



# DARK SECTOR PROSPECTS AT BELLE II

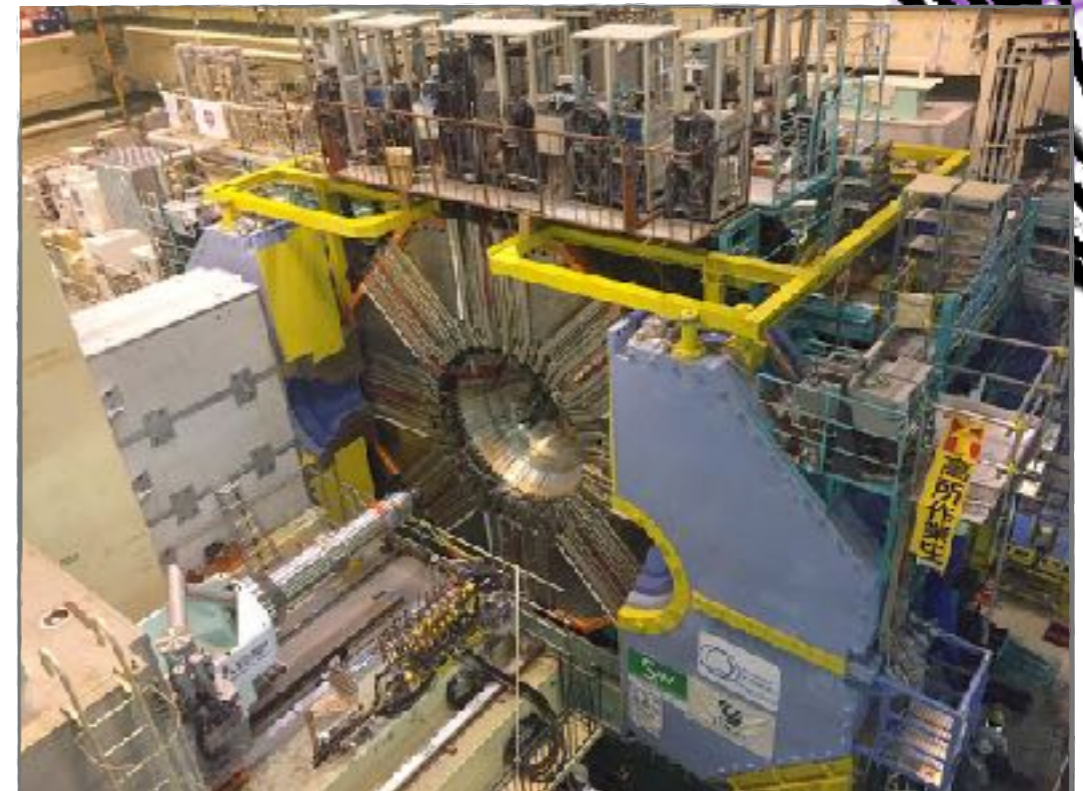
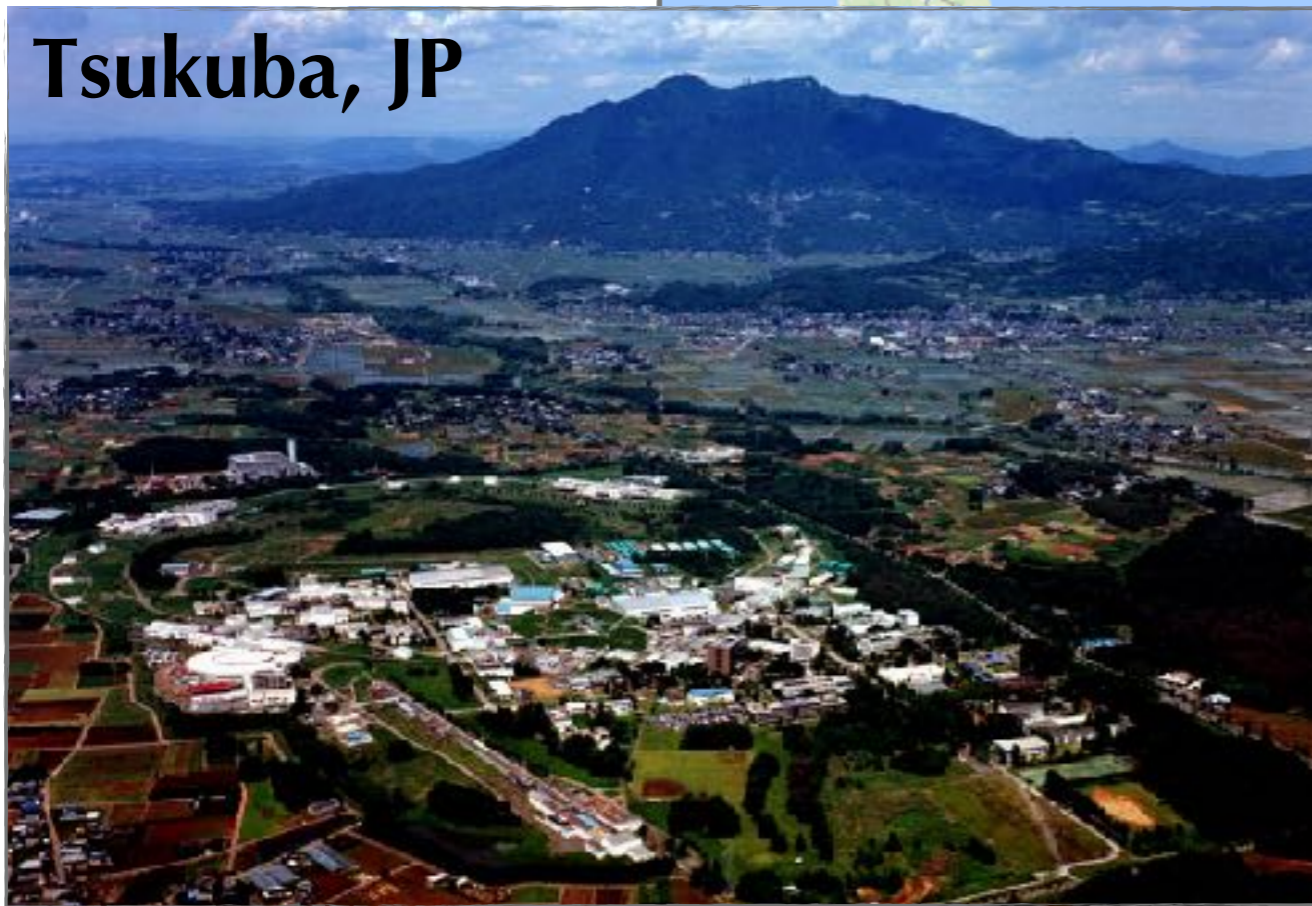
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Caitlin MacQueen ON BEHALF OF THE BELLE II COLLABORATION

Dark Side of the Universe — Daejeon, South Korea — July 10th, 2017

# SUPERKEKB AND BELLE II

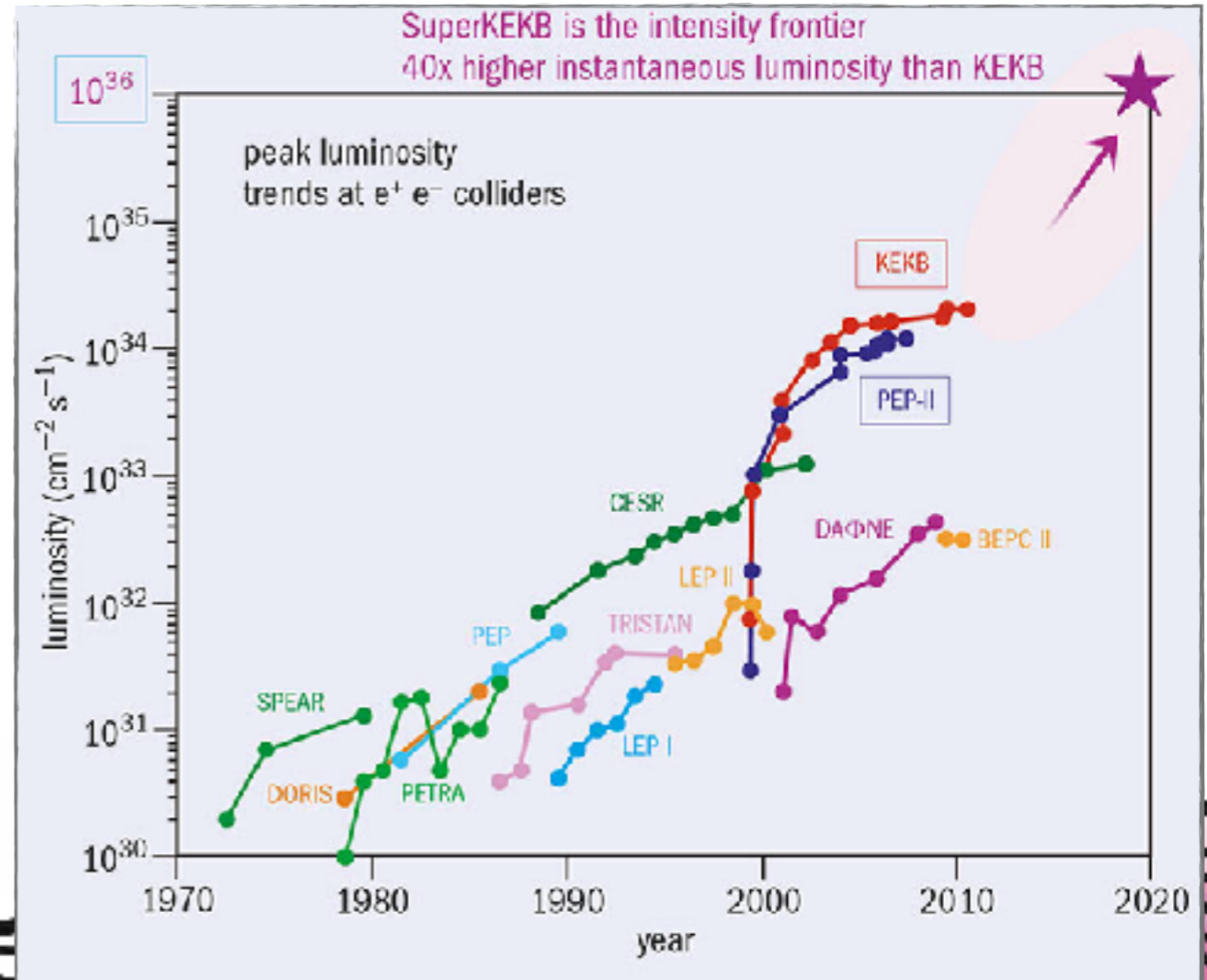
- 700+ scientists
- 100+ institutions
- 24 countries



