



Search for $e^+e^- \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$) at \sqrt{s} near 10.746 GeV at Belle II

W. T. Xiong,^{1,*} S. Jia,^{2,†} and C. P. Shen^{1,3,‡}

(The Belle II Collaboration)

¹*Fudan University, Shanghai, China*

²*Southeast University, Nanjing, China*

³*Zhengzhou University, Henan, China*

Abstract

1 This note describes the study of $e^+e^- \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$) at center-of-mass energies \sqrt{s}
2 = 10.653, 10.701, 10.746, and 10.804 GeV. These data were collected with the Belle II detector
3 during special operations of the SuperKEKB collider, corresponding to 3.5, 1.6, 9.8, and 4.7 fb⁻¹ of
4 integrated luminosity, respectively. We do not find evidence for any process at each energy point.
5 The upper limits at 90% confidence level on the Born cross sections for $e^+e^- \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1,$
6 2) at \sqrt{s} near 10.746 GeV are set.

VERSION LOG FOR MAIN CHANGES

– Updates in Belle II Note v3.1

- Add the source of systematic uncertainty from the signal PDF parameters. Please see Sec. 4.

– Updates in Belle II Note v3.0:

- The box is open and we update the open box result. Please see Sec. 3.3.5.
- Dividing systematic uncertainty into additive systematic uncertainty and multiplicative systematic uncertainty, and separately account them in the the calculations of upper limits. Please see Sec. 4.

– Updates in Belle II Note v2.2:

- Update the fitting procedure and add input/output check to test the stability of this fit. Please see Sec. 3.3.4.

– Updates in Belle II Note v1.3:

- Change the 5C fit into 4C fit for $\gamma\gamma l^+l^-$ final states. Please see Sec. 3.3.1.
- Update the systematic uncertainties from photon reconstruction efficiency and beam-energy calibration. Please see Sec. 4.

– Updates in Belle II Note v1.2:

- Add systematic uncertainty due to the effect of initial state radiative photon in Sec. 4.

– Updates in Belle II Note v1.1:

- Remove the requirement for E_{γ_h} since the E_{γ_h} is correlated to the $M(\gamma\Upsilon(1S))$ and add selection criterion of $|\cos(\theta_{\gamma_h})| < 0.7$. Please see Sec. 3.3.1. All of the results are revised accordingly.

CONTENTS

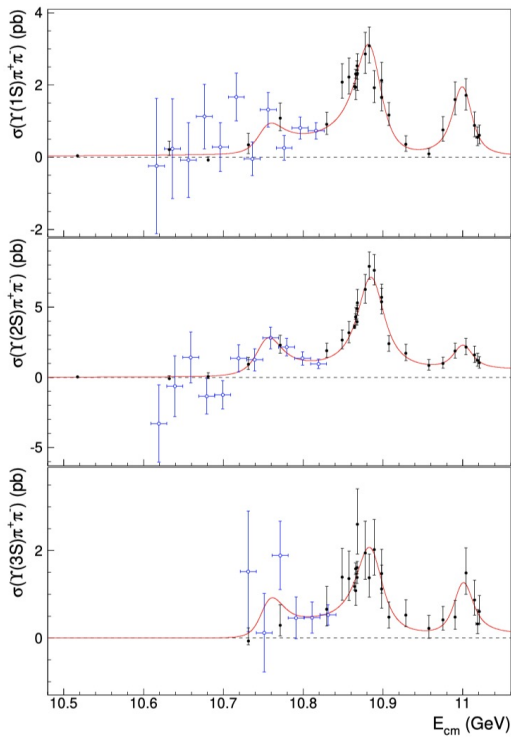
1. Introduction	4
1.1. Motivation	4
1.2. Research method	6
2. Dataset	7
2.1. Signal MC samples	7
2.2. Generic MC samples	7
2.3. Data samples	7
3. Analysis of $\Upsilon(10753) \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$)	8
3.1. Selection criteria	8
3.2. The study of signal MC samples	14
3.2.1. Distribution from signal MC samples	15
3.3. The study of generic MC samples	15
3.4. Signal extraction	17
3.5. The Study of data samples	18
4. Systematic uncertainty	20
5. Summary	25
References	25

7 **1. INTRODUCTION**

8 **1.1. Motivation**

9 We study the process of $e^+e^- \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$) at \sqrt{s} near 10.746 GeV to search for
 10 the radiative decay of $\Upsilon(10753)$, $\Upsilon(10753) \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$).

11 The $e^+e^- \rightarrow \Upsilon(nS)\pi^+\pi^-$ ($n = 1, 2, 3$) cross sections at energies from 10.52 to 11.02
 12 GeV have been measured by Belle, as shown in Fig. 1 [5]. A structure with the mass of
 13 $(10752.7 \pm 5.9_{-1.1}^{+0.7})$ MeV/ c^2 and width of $(35.5_{-11.3}^{+17.6+3.9}_{-3.3})$ MeV/ c^2 was observed with a global
 14 significance of 5.2σ including systematic uncertainty. And recent measurement updates the
 15 mass of $(10756.3 \pm 2.7 \pm 0.6)$ MeV/ c^2 and width of $(29.7 \pm 8.5 \pm 1.1)$ MeV/ c^2 [6]. We call
 16 this state as $\Upsilon(10753)$.



17 FIG. 1. Energy dependency of the $e^+e^- \rightarrow \Upsilon(nS)\pi^+\pi^-$ cross sections ($n = 1, 2, 3$ from top to
 18 bottom). The points with error bars are data; the curves are the results of the fit. Blue open dots
 19 with error bars are cross sections estimated using the ISR process in the $\Upsilon(10860)$ on-resonance
 20 data.

21 The $\Upsilon(10753)$ has been interpreted as a conventional bottomonium [7–16], hybrid [17,
 22 18], or tetraquark state [19–23]. The nature of the $\Upsilon(10753)$ is unclear. To get a better
 23 knowledge of the nature of this new resonance $\Upsilon(10753)$, Belle II experiment conducted a
 24 higher statistics scan near 10.75 GeV in November 2021.

25 Some channels have been measured to study the nature of $\Upsilon(10753)$ energy region. In
 26 2022, the signal of $e^+e^- \rightarrow \omega\chi_{b1}(1P)$ at $\sqrt{s} = 10.746$ GeV and evidence of $e^+e^- \rightarrow \omega\chi_{b2}(1P)$
 have been observed by using full data sample near 10.746 GeV of Belle II experiment, as
 shown in Fig. 2 [24]. Measurements of the $e^+e^- \rightarrow B\bar{B}, B\bar{B}^*$, and $B^*\bar{B}^*$ cross sections at

27 10.653, 10.701, 10.746, and 10.804 GeV were performed. In the first 2 – 5 MeV above $B^*\bar{B}^*$
 28 threshold, the $e^+e^- \rightarrow B^*\bar{B}^*$ cross section increases rapidly. This may indicate the presence
 29 of a pole close to the threshold [26]. The processes of $e^+e^- \rightarrow \omega\eta_b(1S)$ and $e^+e^- \rightarrow \omega\chi_{b0}(1P)$
 30 were searched for using the ω meson’s recoil mass with Belle II scan data at $\sqrt{s} = 10.746$
 31 GeV [27]. No evidence for either process is found, and upper limits on the corresponding
 32 Born cross sections for $e^+e^- \rightarrow \omega\eta_b(1S)$ and $e^+e^- \rightarrow \omega\chi_{b0}(1P)$ were set to be 2.5 pb and
 33 7.8 pb, respectively, at the 90% confidence level.

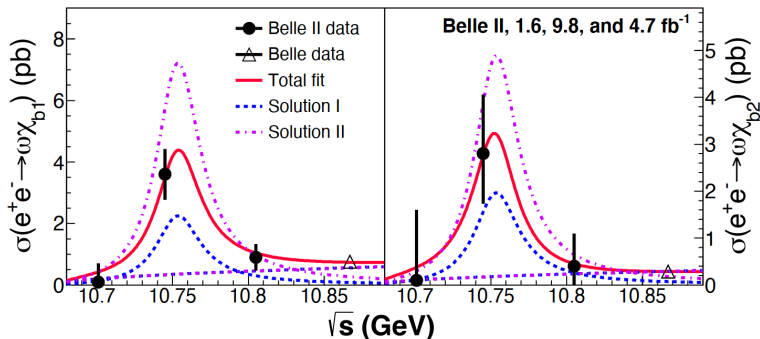


FIG. 2. The Born cross sections for $e^+e^- \rightarrow \omega\chi_{b1}$ (left) and $e^+e^- \rightarrow \omega\chi_{b2}$ (right) as a function of C.M. energy by combining the measurements of $\sigma^B(e^+e^- \rightarrow \omega\chi_{b1})$ ($\sigma^B(e^+e^- \rightarrow \omega\chi_{b2})$) from Belle at $\sqrt{s} = 10.867$ GeV [25] and Belle II at $\sqrt{s} = 10.701, 10.746,$ and 10.804 GeV [24].

34 The radiative decay of $\Upsilon(10753)$ has not been studied experimentally. From theoretical
 35 calculations, the bottomonium with orbital angular momentum $L = 2$ decays into $\gamma\chi_{bJ}(1P)$
 36 abundantly [28, 29]. For examples, the branching fractions for $\Upsilon_1(2^3D_1) \rightarrow \gamma\chi_{bJ}(1P)$ could
 37 be as large as a few percent, as listed in Table I [28, 29]. Therefore, if the $\Upsilon(10753)$ consists
 38 of a D -wave component, the branching fraction for $\Upsilon(10753) \rightarrow \gamma\chi_{bJ}(1P)$ is expected to be
 39 large as well. Measurements of production rates of $\Upsilon(10753) \rightarrow \gamma\chi_{bJ}(1P)$ help distinguish
 40 the nature of $\Upsilon(10753)$.

41 In addition, the cross section distribution for $e^+e^- \rightarrow \pi^+\pi^-\Upsilon(nS)$ [5] exhibits two closely
 42 spaced peaks corresponding to the $\Upsilon(10753)$ and $\Upsilon(10860)$, analogous to the two-peak struc-
 43 ture observed in the cross sections for $e^+e^- \rightarrow \pi^+\pi^-J/\psi$ at $Y(4230)$ and $Y(4320)$ by BE-
 44 SIII [30]. Evidence for $e^+e^- \rightarrow \gamma\chi_{c1}$ and $e^+e^- \rightarrow \gamma\chi_{c2}$ was observed by BESIII with
 45 statistical significances of 3.0σ and 3.4σ , respectively, using combined data samples at \sqrt{s}
 46 = 4.009, 4.230, 4.260, and 4.360 GeV. If $Y(4230)$ and $\Upsilon(10753)$ share similar internal dy-
 47 namics, we would expect enhanced production of $e^+e^- \rightarrow \gamma\chi_{b1}$ and $e^+e^- \rightarrow \gamma\chi_{b2}$ near \sqrt{s}
 48 = 10.746 GeV.

TABLE I. The calculated $\Upsilon_1(2^3D_1) \rightarrow \gamma\chi_{bJ}(1P)$ decay branching fractions [28, 29].

Channels	\mathcal{B} from Ref. [28]	\mathcal{B} from Ref. [29]
$\Upsilon_1(2^3D_1) \rightarrow \gamma\chi_{b0}(1P)$	7.7×10^{-2}	6.88×10^{-2}
$\Upsilon_1(2^3D_1) \rightarrow \gamma\chi_{b1}(1P)$	2.0×10^{-2}	2.37×10^{-2}
$\Upsilon_1(2^3D_1) \rightarrow \gamma\chi_{b2}(1P)$	5.0×10^{-4}	6.90×10^{-4}

49 **1.2. Research method**

50 The Born cross section for $e^+e^- \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$) is calculated by:

$$\sigma^B(e^+e^- \rightarrow \gamma\chi_{bJ}) = \frac{N_{\text{fit}} \times |1 - \Pi|^2}{\mathcal{L} \times (\varepsilon_{e^+e^-} + \varepsilon_{\mu^+\mu^-}) \times \mathcal{B}_{\text{int}} \times (1 + \delta_{\text{ISR}})}, \quad (1)$$

51 where N_{fit} is the signal yield of the fit combined from $\mu^+\mu^-$ and e^+e^- modes, \mathcal{L} is the
 52 integrated luminosity of data sample, $\varepsilon_{e^+e^-}$ and $\varepsilon_{\mu^+\mu^-}$ are reconstruction efficiencies for
 53 $\Upsilon(1S) \rightarrow e^+e^-$ and $\Upsilon(1S) \rightarrow \mu^+\mu^-$ modes, $\mathcal{B}_{\text{int}} = \mathcal{B}(\chi_{bJ} \rightarrow \gamma\Upsilon(1S)) \times \mathcal{B}(\Upsilon(1S) \rightarrow l^+l^-)$ is
 54 the product branching fraction of the intermediate states. The factor $|1 - \Pi|$ is the vacuum
 55 polarization factor [33], which is listed in Table II for different energy points.

