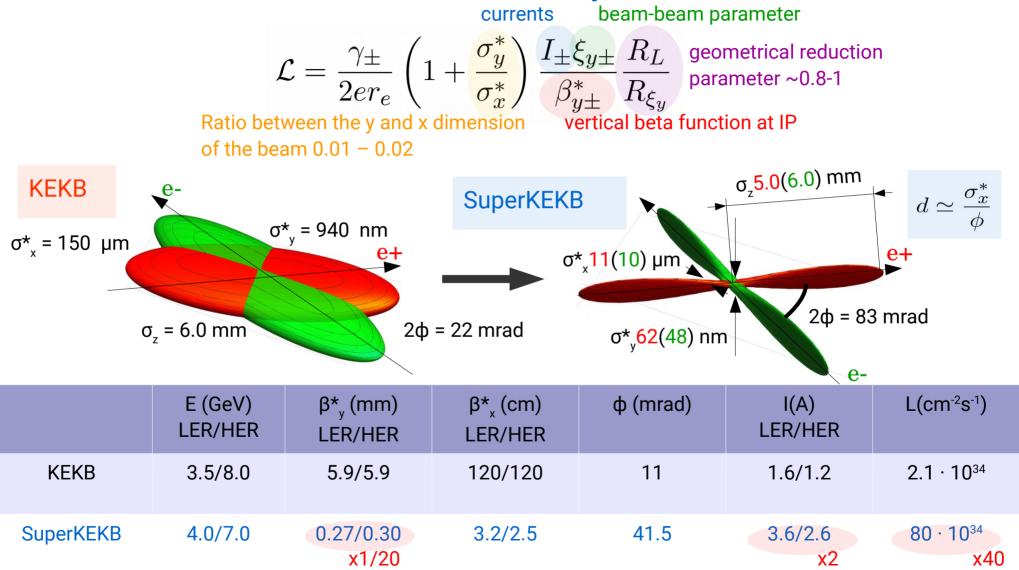
Process	Cross Section [nb]
$e^+e^- \rightarrow \mu^+\mu^-$	$1.148 \pm 0.005$ (full angle)
$e^+e^-  ightarrow \tau^+ \tau^-$	$0.919 \pm 0.003$ (full angle)
$e^+e^- \rightarrow e^+e^-(\gamma)$	$294 \pm 2 \ (10\text{-}170 \ \text{deg})$
$e^+e^-  ightarrow \gamma\gamma(\gamma)$	$4.96 \pm 0.02 \ (10\text{-}170 \ \text{deg})$
$e^+e^- \rightarrow e^+e^-e^+e^-$	$39.74 \pm 0.03$ (full angle)
$e^+e^-  ightarrow e^+e^-\mu^+\mu^-$	$18.87 \pm 0.02$ (full angle)
$e^+e^- \rightarrow u\bar{u}(\gamma)$	1.605 (full angle)
$e^+e^- \rightarrow d\bar{d}(\gamma)$	0.401 (full angle)
$e^+e^- \rightarrow s\bar{s}(\gamma)$	0.383 (full angle)
$e^+e^- \to c\bar{c}(\gamma)$	1.329 (full angle)
$e^+e^- \to \Upsilon(4S) \to B^+B^-$	0.5346 (full angle)
$e^+e^- \to \Upsilon(4S) \to B\bar{B}^0$	0.5654 (full angle)

First generation of B-factories:

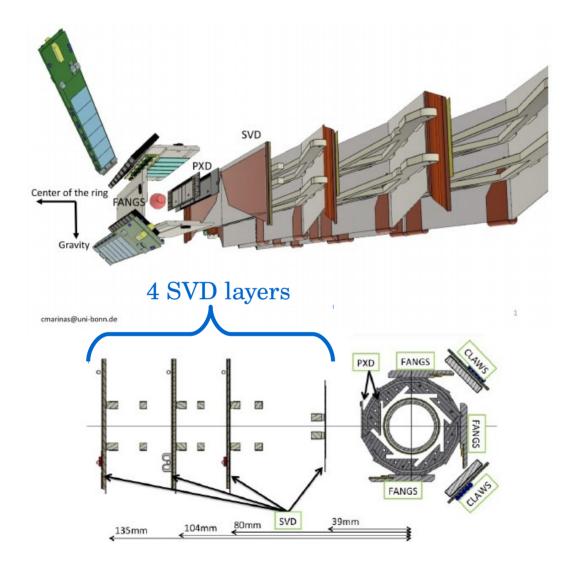
- BaBar at the PEP II collider (SLAC California)
- Belle at the KEKB collider (KEK, Japan)