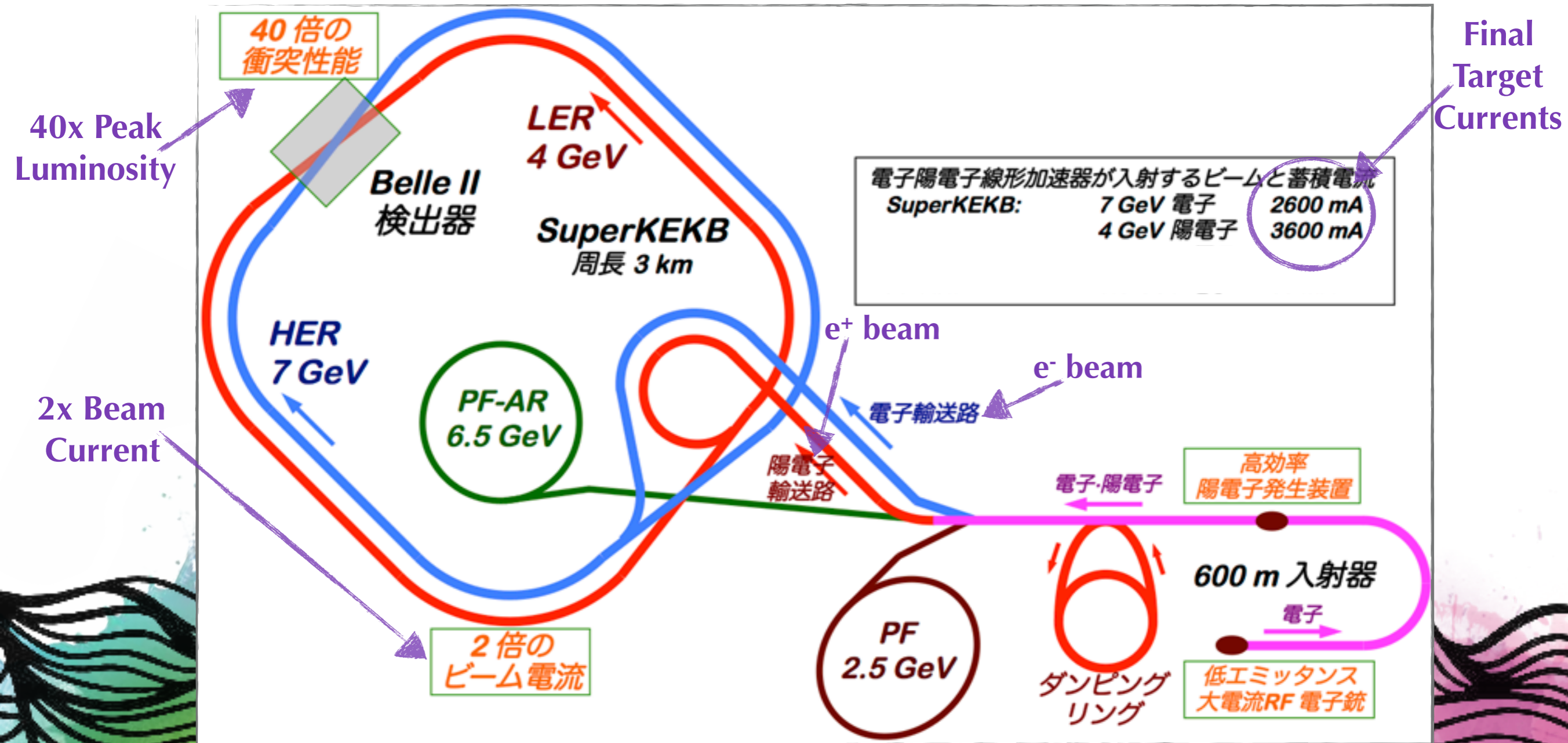
- **40 times** the peak luminosity of KEKB
- 2 times as much current
- 20 times smaller vertical beam size

$$L = \frac{\gamma_{\pm}}{2er_e} \left( 1 + \frac{\sigma_y^*}{\sigma_x^*} \right) \left( \frac{I_{\pm} \zeta_{\pm y}}{\beta_y^*} \right) \left( \frac{R_L}{R_y} \right)$$

Lorentz Factor  $\rightarrow \gamma_{\pm}$   
 Beam Current  $\rightarrow I_{\pm}$   
 Beam-Beam Parameter  $\rightarrow \zeta_{\pm y}$   
 Beam Size Aspect Ratio  $\rightarrow \frac{\sigma_y^*}{\sigma_x^*}$   
 Vertical Beta Function  $\rightarrow \beta_y^*$   
 Geometric Factors  $\rightarrow \frac{R_L}{R_y}$



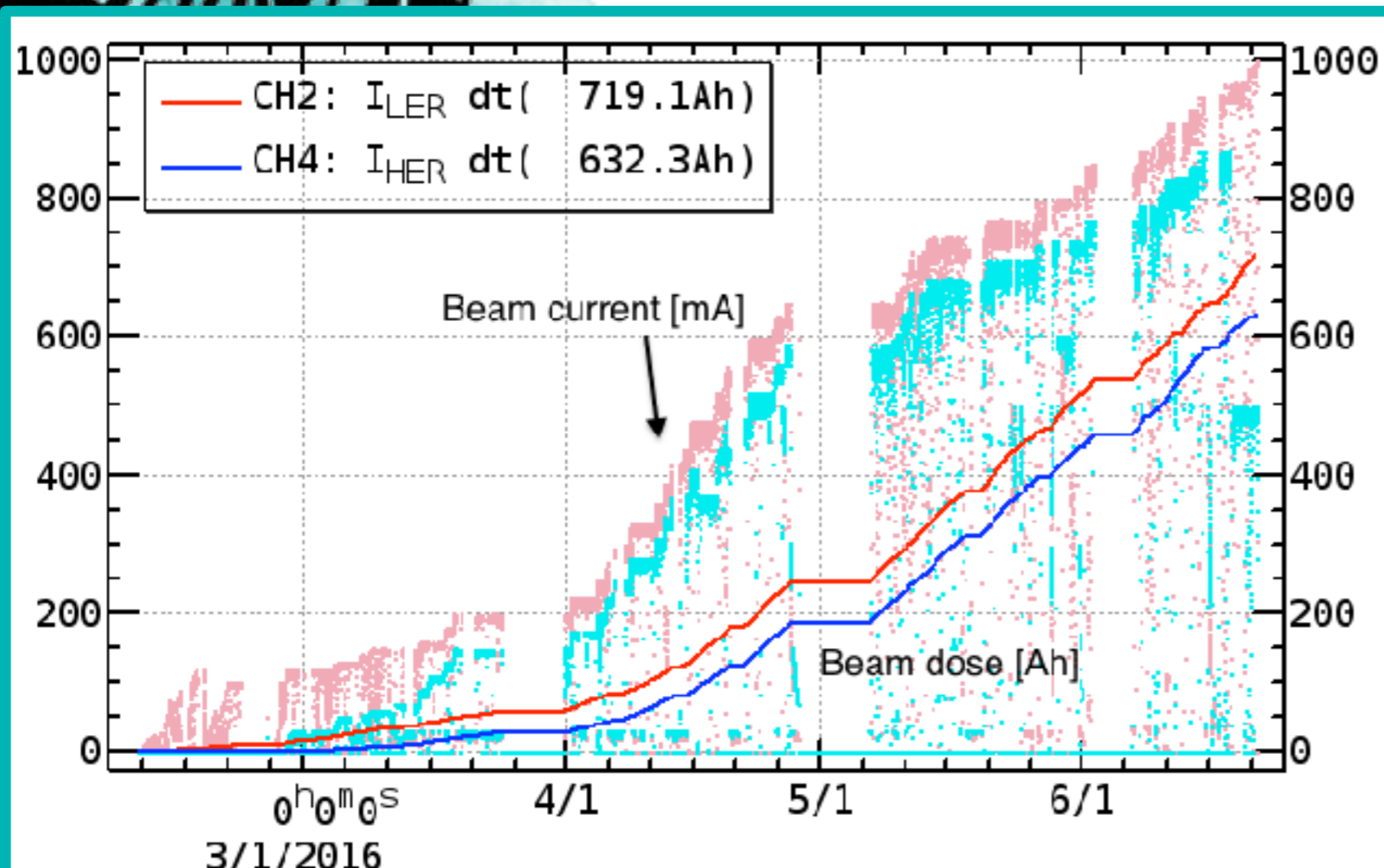
# SUPERKEKB: ON THE INTENSITY FRONTIER





IP - August 2016

- Phase 1: non-collision run
  - February - June, 2016
- Phase 2: **first collisions**
  - Beginning in **February, 2018**
  - Partial detector (1 vertex detector segment), resulting in lower efficiencies
  - **20 +/- 20 fb<sup>-1</sup>** of data collection
- Phase 3: full data-taking period
  - Beginning in late 2018
  - Full detector



5 Months of Operation (2/1 - 6/30)

- Last year, Phase 1 ran — this was the first operation of SuperKEKB.