TABLE II. The values of the vacuum polarization factor at $\sqrt{s} = 10.653, 10.701, 10.746,$ and 10.804 GeV.

e^+e^- center-of-mass energy	Vacuum polarization factor
10.653 GeV	0.930
10.701 GeV	0.931
10.746 GeV	0.931
10.804 GeV	0.932

56

57

58 The $(1 + \delta_{\text{ISR}})$ is the radiative correction factor, which can be calculated by [34–36]:

$$1 + \delta_{\text{ISR}} = \frac{\int \sigma^{\text{dressed}}(s(1-x))F(x,s)dx}{\sigma^{\text{dressed}}}, \quad (2)$$

59 where s is the square of the center-of-mass (C.M.) energy, and x is the fraction of energy
 60 carried by the ISR photon. And $F(x,s)$ is the radiative function obtain from QED calcula-
 61 tion [34–36], which can be described as

$$F(x,s) = v(s)x^{v(s)-1}(1 + \delta(s)) + x^{v(s)}(-v(s) - v(s)^2/4) + x^{v(s)+1}(v(s)/2 - \frac{3}{8}v(s)^2), \quad (3)$$

62 where,

$$v(s) = \frac{2\alpha}{\pi}(\log \frac{s}{m_e^2} - 1), \quad (4)$$

63

$$\delta(s) = \frac{\alpha}{\pi}(\frac{\pi^2}{3} - \frac{1}{2}) + \frac{3}{4}v(s) + v(s)^2(\frac{9}{32} - \frac{\pi^2}{12}), \quad (5)$$

64 where α is the fine structure constant, and m_e is the mass of electron. In Eq. (2), the
 65 integration limit corresponds to the $\gamma\chi_{bJ}$ mass threshold, which corresponds to the left end
 66 of $\Upsilon(10753)$ mass. The σ^{dressed} is the “dressed” cross section, and we assume

$$\sigma^{\text{dressed}}(\sqrt{s}) = \frac{1}{s}, \quad (6)$$

67 The values of radiative correction factors of different C.M. energies and different final states
 68 are calculated, which are listed in Table III.

TABLE III. The values of radiative correction factor for $e^+e^- \rightarrow \gamma\chi_{b0}$, $e^+e^- \rightarrow \gamma\chi_{b1}$, and $e^+e^- \rightarrow \gamma\chi_{b2}$ at $\sqrt{s} = 10.653, 10.701, 10.746,$ and 10.804 GeV, respectively.

$1 + \delta_{\text{ISR}}$	$\sqrt{s} = 10.653$ GeV	$\sqrt{s} = 10.701$ GeV	$\sqrt{s} = 10.746$ GeV	$\sqrt{s} = 10.804$ GeV
$e^+e^- \rightarrow \gamma\chi_{b0}$	0.901	0.905	0.909	0.914
$e^+e^- \rightarrow \gamma\chi_{b1}$	0.898	0.905	0.906	0.911
$e^+e^- \rightarrow \gamma\chi_{b2}$	0.896	0.901	0.904	0.909

69 2. DATASET

70 2.1. Signal MC samples

71 We generate 50k signal Monte Carlo (MC) events for each energy point and each pro-
 72 cess to determine the signal line shapes and the reconstruction efficiencies. A combination
 73 of PHOKHARA [37] and EVTGEN [38] based on the release-06-00-08 is used. We use the
 74 PHOKHARA generator [37] to generate the $e^+e^- \rightarrow V\gamma_{\text{ISR}}$ process at the next to leading
 75 order accuracy. And we simulate the decay chain $V \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$) using EVTGEN
 76 generator [38]. Here the angular distribution is in phase space. And we will change the
 77 angular distribution by $1 \pm \cos^2\theta$ and take the difference in the efficiency as systematic un-
 78 certainty due to the angular distribution. For $\chi_{b0} \rightarrow \gamma\mathcal{Y}(1S)$ decay, HLEAMP mode with
 79 parameters (1. 0. 1. 0.) is used. For $\chi_{b1} \rightarrow \gamma\mathcal{Y}(1S)$ decay, HLEAMP mode with parameters
 80 (1. 0. 1. 0. -1. 0. -1. 0.) is used. For $\chi_{b2} \rightarrow \gamma\mathcal{Y}(1S)$ decay, PHSP mode is used. For
 81 $\mathcal{Y}(1S) \rightarrow l^+l^-$ decay, VLL mode is used. The effect of final-state radiation is taken into
 82 account in the simulation using the PHOTOS package.

83 Run-dependent signal MC samples are produced by including beam background overlay-
 84 ing samples. The GEANT4 [39] is used to simulate the detector response.

85 2.2. Generic MC samples

86 We use official generic MC samples at \sqrt{s} near 10.746 GeV in this analysis. It corresponds
 87 to Run-dependent MC – proc13 + prompt-5S scan. The details are shown in Table IV.
 88 Besides above, we also used e^+e^- type (Bhabha) events with an integrated luminosity of
 89 17.481 fb^{-1} from Run-dependent MC – prompt - 4S exp 20-26.

90 2.3. Data samples

91 $\mathcal{Y}(10750)$ scan data samples from prompt reprocessing bucket#27 are used, and details
 92 are shown in Table V.

93 – C.M. energy: 10.653, 10.701, 10.746, and 10.804 GeV

94 We use the global tag of beamEnergy_bucket27_prompt [40] to correct beam energies and
 95 the C.M. energies in data.

TABLE IV. Event type and size of generic MC samples (Run-dependent MC – proc13 + prompt-5S scan).

Event type	Collected Luminosity (fb ⁻¹)
$u\bar{u}$	200
$d\bar{d}$	200
$s\bar{s}$	200
$c\bar{c}$	200
$\tau^+\tau^-$	1000
$e^+e^-e^+e^-$	100
$e^+e^-\mu^+\mu^-$	100
B^+B^-	200
$B^0\bar{B}^0$	200

TABLE V. The luminosity, experiment number, and run range of $\Upsilon(10750)$ scan data samples.

Energy (GeV)	Collected luminosity (fb ⁻¹)	Exp No.	Run range
10.653	3.52±0.01	21	83-186
10.701	1.63±0.01	21	192-235
10.746	9.82±0.02	21	245-490
10.804	4.69±0.02	21	507-572

96 3. ANALYSIS OF $\Upsilon(10753) \rightarrow \gamma\chi_{bJ} (J = 0, 1, 2)$

97 3.1. Selection criteria

98 The following selection criteria aim to select signal and reduce background as efficiently
 99 as possible. The final states for $e^+e^- \rightarrow \gamma\chi_{bJ} (J = 0, 1, 2)$ are $\gamma\gamma l^+l^-$, where l represents e
 100 or μ .

101 • Charged tracks

102 All charged tracks are required to originate from the vicinity of the interaction point
 103 (IP). We require $|dr| < 1$ cm and $|dz| < 3$ cm, where dr and dz are transverse and
 104 longitudinal distances in respect to IP. Figure 3 shows the distributions of $|dr|$ and $|dz|$
 105 from signal MC and data samples at $\sqrt{s} = 10.746$ GeV. Please note that all “Data” in
 106 the figures in this chapter are real data samples at $\sqrt{s} = 10.746$ GeV. The red arrows
 107 show the requirements of $|dr| < 1$ cm and $|dz| < 3$ cm.

108 - For the lepton pairs l^+l^- used to reconstruct $\Upsilon(1S)$, we require the electronID_noSVD_noTOP
 109 ($\mathcal{R}_e = \frac{\mathcal{L}_e}{\mathcal{L}_e + \mathcal{L}_\mu + \mathcal{L}_\pi + \mathcal{L}_K + \mathcal{L}_p + \mathcal{L}_d}$) > 0.9 in the e^+e^- mode and muonID_noSVD ($\mathcal{R}_\mu =$
 110 $\frac{\mathcal{L}_\mu}{\mathcal{L}_e + \mathcal{L}_\mu + \mathcal{L}_\pi + \mathcal{L}_K + \mathcal{L}_p + \mathcal{L}_d}$) > 0.9 in the $\mu^+\mu^-$ mode. Figure 4 shows the distributions of
 111 electronID_noSVD_noTOP and muonID_noSVD from signal MC and data samples at
 112 $\sqrt{s} = 10.746$ GeV. The red arrows show the requirements of electronID_noSVD_noTOP
 113 > 0.9 and muonID_noSVD > 0.9 .

114 - The number of good charged tracks is required to be two to suppress background.

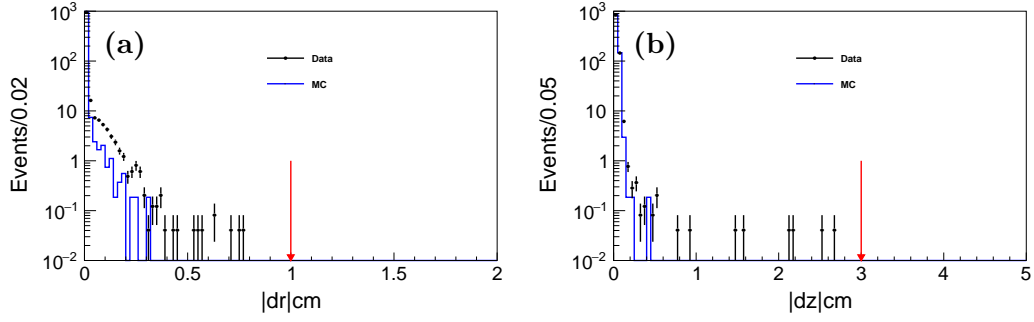


FIG. 3. Distributions of (a) $|dr|$ and (b) $|dz|$ from signal MC and data samples at $\sqrt{s} = 10.746$ GeV.

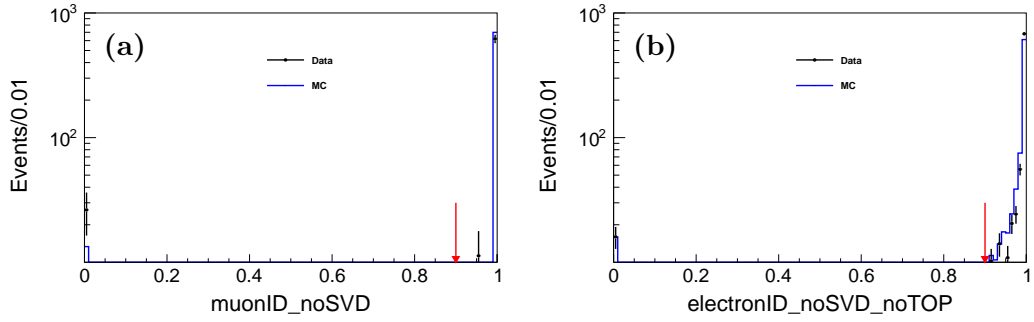


FIG. 4. Distributions of (a) muonID_noSVD (\mathcal{R}_μ) and (b) $\text{electronID_noSVD_noTOP}$ (\mathcal{R}_e) from signal MC and data samples at $\sqrt{s} = 10.746$ GeV.

115

Figure 5 shows the numbers of charged tracks (N_{trk}) from signal MC and data sample

116

at $\sqrt{s} = 10.746$ GeV.