**Some features**: well known initial state, high signal / noise ration, detector with high angular acceptance and composed of several subdetectors

#### Nano-beam scheme and luminosity



#### Phase 2 and Phase 3 VXD geometry (1/2)



Phase 2

CLAWS

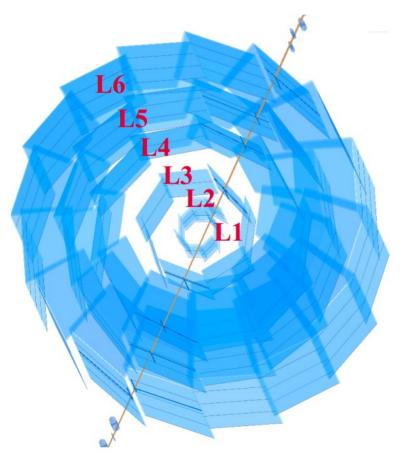
FANGS

PLUME

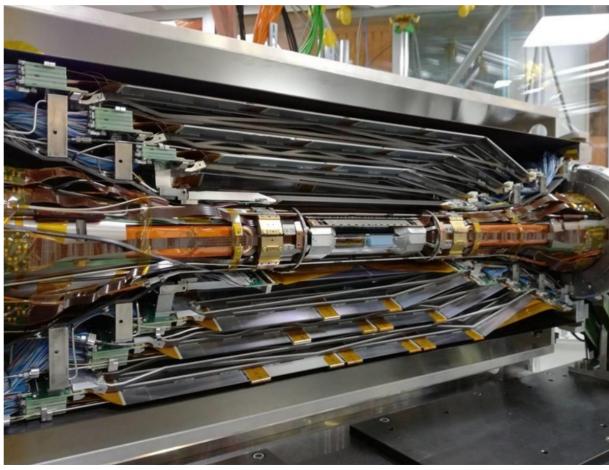
- diamond sensors
  - study and monitor the beam background levels

# Phase 2 and Phase 3 VXD geometry (2/2)

- SVD L3,4,5,6  $\rightarrow$  Low material budget, precise hit time resolution ( $\sigma \sim 3$  ns)
- PXD L1,2  $\rightarrow$  Low material budget, innermost layer at 1.4 cm

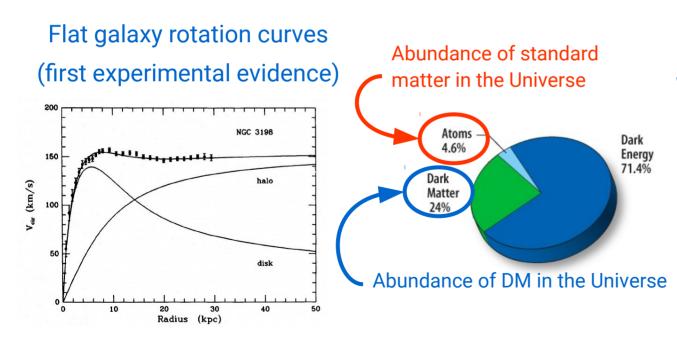


#### Phase 3

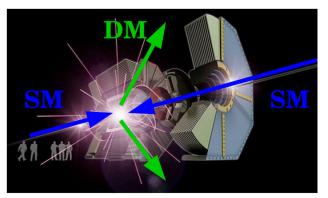


#### Dark Matter (DM): Introduction

- **Massive**  $\rightarrow$  gravitational interation with Standard Model (SM) matter
- **Dark**  $\rightarrow$  does not interact with SM matter thorugh any other interaction
- Many astrophysical observations in agreement with DM existance: flat galaxy rotation curves, gravitational lensing, galaxy velocity dispersion...