**Final Target Currents**

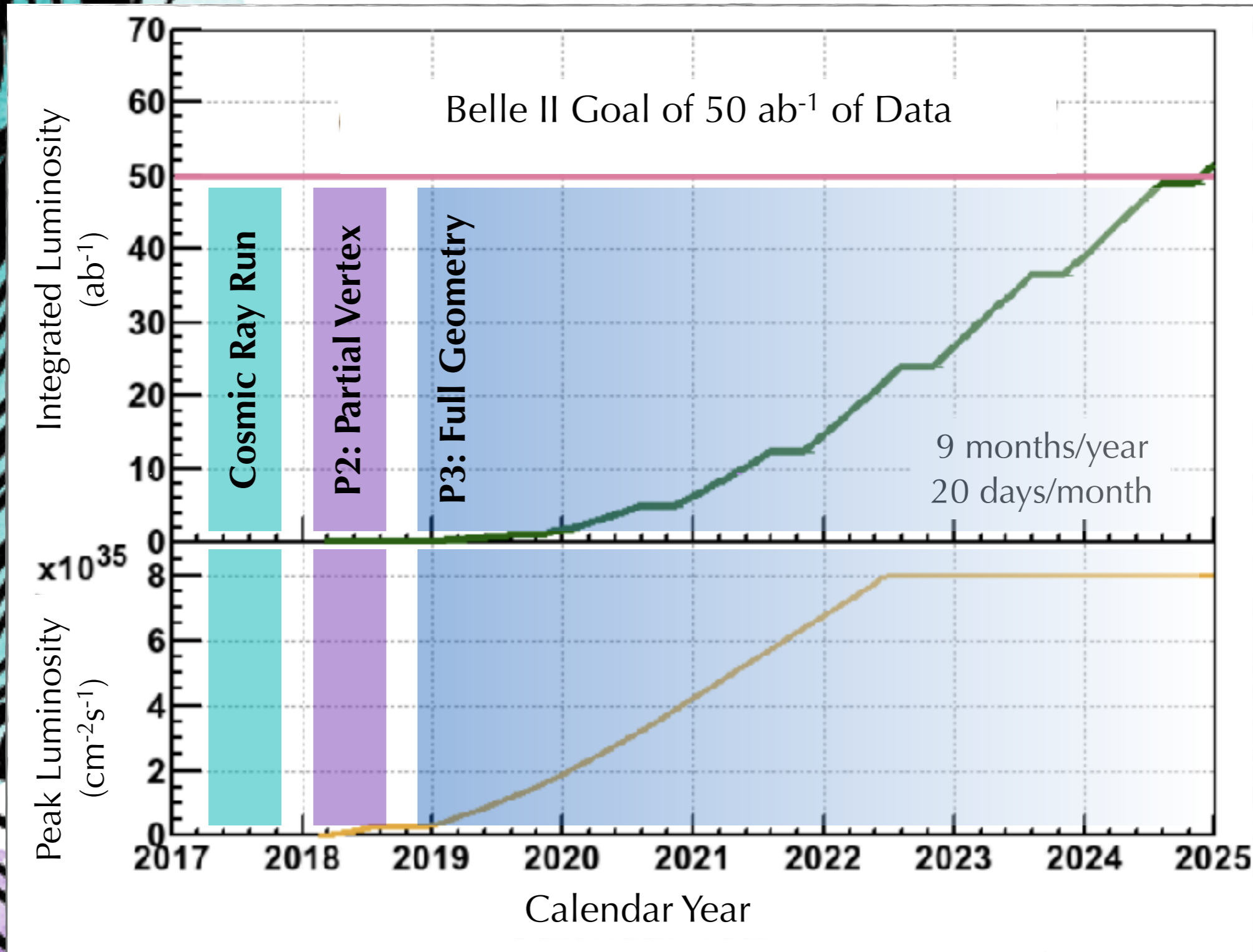
**$e^-$  beam  
(HER = 7 GeV)**

**2600 mA**

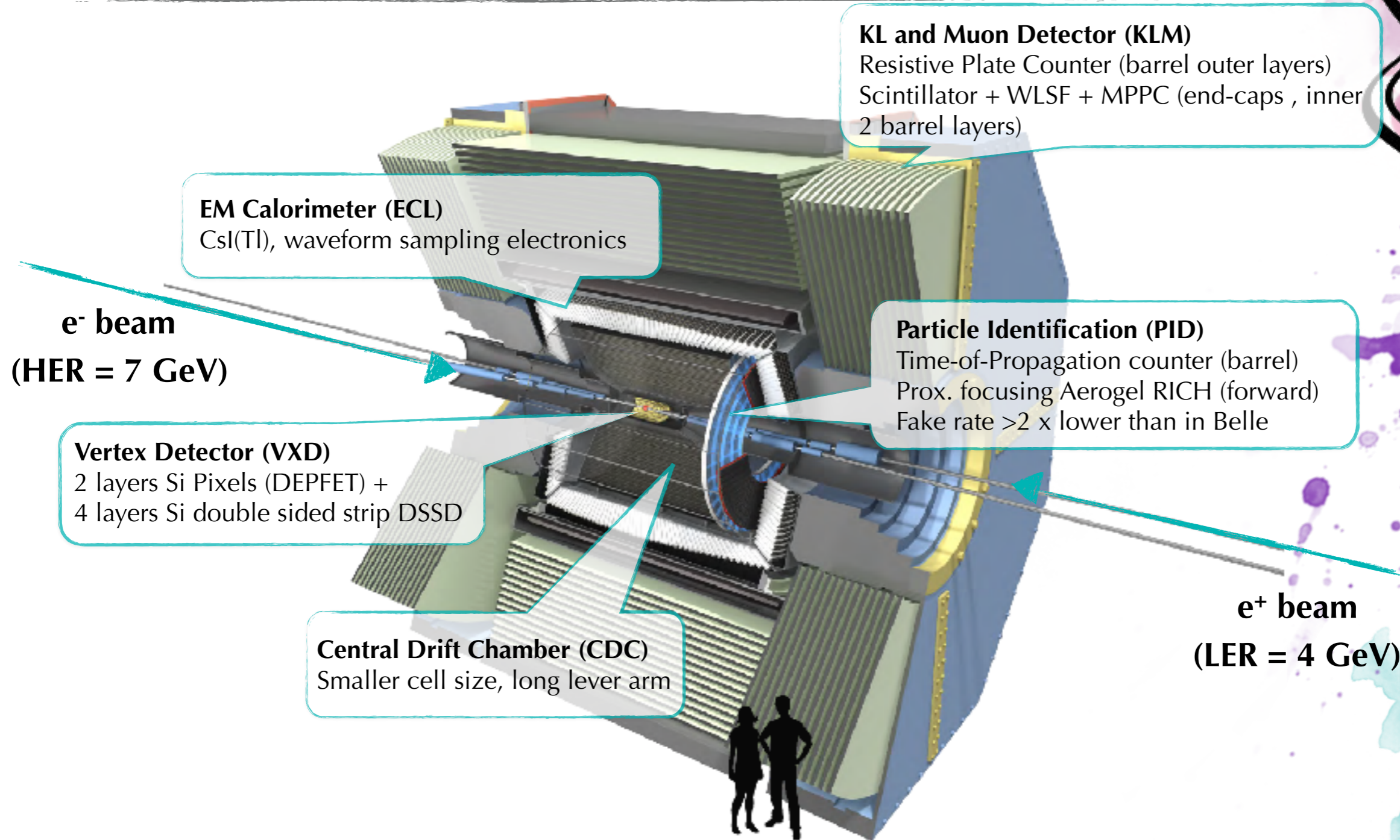
**$e^+$  beam  
(LER = 4 GeV)**

**3600 mA**

# DATA COLLECTION SCHEDULE



# THE BELLE II DETECTOR



**EM Calorimeter (ECL)**  
CsI(Tl), waveform sampling electronics

**KL and Muon Detector (KLM)**  
Resistive Plate Counter (barrel outer layers)  
Scintillator + WLSF + MPPC (end-caps, inner 2 barrel layers)

$e^-$  beam  
(HER = 7 GeV)

**Vertex Detector (VXD)**  
2 layers Si Pixels (DEPFET) +  
4 layers Si double sided strip DSSD

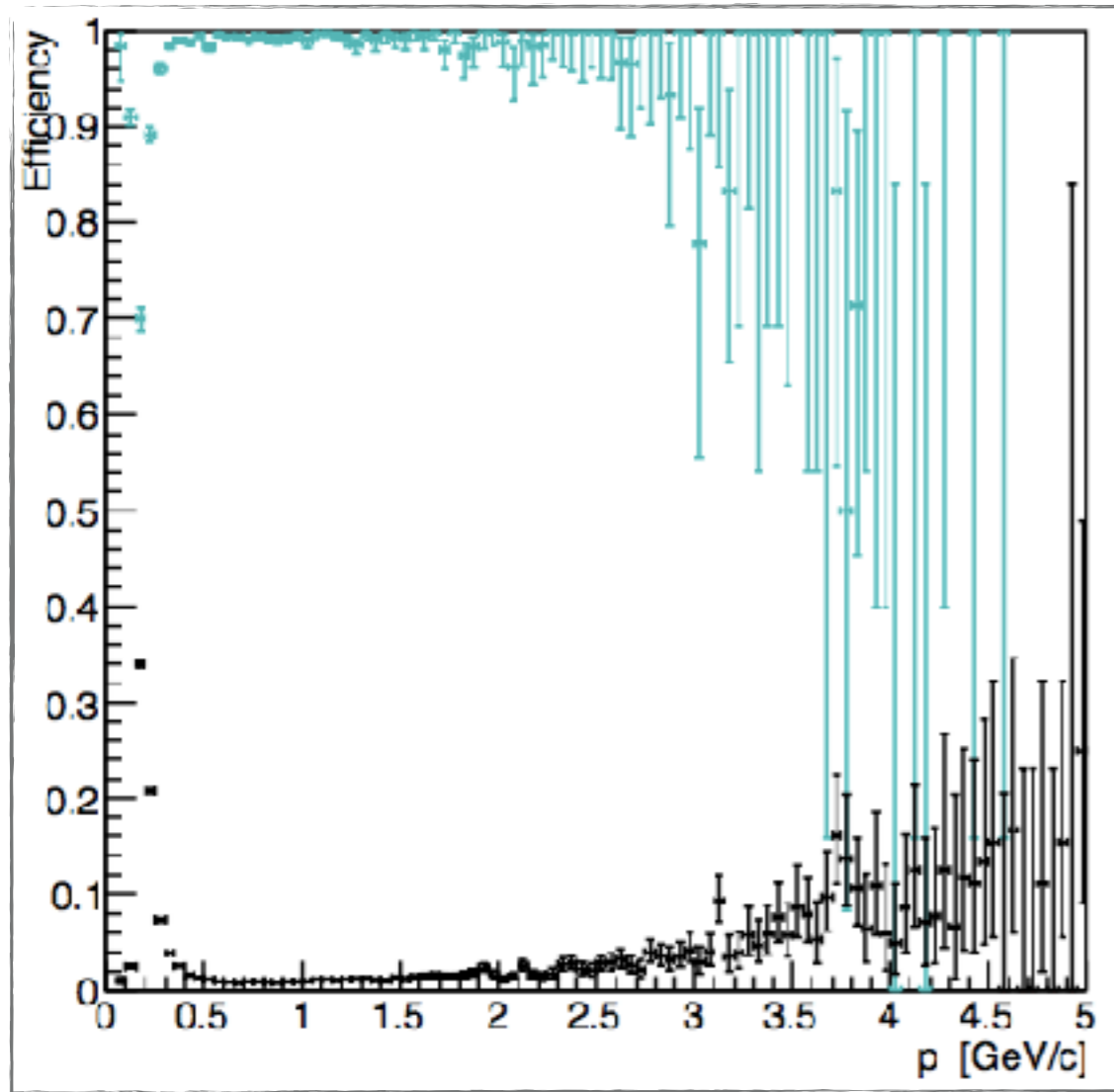
**Central Drift Chamber (CDC)**  
Smaller cell size, long lever arm

**Particle Identification (PID)**  
Time-of-Propagation counter (barrel)  
Prox. focusing Aerogel RICH (forward)  
Fake rate >2 x lower than in Belle