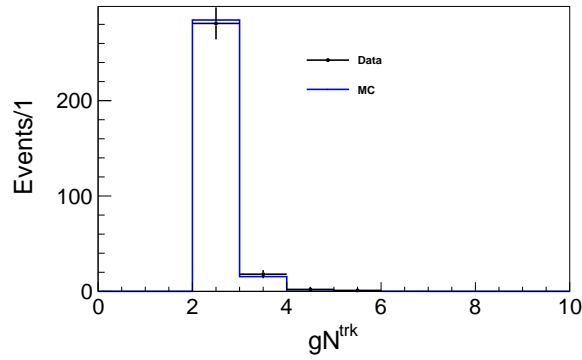


FIG. 5. Distributions of the numbers of good charged tracks (N_{trk}) from signal MC and data samples at $\sqrt{s} = 10.746$ GeV.

117

- Photons

118

There are two photons in $e^+e^- \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$) final states. We require $E9/E21$

119
120
121
122
123
124
125

> 0.8 for all the photons, where the E9/E21 is a ratio of energies in the inner 3×3 crystals and 5×5 crystals around the central crystal without corners. The photon from χ_{bJ} ($J = 0, 1, 2$) decay has lower energy, thus is remarked as γ_l . The photon in $e^+e^- \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$) has higher energy, thus is remarked as γ_h . Thus, we have $e^+e^- \rightarrow \gamma_h\chi_{bJ}$, $\chi_{bJ} \rightarrow \gamma_l\Upsilon(1S)$, and $\Upsilon(1S) \rightarrow \ell^+\ell^-$.

Figure 6 shows the E_{γ_h} distributions in $e^+e^- \rightarrow \gamma_h\chi_{bJ}(\rightarrow \gamma_l\Upsilon(1S))$ ($J = 0, 1, 2$) from signal MC samples at $\sqrt{s} = 10.746$ GeV.

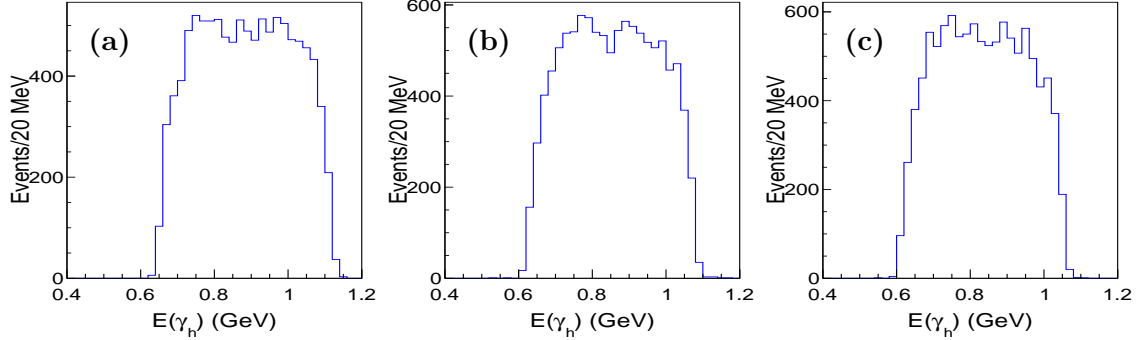


FIG. 6. The E_{γ_h} distributions in (a) $e^+e^- \rightarrow \gamma_h\chi_{b0}$, (b) $e^+e^- \rightarrow \gamma_h\chi_{b1}$, and (c) $e^+e^- \rightarrow \gamma_h\chi_{b2}$ from signal MC samples at $\sqrt{s} = 10.746$ GeV.

126
127

Figure 7 shows the E_{γ_l} distributions in $e^+e^- \rightarrow \gamma_h\chi_{bJ}(\rightarrow \gamma_l\Upsilon(1S))$ ($J = 0, 1, 2$) from signal MC samples at $\sqrt{s} = 10.746$ GeV.

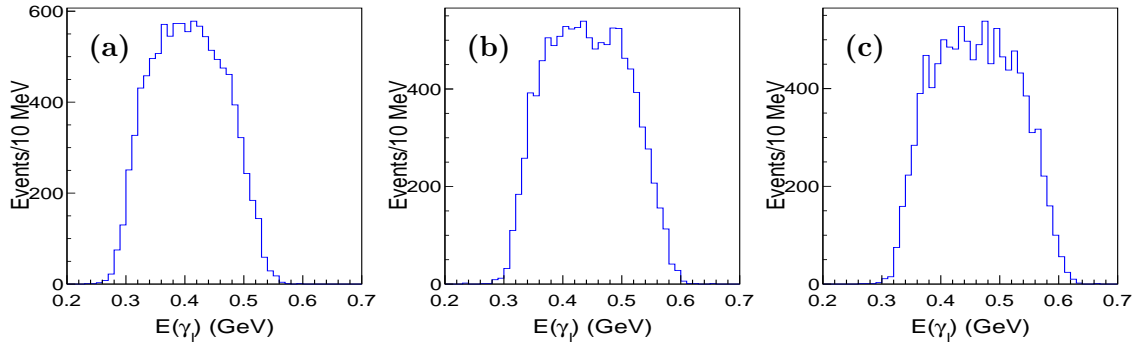


FIG. 7. The E_{γ_l} distributions in (a) $e^+e^- \rightarrow \gamma_h\chi_{b0}$, (b) $e^+e^- \rightarrow \gamma_h\chi_{b1}$, and (c) $e^+e^- \rightarrow \gamma_h\chi_{b2}$ from signal MC samples at $\sqrt{s} = 10.746$ GeV.

128
129
130
131
132
133
134

Figure 8 shows the scattering plots of the $E_{\gamma_h}^*$ (the energy of γ_h in e^+e^- center-of-mass frame) versus $M(\gamma\Upsilon(1S))$ in $e^+e^- \rightarrow \gamma_h\chi_{b0}$. The $E_{\gamma_h}^*$ is highly correlated with the $M(\gamma\Upsilon(1S))$. The selected region of $M(\gamma\Upsilon(1S))$ equals to the application of requirement of E_{γ_h} to some extent. Thus, we just take a loose “protective” selection criterion of $E_{\gamma_h} > 50$ MeV.

We optimize the E_{γ_l} by maximizing the FOM parameter $\frac{\varepsilon}{\frac{3}{2} + \sqrt{B}}$ [41], where ε is the efficiency from signal MC samples and B is the number of background events from

135
136
137
138
139
140
141
142
143
144

normalized generic MC samples. Figure 9 shows the FOM parameter distributions of E_{γ_l} for $\mu^+\mu^-$ mode and e^+e^- mode in $e^+e^- \rightarrow \gamma_h\chi_{b0}(\rightarrow \gamma_l\Upsilon(1S))$.

From Figure 18, the data sidebands are basically consistent with generic MC. If we use the $\Upsilon(1S)$ mass sidebands for estimating the number of background events, the optimized results do not change.

Please note that we identify the signal via the same $M(\gamma_l\Upsilon(1S))$ distribution for different $\chi_{b0,b1,b2}$ decays. Thus, for different $\chi_{b0,b1,b2}$ decays, the requirements should be the same. Here, we chose the $e^+e^- \rightarrow \gamma\chi_{b0}$ channel to optimize the E_{γ_l} requirement to obtain the most conservative result. We require $E_{\gamma_l} > 280$ MeV for $\mu^+\mu^-$ mode and $E_{\gamma_l} > 290$ MeV for e^+e^- mode, respectively, at each energy point.

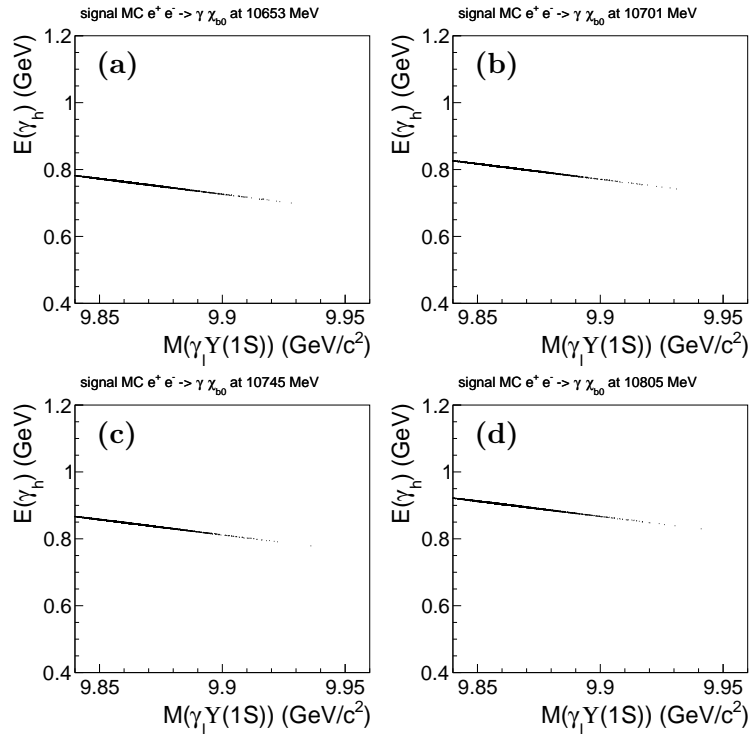


FIG. 8. The scattering plots of the $E_{\gamma_h}^*$ (the energy of γ_h in e^+e^- center-of-mass frame) versus $M(\gamma\Upsilon(1S))$ from MC simulated $e^+e^- \rightarrow \gamma_h\chi_{b0}(\rightarrow \gamma_l\Upsilon(1S))$ at (a) $\sqrt{s} = 10.653$ GeV, (b) $\sqrt{s} = 10.701$ GeV, (c) $\sqrt{s} = 10.746$ GeV, and (d) $\sqrt{s} = 10.804$ GeV, respectively.

145
146
147
148
149
150
151
152
153

- Helicity angle

Figure 10 shows the distributions of the absolute value of the cosine of helicity angle of γ_h from generic MC sample at $\sqrt{s} = 10.746$ GeV. Here, the helicity angle is the polar angle of γ_h in the e^+e^- center-of-mass frame. More Bhabha background events are accumulated in the higher side. We optimize the $|\cos(\theta_{\gamma_h})|$ for the e^+e^- mode by maximizing the FOM parameter $\frac{\varepsilon}{\frac{3}{2} + \sqrt{B}}$ [41], where ε is the efficiency from signal MC samples and B is the number of background events from normalized generic MC samples. Figure 11 shows the distributions of the cosine value of helicity angle of γ_h from MC simulated $e^+e^- \rightarrow \gamma_h\chi_{b0}(\rightarrow \gamma_l\Upsilon(1S))$ assuming $1 + \cos^2(\theta_{\gamma_h})$ distribution and

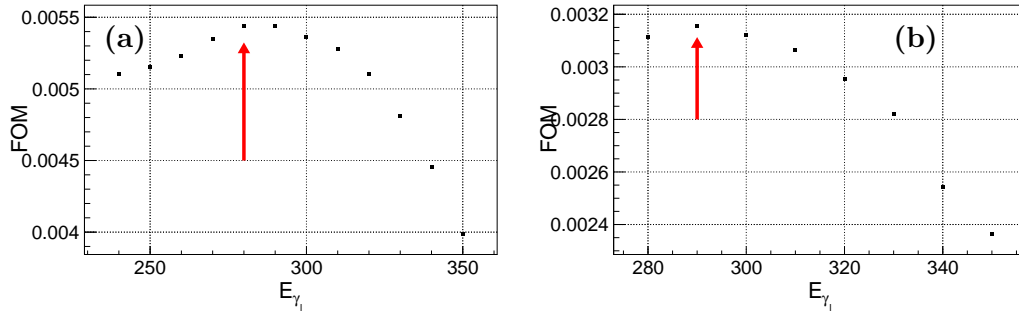


FIG. 9. The FOM parameter distributions of E_{γ_l} from MC simulated $e^+e^- \rightarrow \gamma_h \chi_{b0} (\rightarrow \gamma_l \Upsilon(1S))$ of (a) $\mu^+\mu^-$ final state and (b) e^+e^- final state at $\sqrt{s} = 10.746$ GeV.