If DM exists as particles and interact with SM, although very weakly, it is possible to produce it in colliders



#### **Kinetic mixing**

• Extension of the SM:  $\rightarrow$  additional U(1)' symmetry that mix with the photon

$$A_{\mu} \qquad A'_{\mu} \qquad -\frac{\varepsilon}{2} F_{\mu\nu} F'^{\mu\nu}$$

- Off diagonal kinetic term
- $\varepsilon$  is the strengh of the kinetic mixing ( $\varepsilon \le 10^{-2}$ )
- After the ridefinition of fields ( $A_{\mu} \rightarrow A_{\mu} \epsilon A'_{\mu}$ ) the diagonal kinetic term is restored and the interaction term  $\epsilon J^{\mu}_{SM} A'_{\mu}$  arises in the theory
- The symmetry U(1)' can be broken spontaneously by a dark Higgs mechanism that gives mass to the dark photon

References:

- P. Fayet, Phys. Lett. B 95, 285 (1980),
- P. Fayet Nucl. Phys. B 187, 184 (1981)
- B. Holdom, Phys. Lett. B 166, 196 (1986)

## Other planned dark sector and exotic searches

- Visible dark photon decays
- Off-shell dark photon decays\*\*\*
- Muonic dark force:  $e^+e^- \rightarrow \mu^+\mu^- Z'$ ,  $Z' \rightarrow \mu^+\mu^-$
- Dark sector with Lepton Flavor Violation: Z'
- Dark scalar:  $e^+e^- \rightarrow \tau^+\tau^-S$ ,  $S \rightarrow l^+l^-$
- Magnetic monopoles with small magnetic charges\*\*\*
- Invisible Y(1S) decays via Y(3S) → Y(1S)π<sup>+</sup>π<sup>-</sup> (Requires beam energies at Y(3S))
- Dark Higgs/Higgstrahlung

BaBar: arXiv:1606.03501 (514 fb<sup>-1</sup>)

For further details:

The Belle II Physics Book arXiv:1808.10567

\*\*\*Possible with Phase 2 data

#### Magnetic monopoles

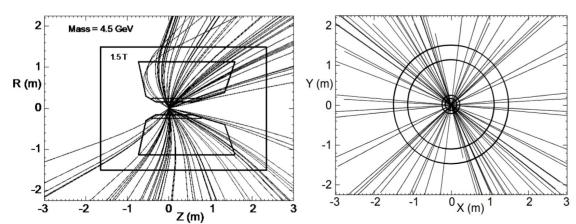
Particle carrying magnetic charge

Distinct signature in drift chamber:

- Tracks are strainght in (x,y) plane
- Tracks are curved in (r,z)

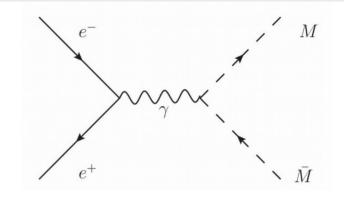
They need a dedicated tracking system

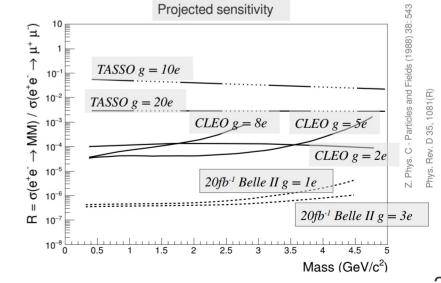
Detection efficiency is high: 40-97%, depending on magneton mass



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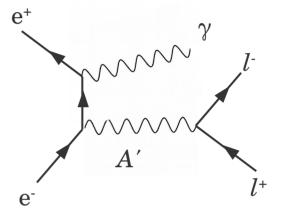
M. K. Sullivan, D. Fryberger arXiv:1707.05295