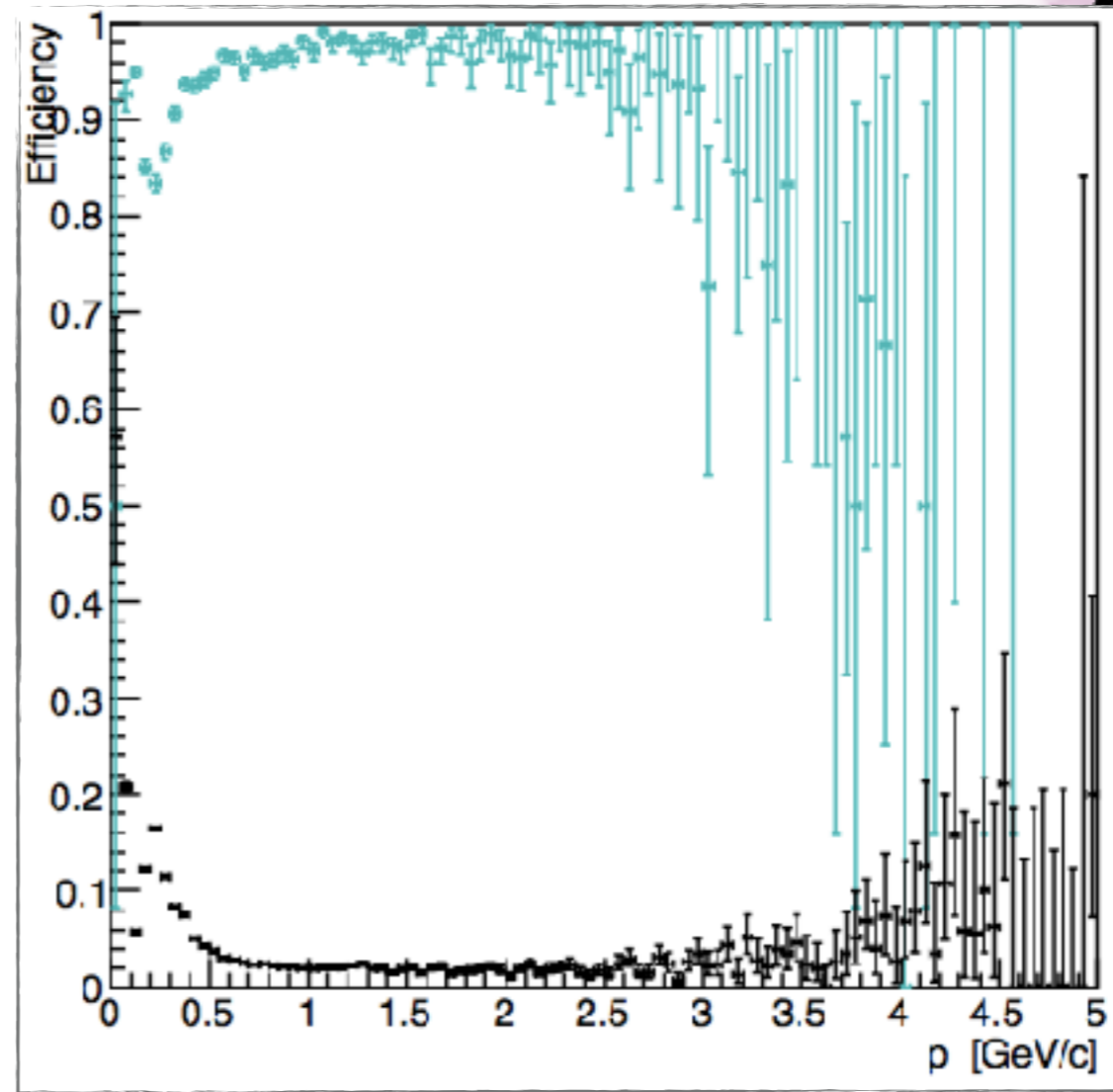
$e^+$  beam  
(LER = 4 GeV)



# DETECTOR PERFORMANCE: ELECTRON ID AND PION FAKE RATE



CDC Likelihood



dE/dx and ECL Likelihood

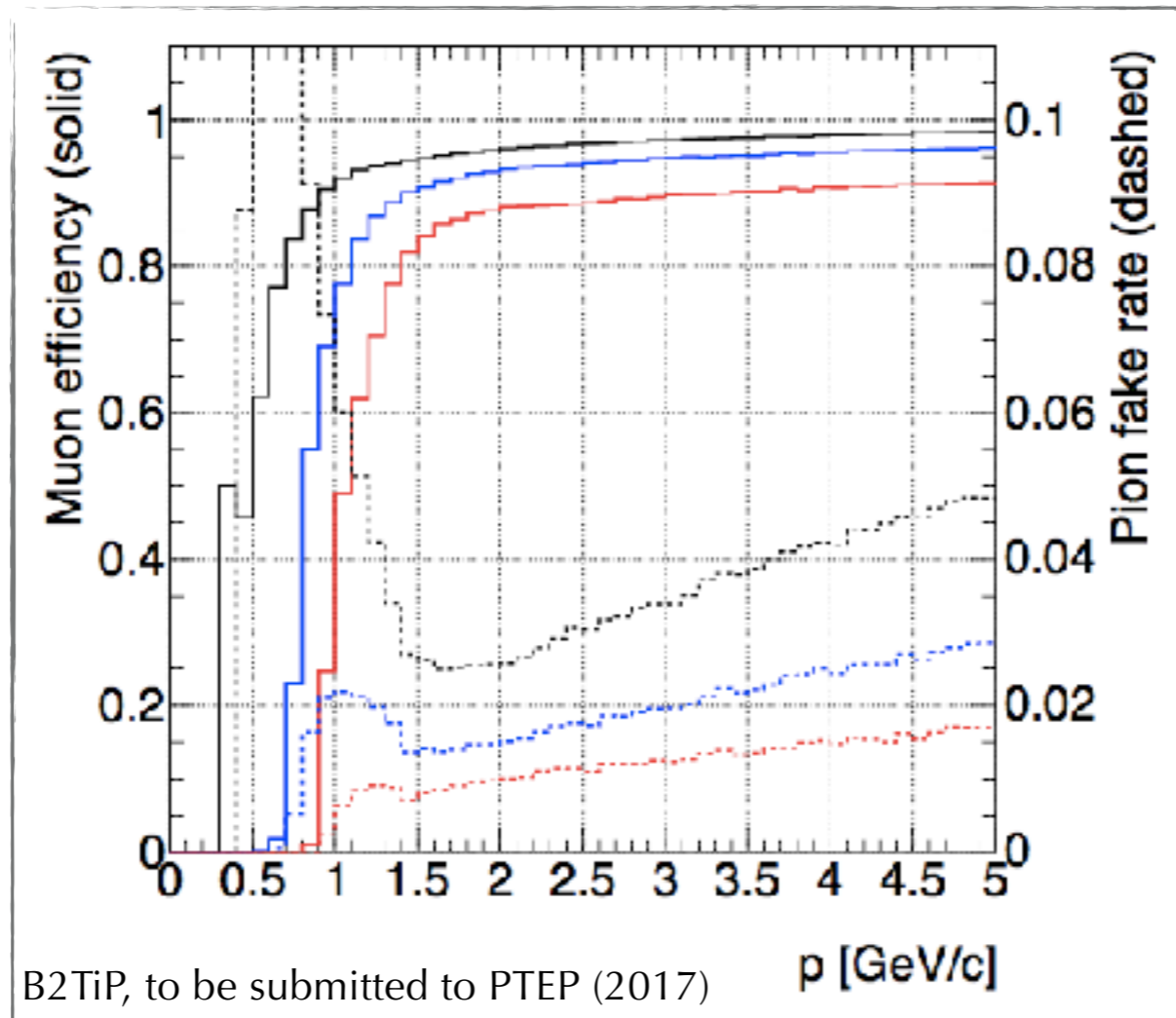
■ Electrons  
■ Pions



Global Electron PID

# DETECTOR PERFORMANCE: MUON ID AND PION FAKE RATE

$$\Delta \equiv \ln(\mathcal{L}(\mu^+; O, \ell, D, \vec{x}, \chi^2, n)) - \ln(\mathcal{L}(\pi^+; O, \ell, D, \vec{x}, \chi^2, n))$$