154
155
156
157

$1 - \cos^2(\theta_{\gamma_h})$ distribution at $\sqrt{s} = 10.746$ GeV. Figure 12 shows the FOM parameter distributions of $|\cos(\theta_{\gamma_h})|$ for $1 + \cos^2(\theta_{\gamma_h})$ distribution and $1 - \cos^2(\theta_{\gamma_h})$ distribution. We require $|\cos(\theta_{\gamma_h})| < 0.7$ in the e^+e^- mode to suppress the Bhabha background events.

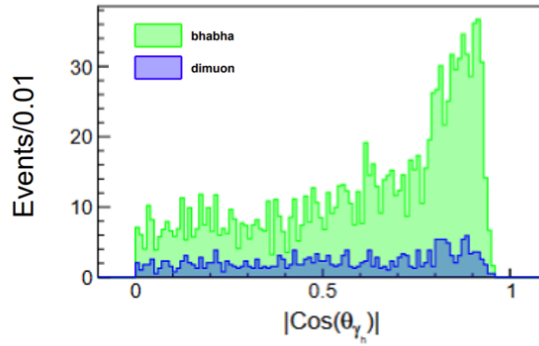


FIG. 10. The distributions of the absolute value of the cosine of helicity angle of γ_h from generic MC sample at $\sqrt{s} = 10.746$ GeV.

158
159
160
161
162
163
164
165
166
167
168

- $\Upsilon(1S) \rightarrow e^+e^-$

In order to consider the effect of bremsstrahlung or final-state radiation,

- $\Upsilon(1S)$ mass

The invariant mass distributions of lepton pairs are shown in Fig. 13. We require the mass of lepton pair larger than $9.44 \text{ GeV}/c^2$ and smaller than $9.49 \text{ GeV}/c^2$ to select $\Upsilon(1S)$ signal candidates. The red lines in Fig. 13 show the $\Upsilon(1S)$ signal region ($9.44 \text{ GeV}/c^2 < M(l^+l^-) < 9.49 \text{ GeV}/c^2$), and the green lines show the $\Upsilon(1S)$ mass sidebands ($9.34 \text{ GeV}/c^2 < M(l^+l^-) < 9.39 \text{ GeV}/c^2$ and $9.54 \text{ GeV}/c^2 < M(l^+l^-) < 9.59 \text{ GeV}/c^2$).

We optimize the $\Upsilon(1S)$ signal region by maximizing the FOM parameter $\frac{\varepsilon}{\frac{3}{2} + \sqrt{B}}$ [41], where ε is the efficiency and B is the number of background events from normalized

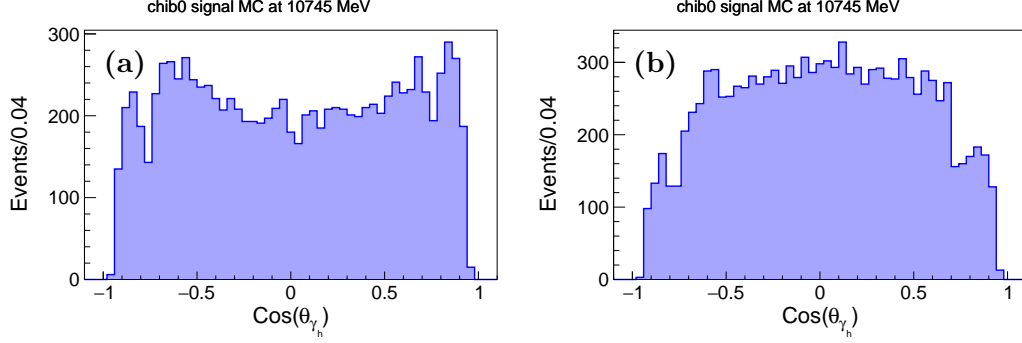


FIG. 11. The distributions of the cosine value of helicity angle of γ_h from MC simulated $e^+e^- \rightarrow \gamma_h \chi_{b0} (\rightarrow \gamma_l \Upsilon(1S))$ assuming (a) $1 + \cos^2(\theta_{\gamma_h})$ distribution and (b) $1 - \cos^2(\theta_{\gamma_h})$ distribution at $\sqrt{s} = 10.746$ GeV.

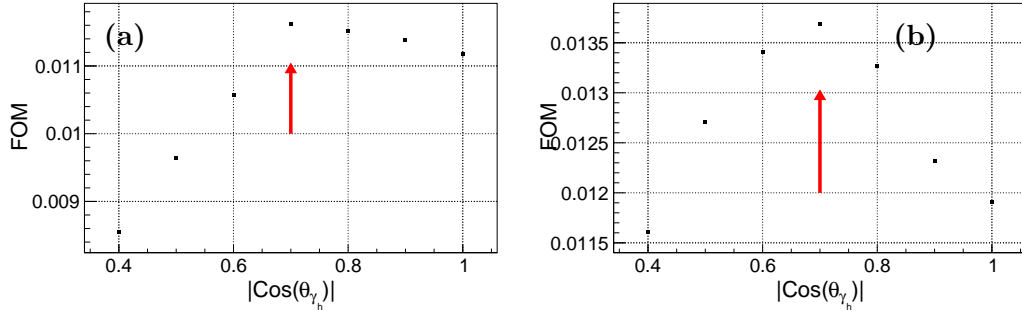


FIG. 12. The FOM parameter distributions of $|\cos(\theta_{\gamma_h})|$ from MC simulated $e^+e^- \rightarrow \gamma_h \chi_{b0} (\rightarrow \gamma_l \Upsilon(1S))$ of (a) $1 + \cos^2(\theta_{\gamma_h})$ distribution and (b) $1 - \cos^2(\theta_{\gamma_h})$ distribution at $\sqrt{s} = 10.746$ GeV.

169
170
171

generic MC samples. Table VI shows the values of the FOM parameters for different $\Upsilon(1S)$ signal regions, and we require the mass of lepton pair larger than 9.44 GeV and smaller than 9.49 GeV according to the maximum value of FOM.

172
173

- We perform a four momentum constraint (4C) fit using KFIT for $\gamma\gamma l^+l^-$ final states to suppress the background with one or more additional or missing particles in the final

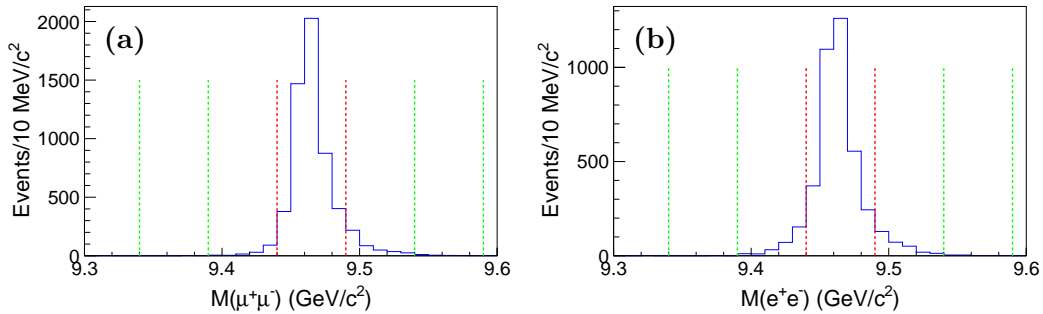


FIG. 13. Distributions of lepton pair mass from signal MC samples for (a) $\mu^+\mu^-$ final state and (b) e^+e^- final state at $\sqrt{s} = 10.746$ GeV.

TABLE VI. The values of FOM parameters with different $\Upsilon(1S)$ mass regions.

$M(l^+l^-)$ region (GeV/ c^2)	FOM value
[9.45, 9.48]	0.0062
[9.44, 9.49] (nominal)	0.0065
[9.43, 9.50]	0.0063
[9.42, 9.51]	0.0062
[9.41, 9.52]	0.0061
[9.40, 9.53]	0.0060
[9.39, 9.54]	0.0059

174 states and improve the mass resolutions, where four-momenta of the final state system
 175 are constrained to the initial e^+e^- C.M. system. Specially, we use the $(M(\gamma\Upsilon(1S))$
 176 - $M(\Upsilon(1S)) + 9.46)$ instead of $M(\gamma\Upsilon(1S))$ to draw the invariant mass distribution
 177 of $\gamma\Upsilon(1S)$ to improve the mass resolution. If there are multiple combinations in one
 178 event, the one with the smallest 4C fit χ^2 is retained. Here the percentage of multiple
 179 candidates per event is 2.2% in data. We optimize the χ_{4C}^2 by maximizing the FOM
 180 parameter $\frac{\varepsilon}{\frac{3}{2} + \sqrt{B}}$ [41], where ε is the efficiency B is the number of background events
 181 from normalized generic MC samples.

182 Figure 14 shows the FOM parameter distribution of 4C fit χ^2 and the distributions
 183 of 4C fit χ^2 from signal MC and data samples at $\sqrt{s} = 10.746$ GeV. The red arrows
 184 show the requirement of $\chi_{4C}^2 < 30$.

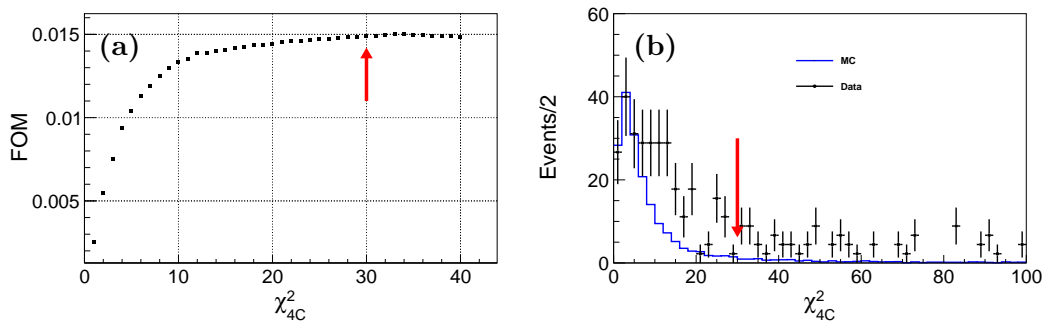


FIG. 14. The (a) FOM parameter distribution of 4C fit χ_{4C}^2 and (b) 4C fit χ_{4C}^2 distributions from signal MC and data samples at $\sqrt{s} = 10.746$ GeV.

185 3.2. The study of signal MC samples

186

188 After applying all selection requirements, the invariant mass distributions of $\gamma\mathcal{Y}(1S)$ from
 189 signal MC samples in $e^+e^- \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$) at $\sqrt{s} = 10.653, 10.701, 10.746,$ and 10.804
 190 GeV are shown in Figs. 15, 16, and 17, respectively. Here we use a Crystal Ball (CB)
 191 function ($\mathcal{F}_{\text{CB}}(M; \mu, \sigma, \alpha, n)$) [42] combined with a Gaussian function to describe the χ_{bJ}
 192 signal shapes. All parameters in CB and Gaussian functions are free.