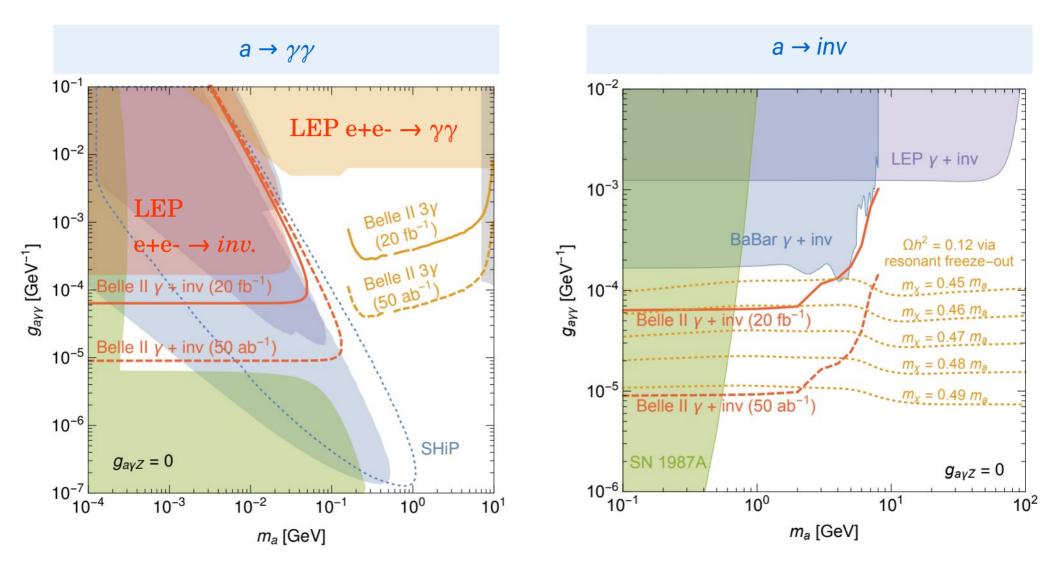
#### Projected upper limits on ε, visible dark photon

 $e^+e^- \rightarrow \gamma A', A' \rightarrow l^+l^ 10^{-2}$ **KLOE 2013** 2012 ω Phenix (g-2) WASA BaBar HADES (g-2) ± 20 favored **10<sup>-3</sup>** elle II 50 fb 500 fb<sup>1</sup> E774 Dark 5 ab<sup>1</sup> 50 ab E141 10-4 **10<sup>-1</sup>**  $10^{-2}$ 10 m<sub>₄'</sub> (GeV)

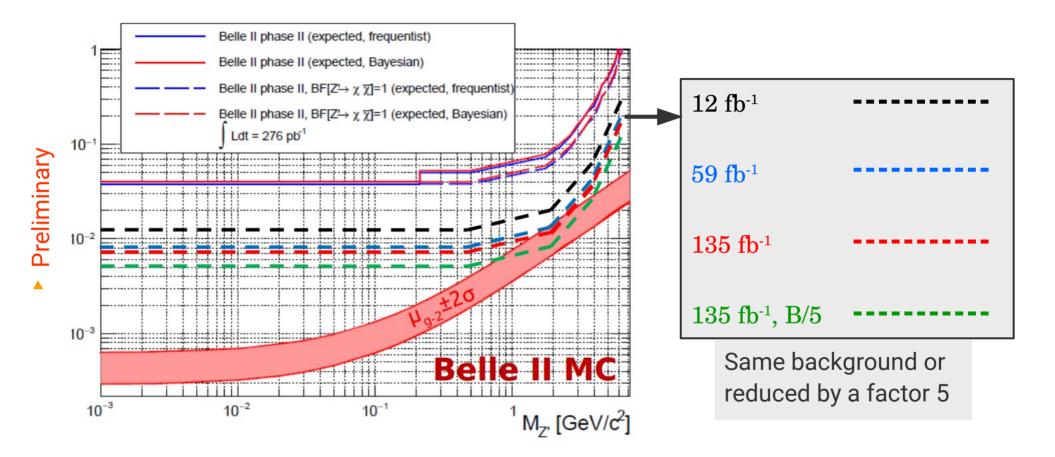


Belle II is competitive only in Phase 3
The Belle II Physics Book arXiv:1808.10567

#### ALPs expected limits



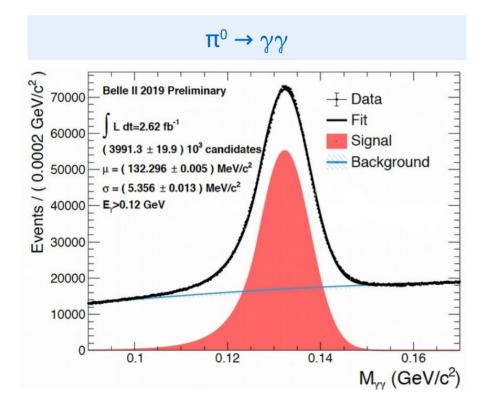
## Projected upper limits on g' with Phase 3 data



#### Some possible factors of improvement: PID, vertex fit (full VXD), Multivariate Analysis

# Highlights from Phase 3

- Results for early Phase 3 data:
  - based on 2.62 fb-1



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#### $J/\psi \rightarrow e^+e^-$