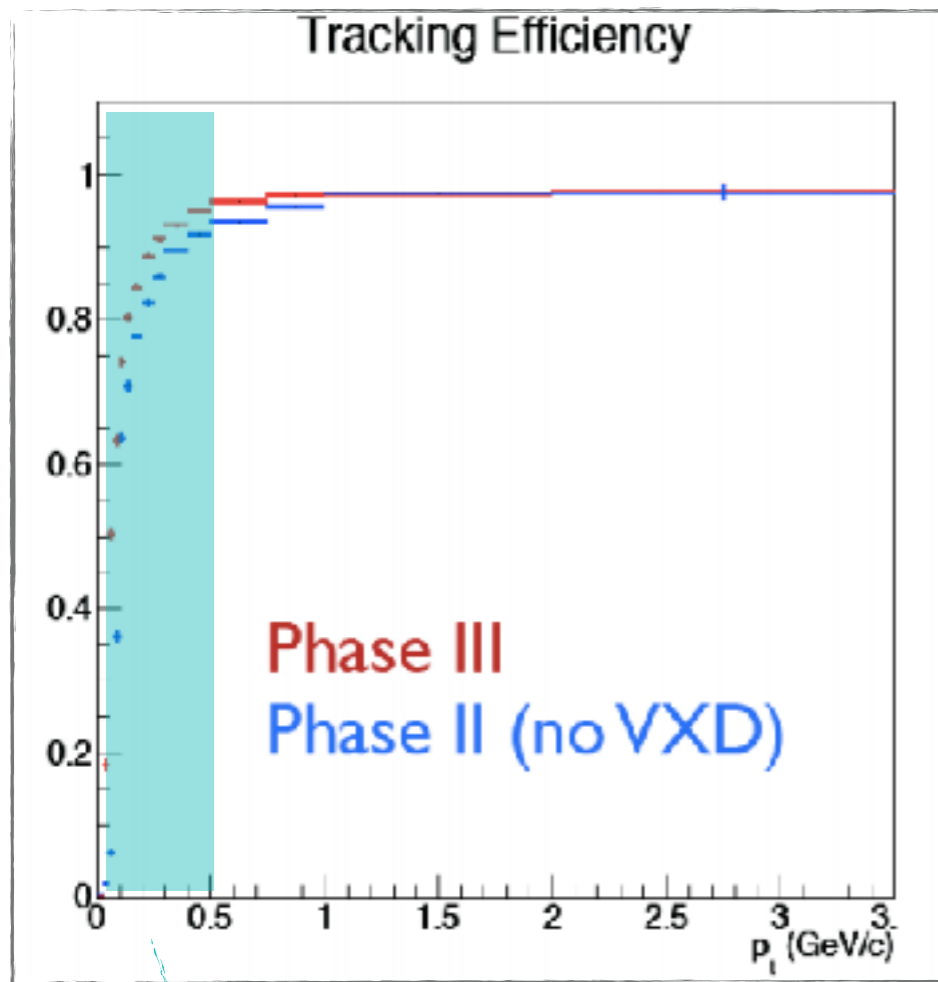


KLM Likelihood

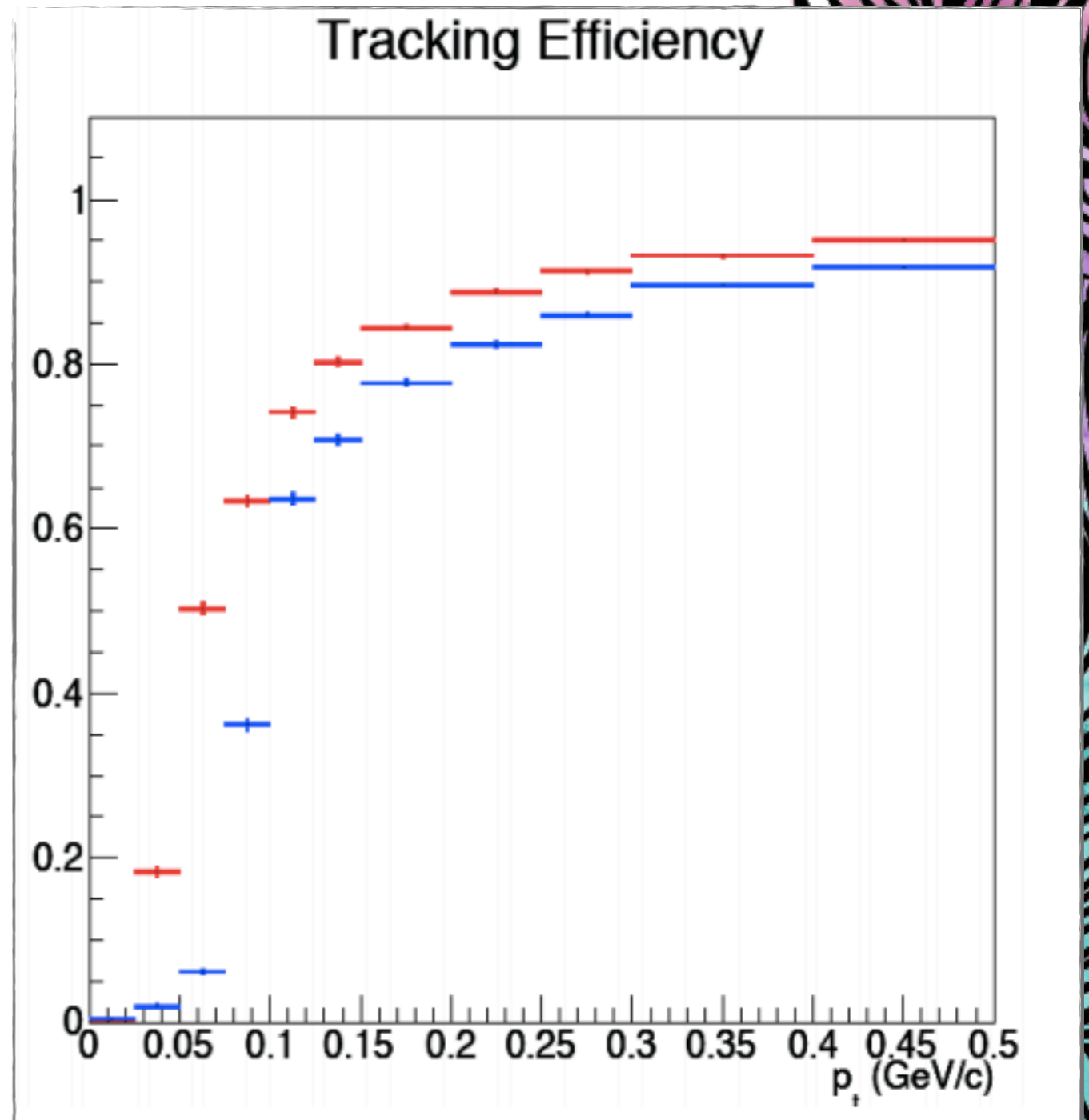


Global Muon PID

- $\Delta > 0$
- $\Delta > 10$
- $\Delta > 20$



Tracking Efficiencies  
are  $>90\%$  for  
 $p_T > 0.3 \text{ GeV/c}$



## Phase 1

### Touschek Scattering

- Intra-bunch scattering process
- Dominant with highly compressed beams

### Beam-Gas Scattering

- Bremsstrahlung (negligible) and Coulomb interactions with residual gas atoms and molecules

### Synchrotron Scattering

- Emission of photons by charged particles when deflected by B-field

## Phase 2

### Radiative Bhabha

- Photon emission prior to or after Bhabha scattering
- Interaction with iron in the magnets

### Two-Photon

- Production of very low momentum electron-positron pairs
- Increased hit occupancy in inner detectors

	Cross Section (acceptance) [nb]
BB	1.1
c	1.3
uds	2.4
$\tau\tau$	0.9
$\mu\mu^*$	0.9
$\gamma\gamma^*$	3.1
$ee^*$	74
eeee/ee $\mu\mu$	60
Total	143.7

- The Belle II trigger system is composed of a **hardware (L1) trigger** and a **software high level trigger (HLT)**
- $e^+e^-$  beams accelerated at 500 MHz, while the output rate must be kept below 10 kHz

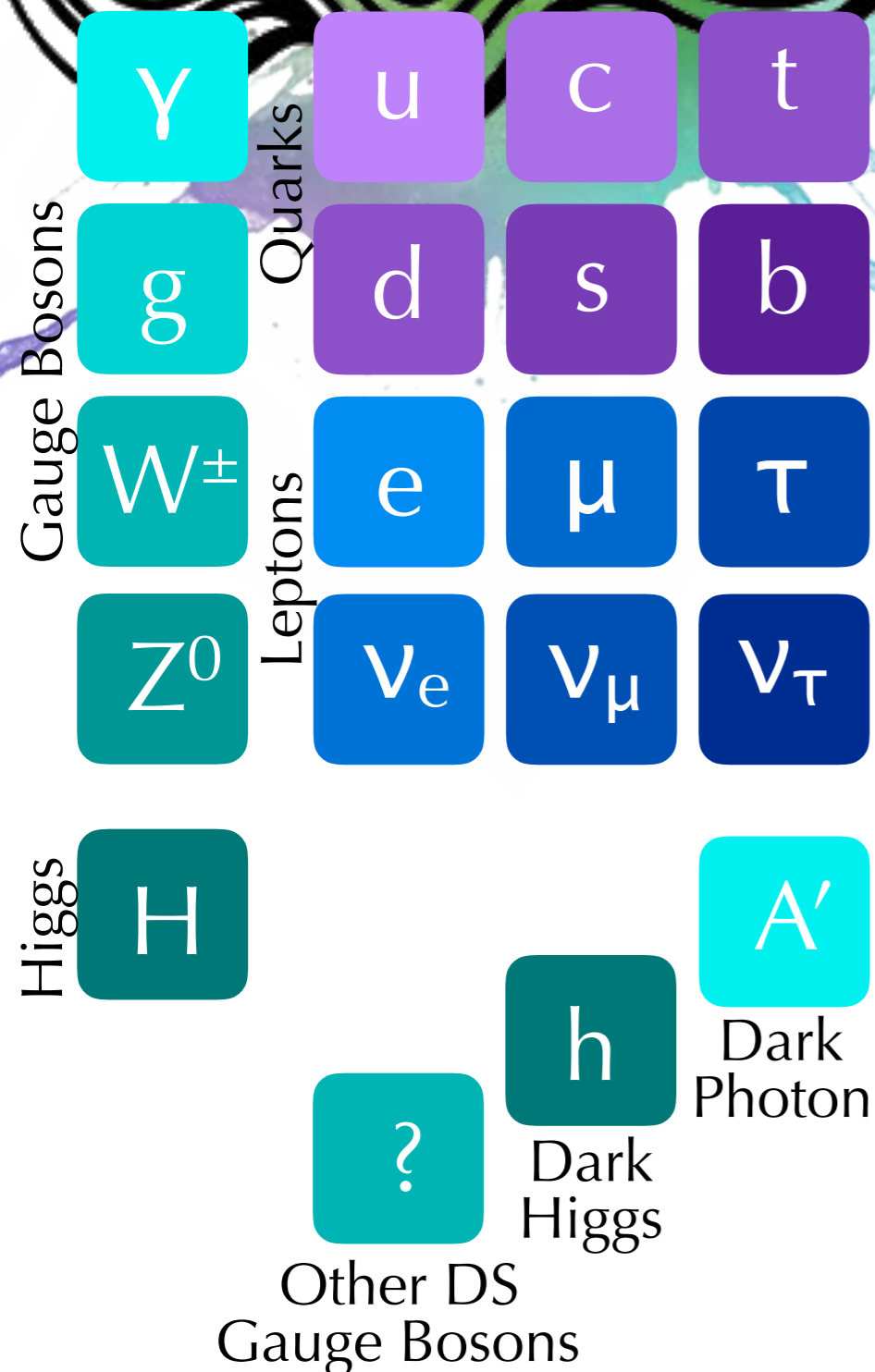
total cross section:  
~140nb

Subdetector Acceptances

CDC:  $17^\circ < \theta < 150^\circ$   
ECL:  $12.4^\circ < \theta < 155.1^\circ$

\*two particles  
in ECL  
coverage

- L1 trigger
  - 3D tracking is implemented on CDC trigger
  - 3D **ECL bhabha veto logic** is available on the ECL trigger to identify bhabha events with high purity
  - Matches between CDC tracks and ECL clusters and other low level reconstruction techniques are implemented on the L1 trigger
- HLT trigger
  - Reconstructs the events online by using the offline reconstruction algorithm
  - 6000 cpu cores at target luminosity
- As we will see, the Belle II trigger system must/will implement **dedicated triggers for DS studies**



## SHORTCOMINGS OF THE SM

- Gauge bosons mediate interactions between visible matter.
- But what mediates interactions between dark matter and itself and dark matter and visible matter?
- If DM is not a WIMP, we must search for a **portal** between the dark sector and visible sector.

- A dark sector gauge boson that couples to the conserved current associated with the dark sector group,  $U(1)_X$ .
- $U(1)_X$  undergoes **kinetic mixing** with the SM  $U(1)_Z$  (falling out of EW breaking) gauge group (our portal to the dark sector).
- The inclusion of this dark sector gauge group provides additional terms to the SM Lagrangian.

$$\mathcal{L} = \mathcal{L}_{SM} - \frac{1}{4} F'_{\mu\nu} F'^{\mu\nu} + \frac{\epsilon}{2} F'_{\mu\nu} F^{\mu\nu}$$

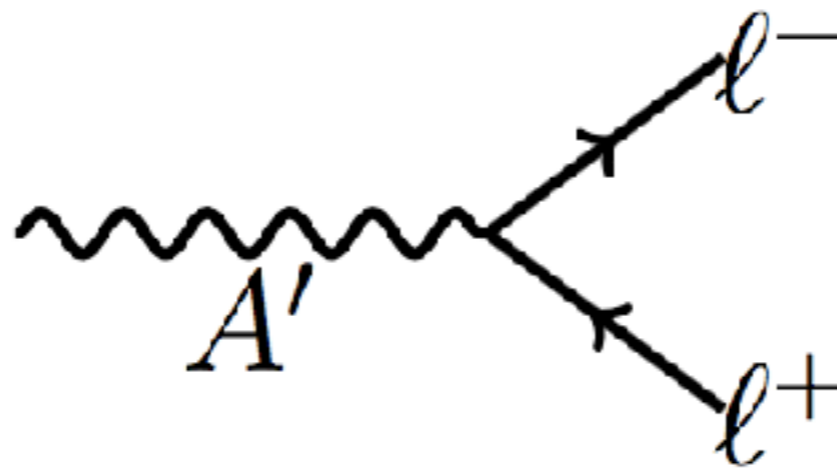
B. HOLDOM, PHYS REV LETT B:166:196 (1986)



# FERMIONIC INTERACTIONS

- $U(1)_X \otimes U(1)_Z$  undergoes SSB similar to that of  $SU(2)_L \otimes U(1)_Y$  in the SM.
- This gives rise to mass eigenstates  $A$  and  $A'$ , where  $m_A = 0$  and  $m_{A'} > 0$ .
- It also follows that the dark photon **interacts with BOTH** the dark sector and **the visible sector**.