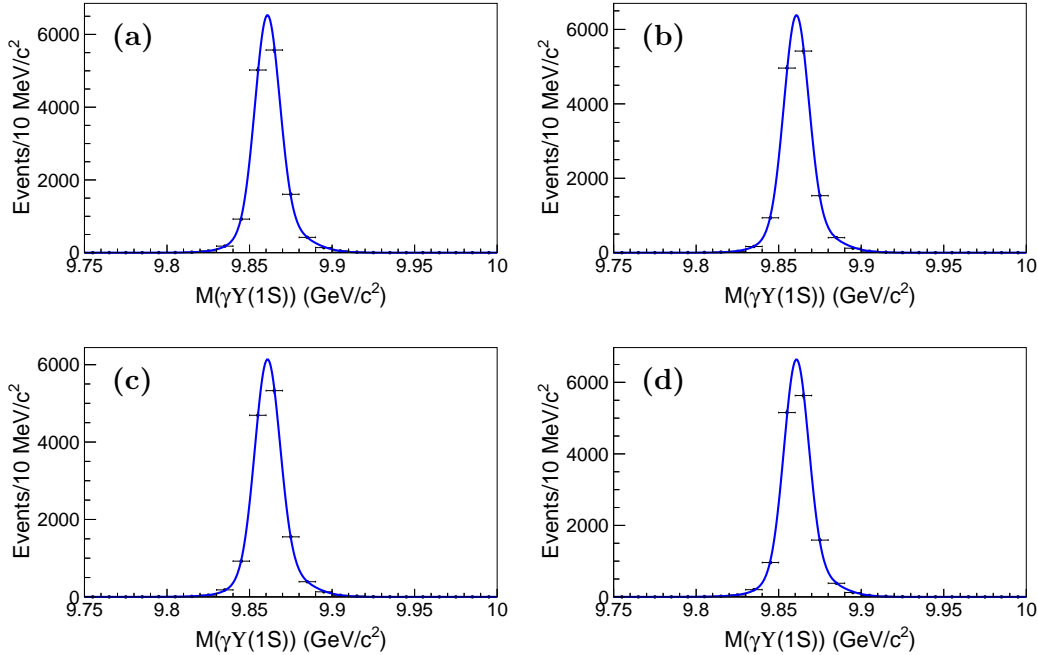


FIG. 15. The invariant mass spectra of $\gamma\mathcal{Y}(1S)$ from MC simulated $e^+e^- \rightarrow \gamma\chi_{b0}$ at (a) $\sqrt{s} = 10.653$ GeV, (b) $\sqrt{s} = 10.701$ GeV, (c) $\sqrt{s} = 10.746$ GeV, and (d) $\sqrt{s} = 10.804$ GeV, respectively.

193 Table VII and table VIII show the reconstruction efficiency from signal MC samples from
 194 e^+e^- mode and $\mu^+\mu^-$ mode, respectively.

TABLE VII. Reconstruction efficiency (%) from signal MC samples of e^+e^- mode.

Energy (GeV)	$\gamma\chi_{b0}$	$\gamma\chi_{b1}$	$\gamma\chi_{b2}$
10.653	21.5	22.3	21.2
10.701	21.4	21.7	21.9
10.746	21.6	21.2	21.8
10.804	22.2	22.4	22.8

3.3. The study of generic MC samples

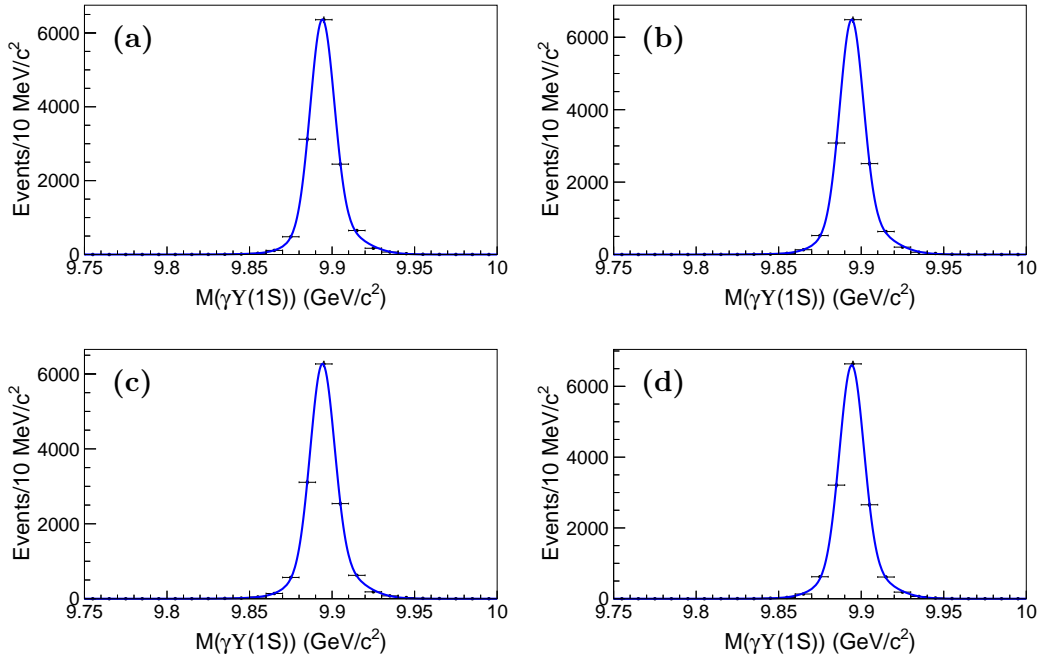


FIG. 16. The invariant mass spectra of $\gamma\mathcal{Y}(1S)$ from MC simulated $e^+e^- \rightarrow \gamma\chi_{b1}$ at (a) $\sqrt{s} = 10.653$ GeV, (b) $\sqrt{s} = 10.701$ GeV, (c) $\sqrt{s} = 10.746$ GeV, and (d) $\sqrt{s} = 10.804$ GeV, respectively.

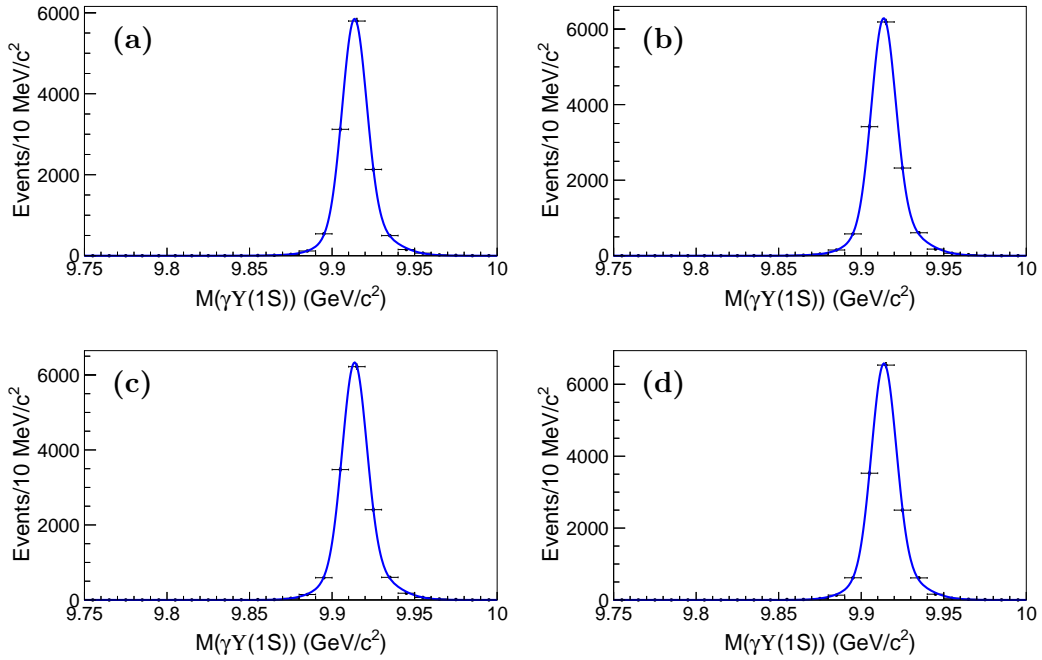


FIG. 17. The invariant mass spectra of $\gamma\mathcal{Y}(1S)$ from MC simulated $e^+e^- \rightarrow \gamma\chi_{b2}$ at (a) $\sqrt{s} = 10.653$ GeV, (b) $\sqrt{s} = 10.701$ GeV, (c) $\sqrt{s} = 10.746$ GeV, and (d) $\sqrt{s} = 10.804$ GeV, respectively.

TABLE VIII. Reconstruction efficiency (%) from signal MC samples of $\mu^+\mu^-$ mode.

Energy (GeV)	$\gamma\chi_{b0}$	$\gamma\chi_{b1}$	$\gamma\chi_{b2}$
10.653	36.1	35.3	34.1
10.701	34.6	35.1	34.1
10.746	35.6	36.1	35.8
10.804	36.1	36.6	36.4

197 After applying all of the above selection criteria, Figure 18 shows the invariant mass
 198 distributions of $\gamma\Upsilon(1S)$ from normalized generic MC samples at $\sqrt{s} = 10.653, 10.701, 10.746,$
 199 and 10.804 GeV, respectively. The remaining backgrounds are mainly from Bhabha and di-
 200 muon processes. No peaking backgrounds are found.

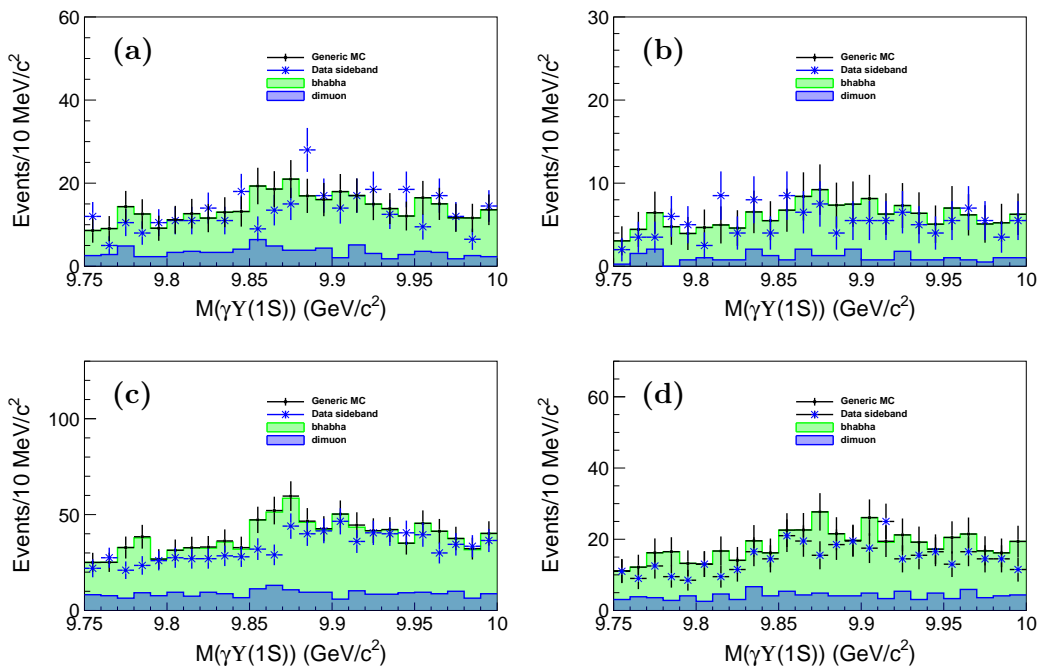


FIG. 18. The invariant mass spectra of $\gamma\Upsilon(1S)$ from normalized generic MC samples and normalized $\Upsilon(1S)$ mass sidebands in data at (a) $\sqrt{s} = 10.653$ GeV, (b) $\sqrt{s} = 10.701$ GeV, (c) $\sqrt{s} = 10.746$ GeV, and (d) $\sqrt{s} = 10.804$ GeV. The black points with error bars are generic MC. The blue histograms show the dimuon process. And the green histograms show the Bhabha process. The blue points with error bars are $\Upsilon(1S)$ mass sidebands in data.

201 3.4. Signal extraction

202 To study the strategies of the fits to the invariant mass spectra of $\gamma\Upsilon(1S)$, we combine
 203 the data sideband samples and signal MC samples to expected signal yields.

204 The unbinned extended maximum-likelihood fits are performed. Here, we simultaneous
 205 fit $\Upsilon(1S) \rightarrow \mu^+\mu^-$ and $\Upsilon(1S) \rightarrow e^+e^-$ mode. In the simultaneous fit, the ratio of yields

206 between $\mu^+\mu^-$ and e^+e^- channel is fixed according to the efficiency. The χ_{bJ} signal shape
 207 is described by a combination of Crystal Ball (CB) and Gaussian functions. All of the
 208 parameters in CB and Gaussian function are fixed according to signal MC simulations. A
 209 second order Chebyshev polynomial will be used to describe the background. The parameters
 210 in the polynomial function are floated.