B. BRAHMACHARI, A. RAYCHAUDHURI, NUCL PHYS B:887 (2014)



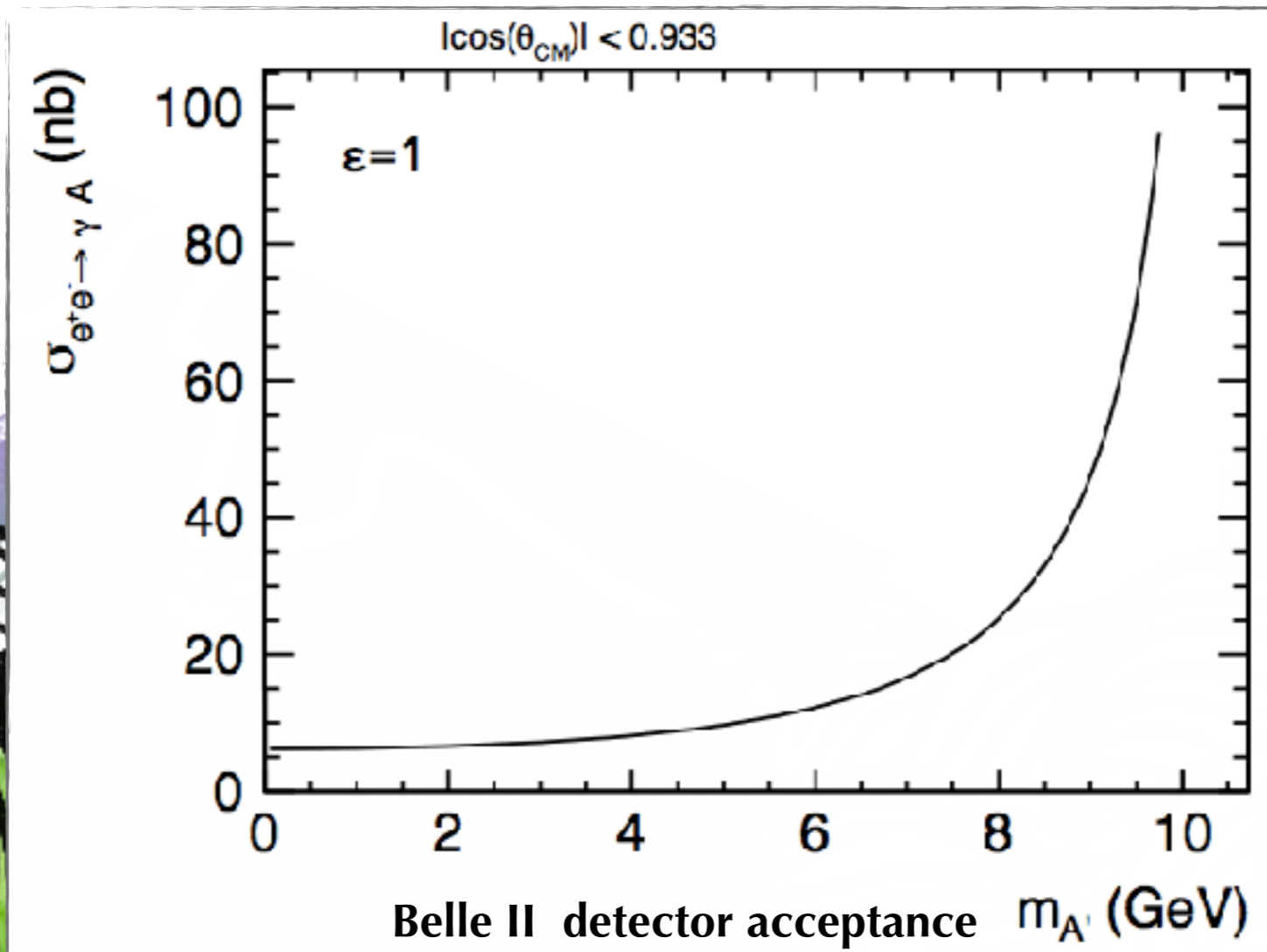
- The mixing strength between the photon and the dark photon is constrained through the cross section as...

$$\sigma = \frac{2\pi\epsilon^2\alpha^2}{E_{cm}^2} \left(1 - \frac{m_{A'}^2}{E_{cm}^2}\right) \left( \left(1 + \frac{2m_{A'}^2/E_{cm}^2}{(1 - m_{A'}^2/E_{cm}^2)^2}\right) \Theta - \cos\theta_{max} + \cos\theta_{min} \right)$$

↖ **Mixing Strength Parameter**

$$\Theta = \log \left( \frac{(1 + \cos\theta_{max})(1 - \cos\theta_{min})}{(1 - \cos\theta_{max})(1 + \cos\theta_{min})} \right)$$

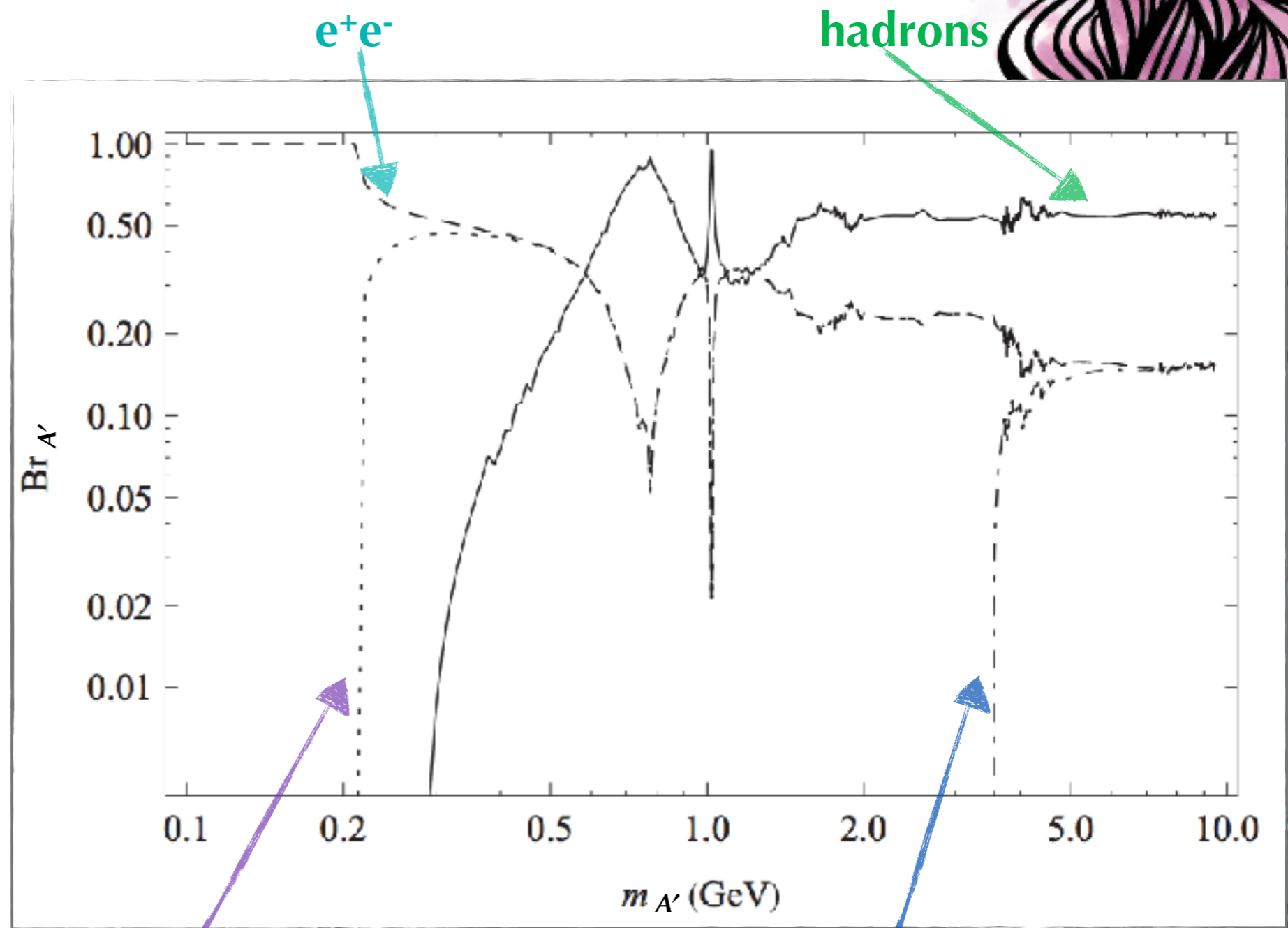
R. ESSIG, P. SCHUSTER, N. TORO, PHYS REV LETT D:80:015003 (2009)



# DARK PHOTON BRANCHING RATIOS

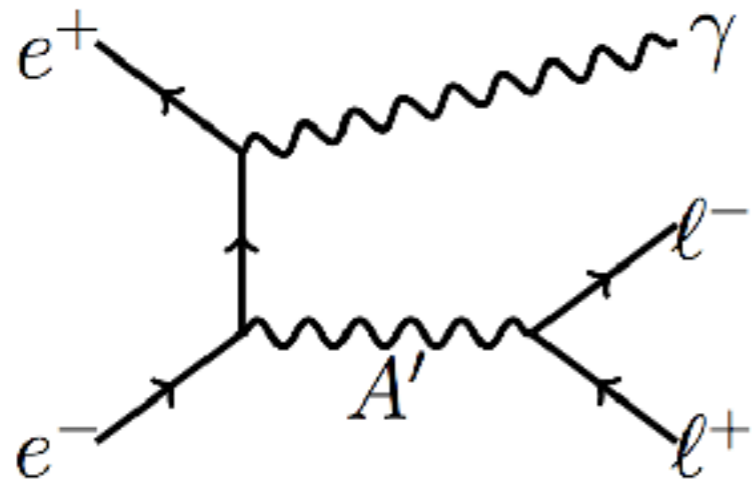
Strictly for Decays  
to Visible Matter

- Dielectron mode has the largest branching fraction.
- Dimuon mode will have higher sensitivity above threshold as there exist far fewer backgrounds.