211 The stability of the fits with this procedure has been tested. Assuming the Born cross
 212 section for $e^+e^- \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$) at $\sqrt{s} = 10.653, 10.701, 10.746,$ and 10.804 GeV of
 213 1 pb, we generate signal MC samples. Figure 19 shows the combined samples and the fit
 214 results at $\sqrt{s} = 10.653, 10.701, 10.746,$ and 10.804 GeV. The comparisons between inputs
 215 and outputs on the Born cross section are listed in Table IX. The input and output results
 216 on the signal yields are consistent within $\pm 1\sigma$.

TABLE IX. Comparisons of the values of inputs and outputs on the Born cross section (pb) from
 the fits at $\sqrt{s} = 10.653, 10.701, 10.746,$ and 10.804 GeV.

Energy (GeV)		$\gamma\chi_{b0}$	$\gamma\chi_{b1}$	$\gamma\chi_{b2}$
10.653	Input	1.0	1.0	1.0
	Output	5.1 ± 4.3	0.9 ± 0.3	1.3 ± 0.5
10.701	Input	1.0	1.0	1.0
	Output	4.3 ± 5.3	1.2 ± 0.4	0.7 ± 0.7
10.746	Input	1.0	1.0	1.0
	Output	2.4 ± 2.3	1.1 ± 0.2	1.1 ± 0.2
10.804	Input	1.0	1.0	1.0
	Output	1.5 ± 1.5	0.9 ± 0.2	0.8 ± 0.3

217 3.5. The Study of data samples

218 After the application of the above requirements, the $M(l^+l^-)$ distribution from a com-
 219 bined $\sqrt{s} = 10.653, 10.701, 10.746,$ and 10.804 GeV data sample is shown in Fig. 20. The red
 220 dash lines in Fig. 20 show the $\Upsilon(1S)$ signal region ($9.44 \text{ GeV}/c^2 < M(l^+l^-) < 9.49 \text{ GeV}/c^2$).
 221 No clear $\Upsilon(1S)$ signal is observed. With the requirement of the $\Upsilon(1S)$ signal region, the
 222 $\gamma\Upsilon(1S)$ invariant mass distributions from data samples at $\sqrt{s} = 10.653, 10.701, 10.746,$ and
 223 10.804 GeV are shown in Fig. 21. The $\gamma\Upsilon(1S)$ invariant mass distributions from a combined
 224 $\sqrt{s} = 10.653, 10.701, 10.746,$ and 10.804 GeV data sample are shown in Fig. 22.

225 We perform unbinned extended maximum-likelihood fits to $\gamma\Upsilon(1S)$ invariant mass distri-
 226 butions to extract signal yields. Here, we simultaneously fit $\Upsilon(1S) \rightarrow \mu^+\mu^-$ and $\Upsilon(1S) \rightarrow$
 227 e^+e^- modes. In the simultaneous fit, the ratio of yields between $\mu^+\mu^-$ and e^+e^- channels is
 228 fixed according to their respective reconstruction efficiencies (we assume the same branching
 229 fraction of $(2.46 \pm 0.04) \times 10^{-2}$ for $\Upsilon(1S) \rightarrow e^+e^-$ and $\Upsilon(1S) \rightarrow \mu^+\mu^-$ [43]). The probability
 230 density function (PDF) of χ_{bJ} signal is modeled as a sum of a Crystal Ball function [42] and
 231 a Gaussian function, sharing a common mean. The mean is fixed to the nominal χ_{bJ} mass
 232 value from Particle Data Group [43], while the remaining parameters are derived from simu-
 233 lation. To account for data-simulation discrepancies, the photon energy in simulated events
 234 is smeared according to the measured photon energy resolution scale factor, 1.01 ± 0.01 ,

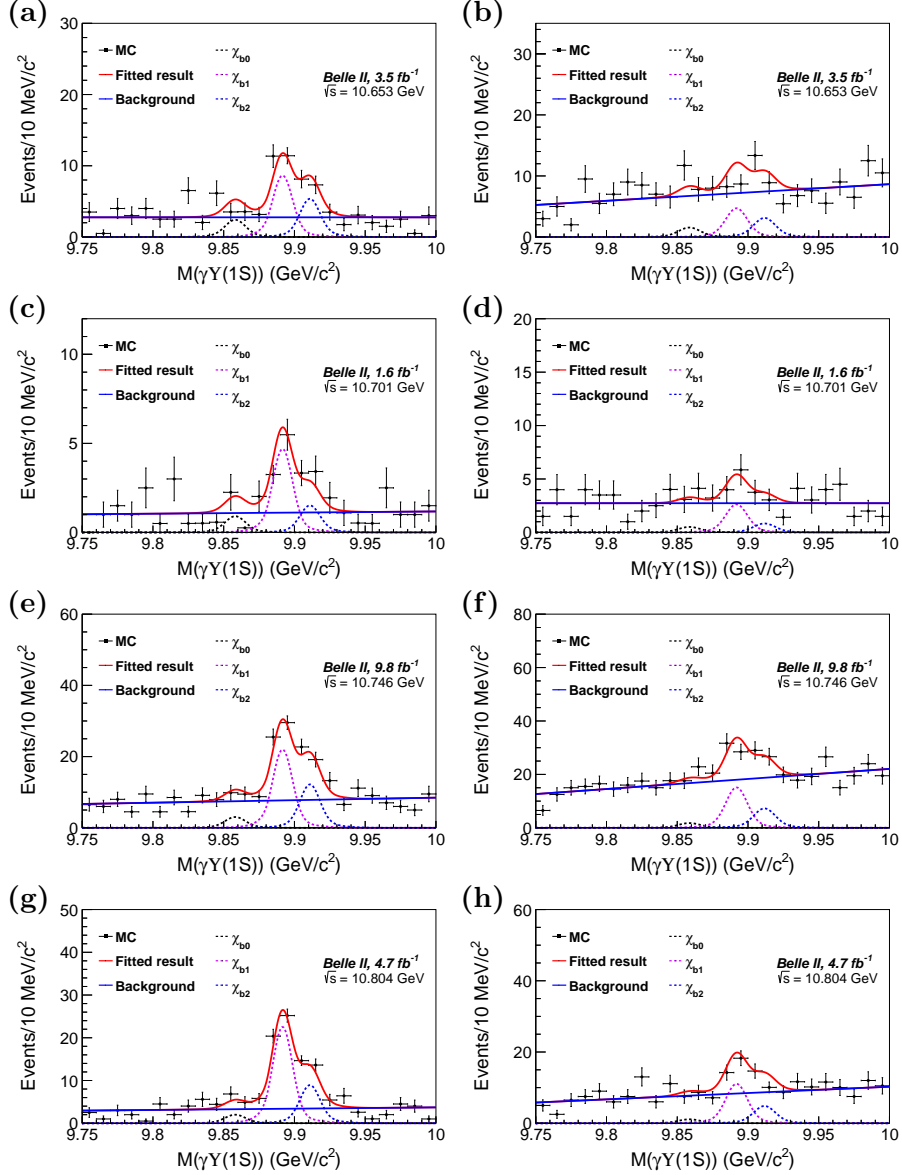


FIG. 19. The simultaneous fitted results to $M(\gamma\Upsilon(1S))$ distributions for combined signal MC samples and data sideband samples in (left) $\mu^+\mu^-$ and (right) e^+e^- modes at $\sqrt{s} =$ (a, b) 10.653, (c, d) 10.701, (e, f) 10.746, and (g, h) 10.804 GeV, respectively. Here the dots with error bars are the combined signal MC samples and data sideband samples. The red curves show the fitted results. The blue cyan curves show the total backgrounds. The black, purple, and blue dashed curves show the χ_{b0} , χ_{b1} , and χ_{b2} signals.

235 obtained from the $\pi^0 \rightarrow \gamma\gamma$ control sample [51]. A second order Chebyshev polynomial will
 236 be used to describe the background. The parameters in the polynomial function are floated.
 237 We calculate the signal significance according to

$$\sqrt{-2 \ln(\mathcal{L}_0/\mathcal{L}_{\max})}, \quad (7)$$

238 where \mathcal{L}_0 and \mathcal{L}_{\max} are the maximized likelihoods without and with the signal, respectively.
 239 The signal yields at each energy point are listed in Table X. No significant χ_{bJ} ($J = 0, 1,$

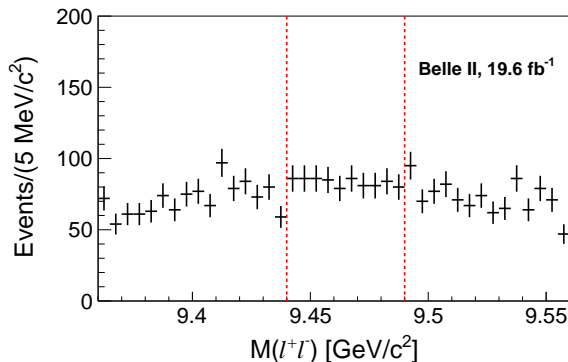


FIG. 20. The invariant mass spectrum of l^+l^- from a combined $\sqrt{s} = 10,653, 10.701, 10.746,$ and 10.804 GeV data sample. The red dash lines show the signal region ($9.44 \text{ GeV}/c^2 < M(l^+l^-) < 9.49 \text{ GeV}/c^2$).

240 2) signal was found. We thus compute 90% C.L. upper limits x^{UL} on the signal yields and
 241 Born cross sections by solving the equation

$$\int_0^{x^{\text{UL}}} \mathcal{L}(x) dx / \int_0^{+\infty} \mathcal{L}(x) dx = 0.90, \quad (8)$$

242 where x is the assumed signal yield or Born cross section, and $\mathcal{L}(x)$ is the corresponding
 243 maximized likelihood of the fit to the assumption. The upper limits on the signal yields
 244 including systematic uncertainties (discussed below) are listed in Table X.

245 The upper limits on the Born cross sections are calculated according to Eq.(1), the
 246 values used to determine the Born cross sections as well as the upper limits on the Born
 247 cross sections including systematic uncertainties are listed in Table X.

248 4. SYSTEMATIC UNCERTAINTY

249 The systematic uncertainties in the cross-section measurements of $e^+e^- \rightarrow \gamma\chi_{bJ}$ include
 250 detection efficiency uncertainties (tracking, particle ID, photon reconstruction), trigger sim-
 251 ulation, MC statistics, branching fractions of intermediate states, integrated luminosity, an-
 252 gular distributions, beam-energy calibration, the radiative correction factor, the signal PDF
 253 parameters, and fit model. The systematic uncertainties from the signal PDF parameters
 254 and fit model are additive. The other sources of systematic uncertainties are multiplicative,
 255 which are listed in Table XII.

- 256 • Detection efficiency

- 257 (a) Tracking

258 For a lepton track, a systematic uncertainty of 0.3% per track is included [44].

- 259 (b) Particle ID

260 In our event selection, we only require at least one lepton to be identified. Con-
 261 sequently, the lepton identification efficiency is so high ($> 99\%$) that the corre-
 262 sponding systematic uncertainty is negligible.

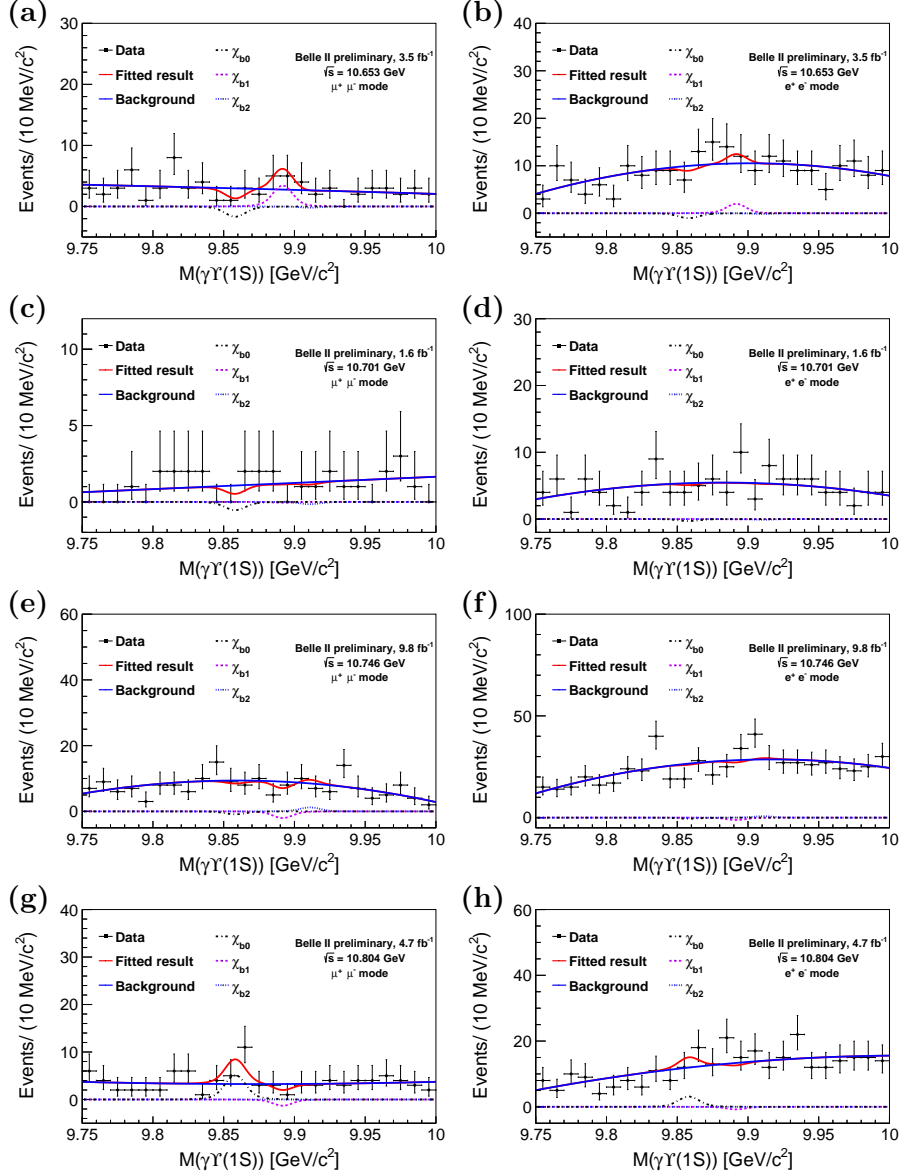


FIG. 21. The simultaneous fitted results to $M(\gamma\Upsilon(1S))$ distributions from data samples in (left) $\mu^+\mu^-$ and (right) e^+e^- modes at $\sqrt{s} =$ (a, b) 10.653, (c, d) 10.701, (e, f) 10.746, and (g, h) 10.804 GeV, respectively. Here the dots with error bars are the data samples. The red curves show the fitted results. The blue cyan curves show the total backgrounds. The black, purple, and blue dashed curves show the χ_{b0} , χ_{b1} , and χ_{b2} signals.

(c) Photon reconstruction

The Data/MC ratio of photon reconstruction efficiency of the Belle II calorimeter using di-muon events has been studied [49]. By weighting events in different regions of (θ, ϕ, E) , we find the Data/MC ratio is $(97.0 \pm 1.1)\%$ for γ_l and $(98.2 \pm 0.7)\%$ for γ_h . The corrections are applied to the efficiencies of signal MC. The uncertainties on corrections are included as systematic uncertainties due to photon reconstruction.

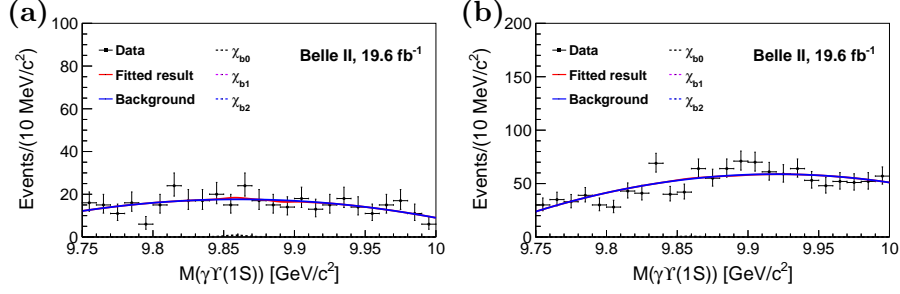


FIG. 22. The simultaneous fitted results to $M(\gamma\Upsilon(1S))$ distributions from a combined $\sqrt{s} = 10.653, 10.701, 10.746,$ and 10.804 GeV data sample in (a) $\mu^+\mu^-$ and (b) e^+e^- modes. Here the dots with error bars are the data samples. The red curves show the fitted results. The blue cyan curves show the total backgrounds. The black, purple, and blue dashed curves show the χ_{b0} , χ_{b1} , and χ_{b2} signals.

TABLE X. The values used to determine the Born cross sections for $e^+e^- \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$) at $\sqrt{s} = 10.653, 10.701, 10.746,$ and 10.804 GeV, respectively. Here \mathcal{L} is the integrated luminosity, N^{fit} is the sum of signal yields of $\Upsilon(1S) \rightarrow e^+e^-$ and $\Upsilon(1S) \rightarrow \mu^+\mu^-$ modes with purely statistical uncertainty, N^{UL} is the upper limit at 90% C.L. on the signal yield, $\varepsilon_{e^+e^-}$ and $\varepsilon_{\mu^+\mu^-}$ are reconstruction efficiencies for $\Upsilon(1S) \rightarrow e^+e^-$ and $\Upsilon(1S) \rightarrow \mu^+\mu^-$ modes, $\mathcal{B}_{e^+e^-}^{\text{int}}$ and $\mathcal{B}_{\mu^+\mu^-}^{\text{int}}$ are the product of the branching fractions of the intermediate states for $\Upsilon(1S) \rightarrow e^+e^-$ and $\Upsilon(1S) \rightarrow \mu^+\mu^-$ modes, $|1 - \Pi|^2$ is the vacuum polarization factor, $1 + \delta_{\text{ISR}}$ is the radiative correction factor, syst is the systematic uncertainty, and $\sigma_{\text{Born}}^{\text{UL}}$ is the upper limit at 90% C.L. on the Born cross section.

\sqrt{s} (GeV)	\mathcal{L} (fb $^{-1}$)	Channel	N^{fit}	N^{UL}	$\varepsilon_{e^+e^-}$	$\varepsilon_{\mu^+\mu^-}$	$\mathcal{B}_{e^+e^-}^{\text{int}}$	$\mathcal{B}_{\mu^+\mu^-}^{\text{int}}$	$ 1 - \Pi ^2$	$1 + \delta_{\text{ISR}}$	syst (%)	$\sigma_{\text{Born}}^{\text{UL}}$ (pb)
10.653	3.512	$\gamma\chi_{b0}$	-5.5 ± 4.9	8.5	0.215	0.361	0.00046	0.00048	0.930	0.901	15.5	9.18
		$\gamma\chi_{b1}$	10.4 ± 7.0	21.5	0.223	0.353	0.00841	0.00873	0.930	0.898	9.0	1.28
		$\gamma\chi_{b2}$	-0.7 ± 5.5	10.5	0.212	0.341	0.00430	0.00446	0.930	0.896	8.9	1.28
10.701	1.632	$\gamma\chi_{b0}$	-1.7 ± 2.8	7.0	0.214	0.346	0.00046	0.00048	0.931	0.905	15.5	16.68
		$\gamma\chi_{b1}$	-0.2 ± 3.7	8.6	0.217	0.351	0.00841	0.00873	0.931	0.903	9.0	1.11
		$\gamma\chi_{b2}$	-0.5 ± 3.4	8.0	0.219	0.341	0.00430	0.00446	0.931	0.901	8.9	2.06
10.746	9.818	$\gamma\chi_{b0}$	-2.9 ± 9.5	13.9	0.216	0.356	0.00046	0.00048	0.931	0.909	15.5	5.37
		$\gamma\chi_{b1}$	-6.1 ± 8.5	12.1	0.212	0.361	0.00841	0.00873	0.931	0.906	9.0	0.26
		$\gamma\chi_{b2}$	3.8 ± 9.7	19.1	0.218	0.358	0.00430	0.00446	0.931	0.904	8.9	0.79
10.804	4.689	$\gamma\chi_{b0}$	16.7 ± 8.1	30.1	0.222	0.361	0.00046	0.00048	0.932	0.914	15.5	23.77
		$\gamma\chi_{b1}$	-4.0 ± 5.6	10.0	0.224	0.366	0.00841	0.00873	0.932	0.911	9.0	0.43
		$\gamma\chi_{b2}$	-0.1 ± 5.7	11.2	0.228	0.364	0.00430	0.00446	0.932	0.909	8.9	0.94

270
271

We add individual uncertainties in quadrature to obtain the final uncertainties for $\mu^+\mu^-$ and e^+e^- modes related to the detection efficiency, which is 1.9%.

272

- Trigger simulation

273
274
275
276
277

We chose L1 CDC triggers of hie, c4, lml1, and ECL triggers of fyo, fy30, stt for our analysis. We study the uncertainty of these trigger bits by the Bhabha process and di-muon process. The efficiencies from data and MC are 99.8% and 99.6% for dimuon process, and 98.6% and 96.5% for Bhabha process. As a result, we include 0.2% uncertainty for $\mu\mu$ final state and 2.1% uncertainty for ee final state due to trigger simulation.

278 The above two sources of systematic uncertainties (Detection efficiency and Trigger sim-
 279 ulation) for $\mu^+\mu^-$ and e^+e^- modes are different, thus we combine them to obtain the total
 280 systematic uncertainty by taking into account the detection efficiencies for $\mu^+\mu^-$ and e^+e^-
 281 modes as the weight factors.

- 282 • MC statistics

283 The statistical error in the MC efficiency can be calculated using $\Delta_\varepsilon = \sqrt{(\varepsilon(1-\varepsilon))/N}$,
 284 where ε is the reconstruction efficiency, and N is the total number of generated events. The
 285 uncertainty in the MC statistics is estimated by $\Delta_\varepsilon/\varepsilon$, which is 1.0%.

- 286 • Branching fractions of intermediate states

287 The uncertainties of branching fractions for $\chi_{b0} \rightarrow \gamma\mathcal{Y}(1S)$, $\chi_{b1} \rightarrow \gamma\mathcal{Y}(1S)$, $\chi_{b2} \rightarrow$
 288 $\gamma\mathcal{Y}(1S)$, $\mathcal{Y}(1S) \rightarrow e^+e^-$, $\mathcal{Y}(1S) \rightarrow \mu^+\mu^-$ are 13.9%, 5.7%, 5.6%, 4.6%, and 2.0%. The
 289 total uncertainties on branching fractions in $e^+e^- \rightarrow \gamma\chi_{b0}$, $e^+e^- \rightarrow \gamma\chi_{b1}$, and $e^+e^- \rightarrow \gamma\chi_{b2}$
 290 are 14.6%, 7.3%, and 7.2%, respectively.

- 291 • Integrated luminosity

292 Belle II measures the luminosity at 1.2% precision using Bhabha, digamma, and mupair
 293 events [52].

- 294 • Angular distribution for $V \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$)

295 We require the $|\cos(\theta_{\gamma_h})| < 0.7$ to suppress Bhabha background events. We change the
 296 uniform angular distribution to the distribution of $1 \pm \cos^2\theta$ for $V \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$)
 297 and take the maximum difference of efficiency (4 %) as the uncertainty of different angular
 298 distributions.

- 299 • Beam-energy calibration

300 Table XI shows the measurements for the e^+e^- C.M. energy in data. We change the
 301 collision energies according to Table XI in the 4C kinematic fit and take the largest dif-
 302 ference (C.M. energy(\pm uncertainty) relative to nominal C.M. energy) in efficiency as the
 303 uncertainty, which is less than 1%. We will take the difference of 1% in the signal efficiency
 304 as an uncertainty due to the beam-energy calibration.

TABLE XI. Measurement results for the energy of the scan data samples. [26]

point	Collected luminosity (fb^{-1})	E_{cm} (MeV)
1	3.521	$10653.30 \pm 0.71 \pm 0.89 \pm 0.50$
2	1.633	$10700.90 \pm 0.61 \pm 0.14 \pm 0.50$
3	9.818	$10746.30 \pm 0.46 \pm 0.15 \pm 0.50$
4	4.690	$10804.50 \pm 0.65 \pm 0.25 \pm 0.50$

- 305 • The radiative correction factor

306 To calculate the radiative correction factor of $e^+e^- \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$), we assume the
 307 dressed cross section follows the $1/s$ line shape. We change the $1/s$ to $1/s^2$ [50] and take
 308 the difference on the radiative correction factor as the systematic uncertainty.