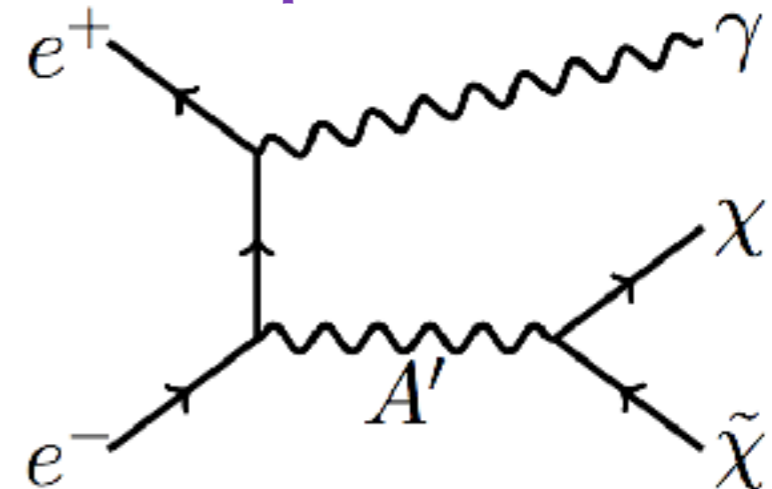


$\mu^+\mu^-$  B. BATELL, B. POSPELOV, A. RITZ, "PROBING  $\tau^+\tau^-$   
A SECLUDED U(1) AT B-FACTORIES (2009)"

## Dilepton Mode

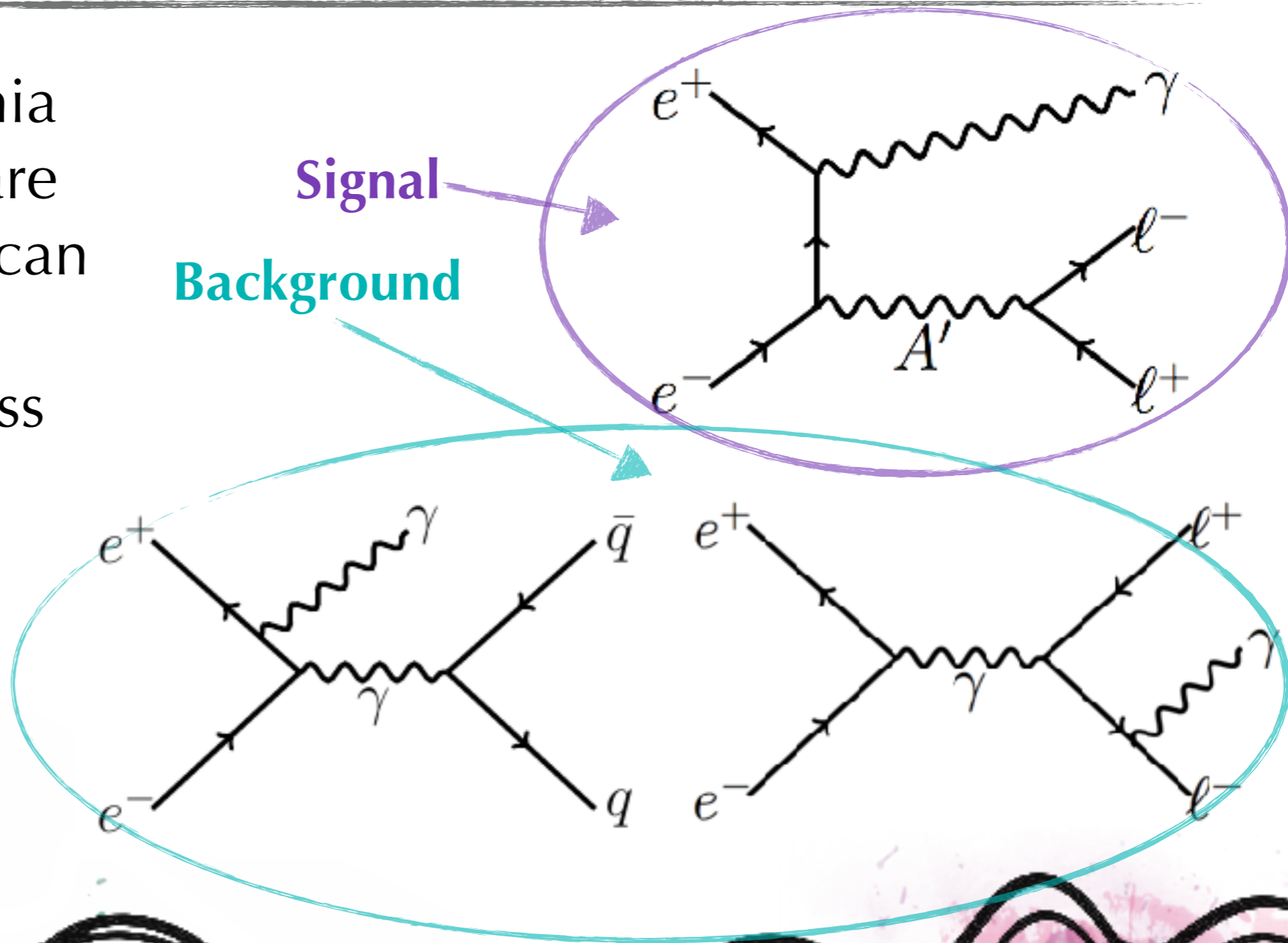


## Monophoton Mode



- If dark photon is the lightest DM particle, we expect to mostly observe decays to SM leptons (e or  $\mu$ ).
- Dark Photon **invariant mass “bump search”**.
- More difficult to detect (larger irreducible backgrounds).
- Search relies heavily on **special triggers** being employed at Belle II.
- Otherwise, the dark photon will likely decay to light DM, signal of a monoenergetic photon.

- **ISR** results in quarkonia resonances — these are well understood and can be included in our dilepton invariant mass fits
- **FSR** is also a large background to be concerned with



# HOW DOES BACKGROUND COMPARE TO SIGNAL?

	Belle 1ab <sup>-1</sup>	Belle 2 Phase II 20fb <sup>-1</sup>	Belle 2 Phase III 50ab <sup>-1</sup>	
<b>Signal</b>	$\sigma_{A'\gamma} \approx 1\text{fb}$	$1 \times 10^3$	$2 \times 10$	$5 \times 10^4$
	$\sigma_{A'\gamma} \approx 10\text{fb}$	$1 \times 10^4$	$2 \times 10^2$	$5 \times 10^5$
<b>Bkg</b>	$\sigma_{e^+e^-\gamma} = 300\text{nb}$	$3 \times 10^{11}$	$6 \times 10^9$	$1.15 \times 10^{13}$
	$\sigma_{\mu^+\mu^-\gamma} = 1.148\text{nb}$	$1.148 \times 10^9$	$2.296 \times 10^7$	$5.74 \times 10^{10}$

$$\sigma = \frac{N_{\text{sig}}}{\epsilon L}$$

Assuming  
Perfect  
Efficiency

*\* Dielectron background cross section includes requirement that final state particles be in the detector acceptance*

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Example cross-sections  
arbitrarily chosen

*\* Dielectron background cross section includes requirement that final state particles be in the detector acceptance*

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Assuming Perfect Efficiency

The signal resolution in Belle II is better than Belle and thus, (for similar luminosities) we expect better sensitivity

*\* Dielectron background cross section includes requirement that final state particles be in the detector acceptance*



# HOW DOES BACKGROUND COMPARE TO SIGNAL?

	Belle 1ab <sup>-1</sup>	Belle 2 Phase II 20fb <sup>-1</sup>	Belle 2 Phase III 50ab <sup>-1</sup>
<b>Signal</b>			
$\sigma_{A'\gamma} \approx 1\text{fb}$	$1 \times 10^3$	$2 \times 10$	$5 \times 10^4$
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$$\sigma = \frac{N_{\text{sig}}}{\epsilon L}$$

Assuming Perfect Efficiency

Large amount of background highlights the necessity for an improved triggering system

\* Dielectron background cross section includes requirement that final state particles be in the detector acceptance

- Our main background  $ee \rightarrow \gamma\gamma(\gamma)$ ,  $ee(\gamma)$  which is proportional to the instantaneous luminosity
- Optimize the trigger efficiency with **dedicated trigger schemes** for different luminosity periods
- In Belle, our **Bhabha veto trigger** was highly efficient but also threw out many low-multiplicity events with the same signature as our DS searches.