309 All the multiplicative uncertainties in the measurements of $\sigma_{\text{Born}}(e^+e^- \rightarrow \gamma\chi_{bJ})$ are added
 310 in quadrature to obtain the total systematic uncertainty, which are summarized in Table XII.

- 311 • signal PDF parameters

312 We considered the following systematic uncertainties associated with the signal PDF
 313 parameters. We varied the mean of the signal PDF according to the uncertainties of the χ_{bJ}
 314 masses from Particle Data Group [43], which are ± 0.42 MeV, ± 0.26 MeV, and ± 0.26 MeV
 315 for χ_{b0} , χ_{b1} , and χ_{b2} , respectively. To account for data-simulation discrepancies, the photon
 316 energy in simulated events is smeared according to the measured photon energy resolution
 317 scale factor, 1.01 ± 0.01 , obtained from the $\pi^0 \rightarrow \gamma\gamma$ control sample [51]. We repeated the
 318 fit while varying the photon energy resolution scale factor within its uncertainty.

TABLE XII. Relative multiplicative systematic uncertainties of the measurements of the Born cross sections for $e^+e^- \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$) at $\sqrt{s} = 10.653, 10.701, 10.746$, and 10.804 GeV, respectively.

Sources	$\gamma\chi_{b0}$	$\gamma\chi_{b1}$	$\gamma\chi_{b2}$
Branching fractions	14.6	7.3	7.2
Detection efficiency	1.9	1.9	1.9
Trigger simulation	1.7	1.7	1.7
MC statistics	1.0	1.0	1.0
Integrated luminosity	1.2	1.2	1.2
Angular distributions	3.9	3.9	4.0
Beam-energy calibration	1.0	1.0	1.0
Radiative correction factor	1.5	1.4	1.4
Sum	15.5	9.0	8.9

- 319 • Fit model

320 The systematic uncertainties associated with the fitting procedure are estimated by
 321 changing the order of the background polynomial and the range of the fit, and comparing the
 322 fit results without a χ_{b0} component due to the small branching fraction for $\chi_{b0} \rightarrow \gamma\mathcal{T}(1S)$.

323 In the calculations of upper limits, the systematic uncertainties are taken into account in
 324 two steps. First, when we study the additive systematic uncertainty from from the signal
 325 PDF parameters and fit model (as described above), we calculate the upper limit for each
 326 possible case and take the most conservative upper limit at 90% C.L. on the number of
 327 signal events. Then, the likelihood with the most conservative upper limit is convolved
 328 with a Gaussian function whose width is equal to the corresponding total multiplicative
 329 uncertainty.

330 **5. SUMMARY**

331 In summary, we report a search for the $e^+e^- \rightarrow \gamma\chi_{bJ}$ ($J = 0, 1, 2$) processes at $\sqrt{s} =$
 332 10.653, 10.701, 10.746 and 10.804 GeV. We do not find evidence for any process, and set
 333 upper limits at 90% C.L. on the corresponding Born cross sections.

334 The upper limit at 90% C.L. on the Born cross section for $e^+e^- \rightarrow \gamma\chi_{b1}$ at $\sqrt{s} = 10.746$
 335 GeV is 0.26 pb. This value is significantly smaller than the values of (3.6 ± 0.9) pb and
 336 (1.4 ± 0.2) pb for $e^+e^- \rightarrow \omega\chi_{b1}$ and $e^+e^- \rightarrow \pi^+\pi^-\Upsilon(2S)$ at $\sqrt{s} = 10.746$ GeV [6, 24].
 337 Interpretations as an admixture of the conventional $4S$ and $3D$ states predict comparable
 338 branching fractions of 10^{-3} for $\Upsilon(10753) \rightarrow \omega\chi_{b1}$ and $\Upsilon(10753) \rightarrow \pi^+\pi^-\Upsilon(2S)$ [7, 8]. There-
 339 fore, the branching fraction for $\Upsilon(10753) \rightarrow \gamma\chi_{b1}$ is less than 10^{-3} . If the $\Upsilon(10753)$ consists
 340 of a significant D -wave component, the branching fraction for $\Upsilon(10753) \rightarrow \gamma\chi_{b1}$ can reach
 341 the magnitude of 10^{-2} [28, 29]. Therefore, our measurement indicates that the D -wave
 342 component in the $\Upsilon(10753)$ cannot be large.

343 * 22210200014@m.fudan.edu.cn

344 † jiasen@seu.edu.cn

345 ‡ shencp@fudan.edu.cn

- 346 [1] M. Artuso *et al.* [CLEO], Phys. Rev. Lett. **94** (2005), 032001
 347 doi:10.1103/PhysRevLett.94.032001 [arXiv:hep-ex/0411068 [hep-ex]].
- 348 [2] D. Besson *et al.* [CLEO], Phys. Rev. D **74** (2006), 012003 doi:10.1103/PhysRevD.74.012003
 349 [arXiv:hep-ex/0512061 [hep-ex]].
- 350 [3] M. Abumusabh *et al.* [Belle-II], [arXiv:2508.16036 [hep-ex]].
- 351 [4] B. G. Fulsom *et al.* [Belle], Phys. Rev. Lett. **121** (2018) no.23, 232001
 352 doi:10.1103/PhysRevLett.121.232001 [arXiv:1807.01201 [hep-ex]].
- 353 [5] R. Mizuk *et al.* (Belle Collaboration), JHEP **10**, 220 (2019).
- 354 [6] I. Adachi *et al.* (Belle II Collaboration), arXiv:2401.12021.
- 355 [7] Z.Y. Bai, Y.S. Li, Q. Huang, X. Liu, and T. Matsuki, Phys. Rev. D **105**, 074007 (2022).
- 356 [8] Y.S. Li, Z.Y. Bai, Q. Huang, and X. Liu, Phys. Rev. D **104**, 034036 (2021).
- 357 [9] Q. Li, M.S. Liu, Q.F. Lü, L.C. Gui, and X.H. Zhong, Eur. Phys. J. C **80**, 59 (2020).
- 358 [10] B. Chen, A.L. Zhang, and J. He, Phys. Rev. D **101**, 014020 (2020).
- 359 [11] J.F. Giron and R.F. Lebed, Phys. Rev. D **102**, 014036 (2020).
- 360 [12] V. Kher, R. Chaturvedi, N. Devlani, and A.K. Rai, Eur. Phys. J. Plus **137**, 357 (2022).
- 361 [13] Y.S. Li, Z.Y. Bai, and X. Liu, Phys. Rev. D **105**, 114041 (2022).
- 362 [14] W.H. Liang, N. Ikeno, and E. Oset, Phys. Lett. B **803**, 135340 (2020).
- 363 [15] N.Hüsken, R.E. Mitchell, and E.S. Swanson, Phys. Rev. D **106**, 094013 (2022).
- 364 [16] E.V. Beveren and G. Rupp, Prog. Part. Nucl. Phys. **117**, 103845 (2021).
- 365 [17] J.T. Castellà and E. Passemar, Phys. Rev. D **104**, 034019 (2021).
- 366 [18] N. Brambilla, S. Eidelman, C. Hanhart, A. Nefediev, C.P. Shen, C.E. Thomas, A. Vairo, and
 367 C.Z. Yuan, Phys. Rept. **873**, 1 (2020).
- 368 [19] A. Ali, L. Maiani, A.Y. Parkhomenko, and W. Wang, Phys. Lett. B **802**, 135217 (2020).
- 369 [20] P. Bicudo, N. Cardoso, L. Müller, and M. Wagner, Phys. Rev. D **103**, 074507 (2021).
- 370 [21] P. Bicudo, N. Cardoso, L. Müller, and M. Wagner, Phys. Rev. D **107**, 094515 (2023).
- 371 [22] Z. G. Wang, Chin. Phys. C **43**, 123102 (2019).

- 372 [23] A.Y. Parkhomenko and W. Wang, Phys. Part. Nucl. Lett. **20**, 381 (2023).
- 373 [24] I. Adachi *et al.* (Belle II Collaboration), Phys. Rev. Lett. **130**, 091902 (2023).
- 374 [25] X.H. He *et al.* (Belle Collaboration), Phys. Rev. Lett. **113**, 142001 (2014).
- 375 [26] I. Adachi *et al.* [Belle-II], [arXiv:2405.18928 [hep-ex]].
- 376 [27] I. Adachi *et al.* (Belle II Collaboration), Phys. Rev. D **109**, 072013 (2024).
- 377 [28] S. Godfrey and K. Moats, Phys. Rev. D **92**, 054034 (2015).
- 378 [29] J.Z. Wang, Z.F. Sun, X. Liu, and T. Matsuki, Eur. Phys. J. C **78**, 915 (2018).
- 379 [30] M. Ablikim *et al.* (BESIII Collaboration), Phys. Rev. Lett. **118**, 092001 (2017).
- 380 [31] M. Ablikim *et al.* (BESIII Collaboration), Chin. Phys. C **39**, 041001 (2015).
- 381 [32] D. Santel *et al.* (Belle Collaboration), Phys. Rev. D **93**, 011101 (2016).
- 382 [33] S. Actis *et al.* (Working Group on Radiative Corrections and Monte Carlo Generators for Low
383 Energies Collaborations), Eur. Phys. J. C **66**, 585 (2010).
- 384 [34] E.A. Kuraev and V.S. Fadin, Yad. Fiz. **41**, 733 (1985) [Sov. J. Nucl. Phys. **41**, 466 (1985)].
- 385 [35] S. Jadach, B.F.L. Ward, and Z. Wc as, Phys. Rev. D **63**, 113009 (2001); Comput. Phys.
386 Commun. **130**, 260 (2000).
- 387 [36] M. Benayoun, S.I. Eidelman, V.N. Ivanchenko, and Z.K. Silagadze, Mod. Phys. Lett. A **14**,
388 2605 (1999).
- 389 [37] G. Rodrigo, H. Czyż, J. H. Kühn and M. Szopa, Eur. Phys. J. C **24**, 71 (2002).
- 390 [38] D.J. Lange, Nucl. Instr. and Methods Phys. Res. Sect. A **462**, 152 (2001).
- 391 [39] S. Agostinelli *et al.*, Nucl. Instrum. Meth. A **506**, 250 (2003).
- 392 [40] Please see details via this link: <https://cdbweb.sdcc.bnl.gov/GlobalTag?id=2483>.
- 393 [41] G. Punzi, eConf C030908, MODT002 (2003).
- 394 [42] T. Skwarnicki, Ph.D. thesis, Institute for Nuclear Physics, 1986; DESY Report No. DESY
395 F31-86-02, 1986.
- 396 [43] R.L. Workman *et al.* (Particle Data Group), Prog. Theor. Exp. Phys. **2022**, 083C01 (2022)
397 and 2023 update.
- 398 [44] A. Glazov, P. Rados, A. Rostomyan, E. Paoloni, and L. Zani, BELLE2-NOTE-PH-2020-006.
- 399 [45] D. Ferlewicz, M. Milesi, and P. Rocchetti, BELLE2-NOTE-PH-2020-010.
- 400 [46] A. Martini and A.N. Charan, BELLE2-NOTE-TE-2020-003.
- 401 [47] S. Yo, K. Uno, and S. Yo, BELLE2-NOTE-PH-2019-043.
- 402 [48] J. Skorupa, BELLE2-NOTE-TE-2019-016.
- 403 [49] Lu Cao and K. Tackmann, BELLE2-NOTE-TE-2024-002.
- 404 [50] S. Jia *et al.* (Belle Collaboration), Phys. Rev. D **98**, 092015 (2018).
- 405 [51] Q.P. Ji, S. Jia, and C.P. Shen, BELLE2-NOTE-PH-2021-021.
- 406 [52] I. Adachi *et al.* (Belle II Collaboration), arXiv:2407.00965.