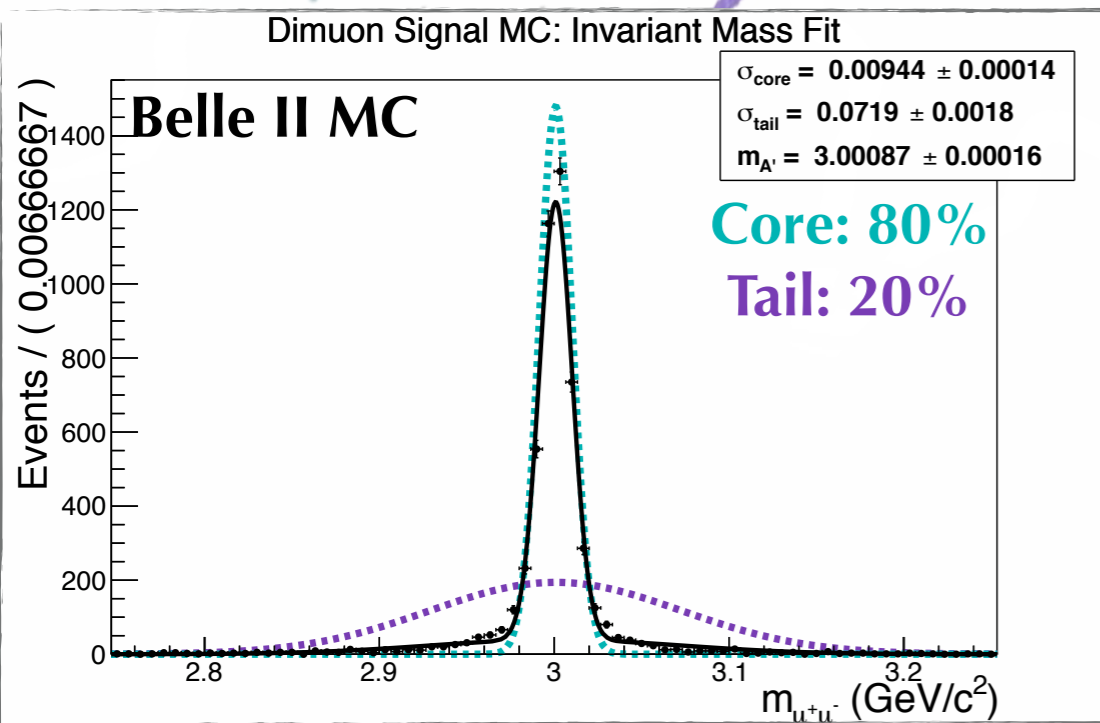
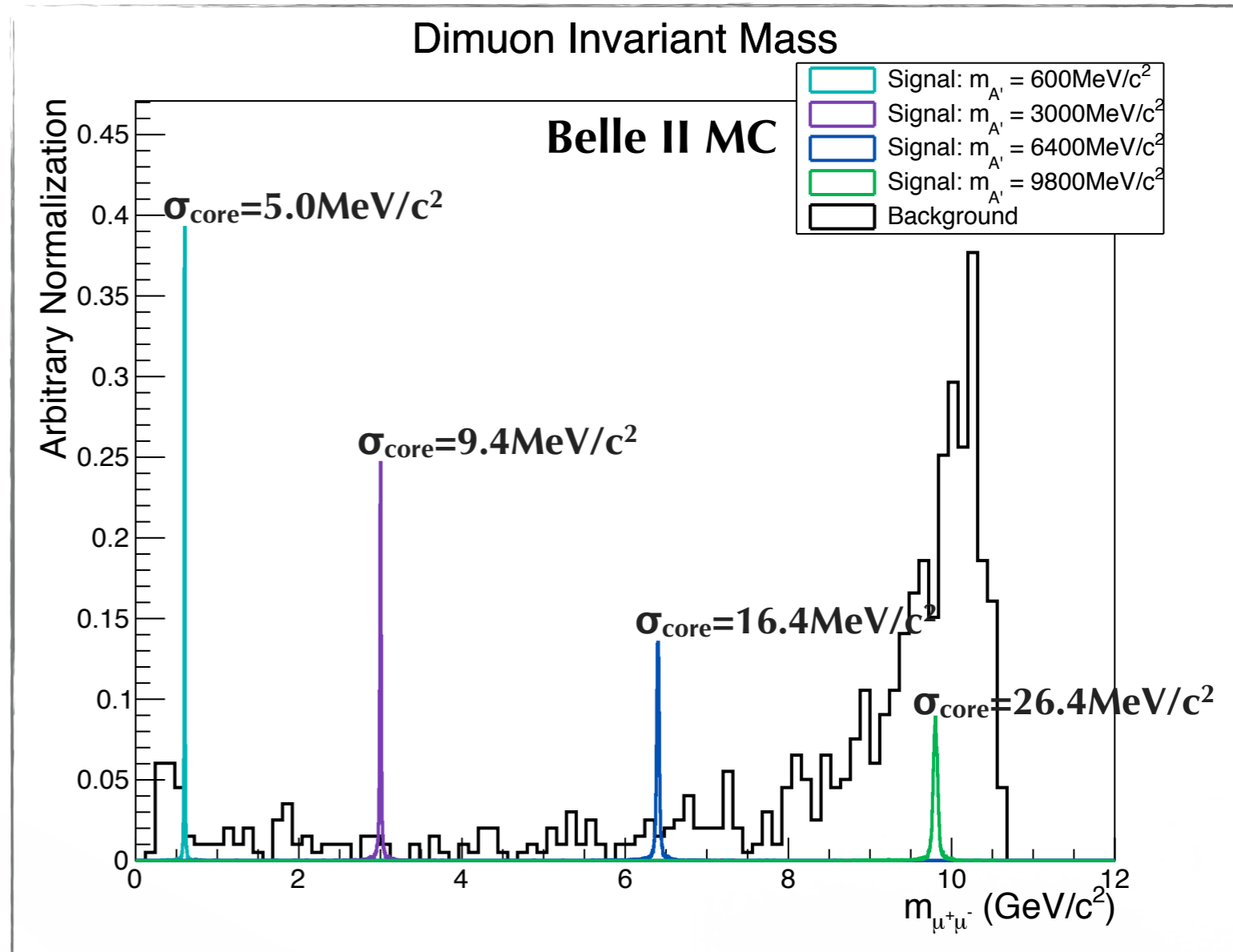
1. Two oppositely charged, “well-identified” ( $\epsilon_{id}$  or  $\mu_{id} > 0.1$ ) lepton tracks
2. Total invariant mass (two tracks AND photon) near the  $\Upsilon(4S)$  resonance ( $9.5 \leq m_{(\mu+\mu^- \text{ or } e+e^-)\gamma} \leq 10.8 \text{ GeV}/c^2$ )
3. One photon with  $E_\gamma \geq 0.2 \text{ GeV}$  (reduced background from bremsstrahlung)

<b>Dimuon mode Belle II MC</b>			
$A'$ Mass Hypothesis*	Cut 1 (%)	Cut 2 (%)	Cut 3 (%)
0.6 GeV	67.81	56.64	56.63
3.0 GeV	61.41	52.86	52.86
6.4 GeV	74.60	65.57	65.57
9.8 GeV	80.56	75.07	73.00
Background	20.57	14.23	7.96

\* 52 Mass Hypotheses from 0.2 to 10.4 GeV were examined. 4 have been arbitrarily chosen to show here.



- Mass resolutions are determined by a double Gaussian fit
- Increased  $A'$  mass degrades the signal resolution
- Overall, Belle II signal resolutions are **improved over Belle by ~30%**



# GLOBAL EXPERIMENTAL EFFORTS: $A'$ $\longrightarrow$ VISIBLE

Experiment	Decay Mode
BaBar	$e^+e^- \rightarrow \gamma A'$
BESIII	$J/\psi \rightarrow \gamma A'$
KLOE 2013	$\phi \rightarrow \eta A'$
KLOE 2014	$e^+e^- \rightarrow \gamma A'$
E141	$e_{\text{beam}}^- (\text{nucleus}) \Rightarrow e^- A'$
E774	$e_{\text{beam}}^- (\text{nucleus}) \Rightarrow e^- A'$
A1	$e_{\text{beam}}^- (\text{nucleus}) \Rightarrow e^+e^- \rightarrow A'$
APEX	$e_{\text{beam}}^- (\text{nucleus}) \Rightarrow e^+e^- \rightarrow A'$
HADES	$p_{\text{beam}}^+ (\text{nucleus}) \Rightarrow K^+K^- \Rightarrow \pi^0 \rightarrow \gamma A'$
NA48/2	$p_{\text{beam}}^+ (\text{nucleus}) \Rightarrow K^+K^- \Rightarrow \pi^0 \rightarrow \gamma A'$
WASA	$p_{\text{beam}}^+ (\text{nucleus}) \Rightarrow K^+K^- \Rightarrow \pi^0 \rightarrow \gamma A'$

**$e^+e^-$  Collider Experiments**  
 Cover the largest parameter space of experiments searching for Dark Photons  
 Consequence of high  $e^+e^- \rightarrow \gamma\gamma$  cross-sections

**Electron Beam-Dump Experiments**  
 Dark Photons are emitted via Bremsstrahlung  
 Secondary to original processes being explored

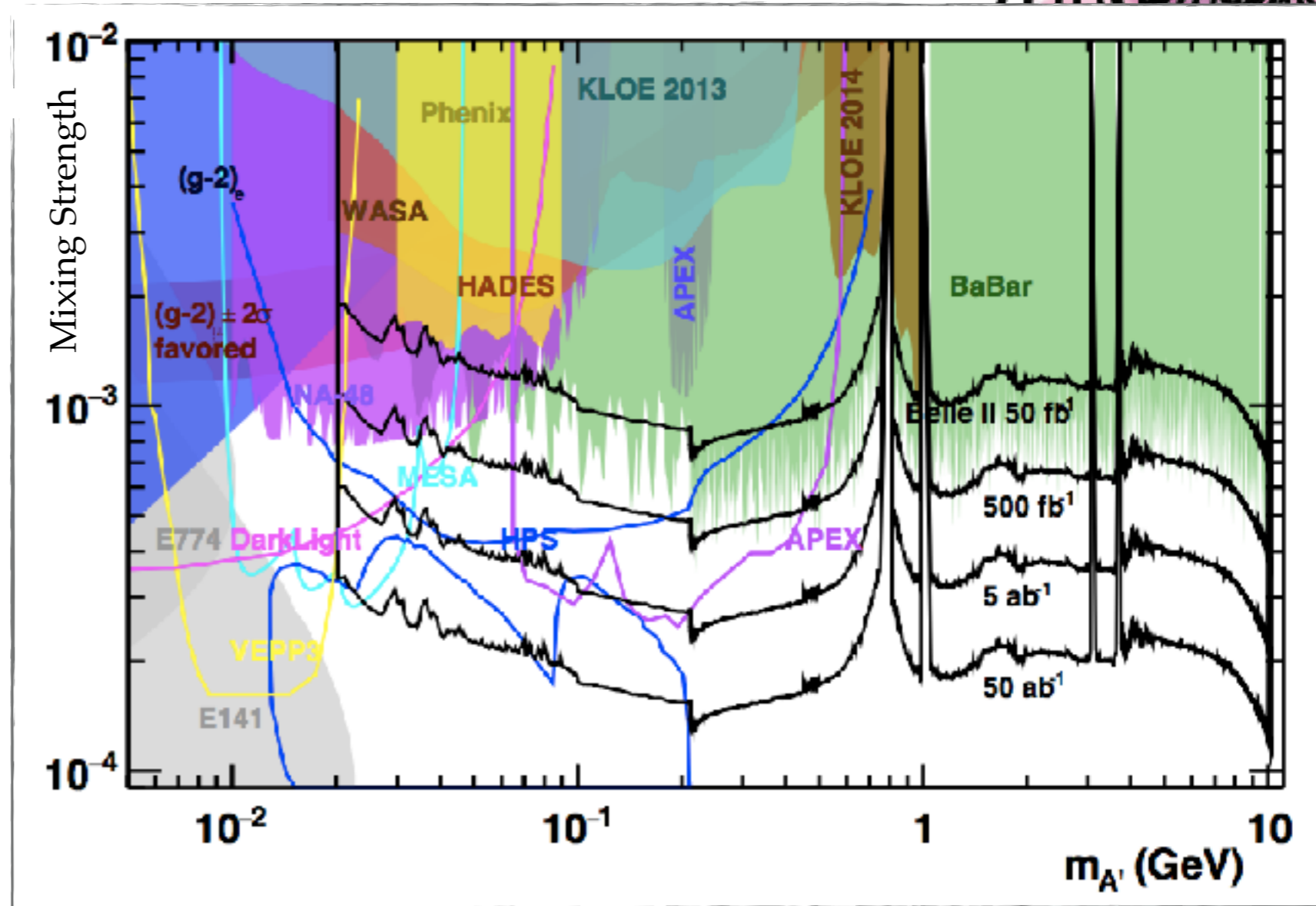
**Proton Fixed Target Experiments**  
 Proton beams scatter off of niobium or liquid hydrogen targets  
 Dark Photons are produced and decay similarly as in electron fixed target experiments

**Electron Fixed Target Experiments**  
 Electron beams scatter off heavy nuclei  
 Dark Photons are produced and decay into  $e^+e^-$  pairs

# PROPOSED REACH OF BELLE II: $A' \rightarrow \text{VISIBLE}$

90% CL upper limits on the mixing strength — projections made from the results of BaBar's 2014 search\*

Total BaBar Data Set:  $514 \text{ fb}^{-1}$   
 Total Belle Data Set:  $1000 \text{ fb}^{-1}$   
 Expected Belle II Data Set:  $50 \text{ ab}^{-1}$

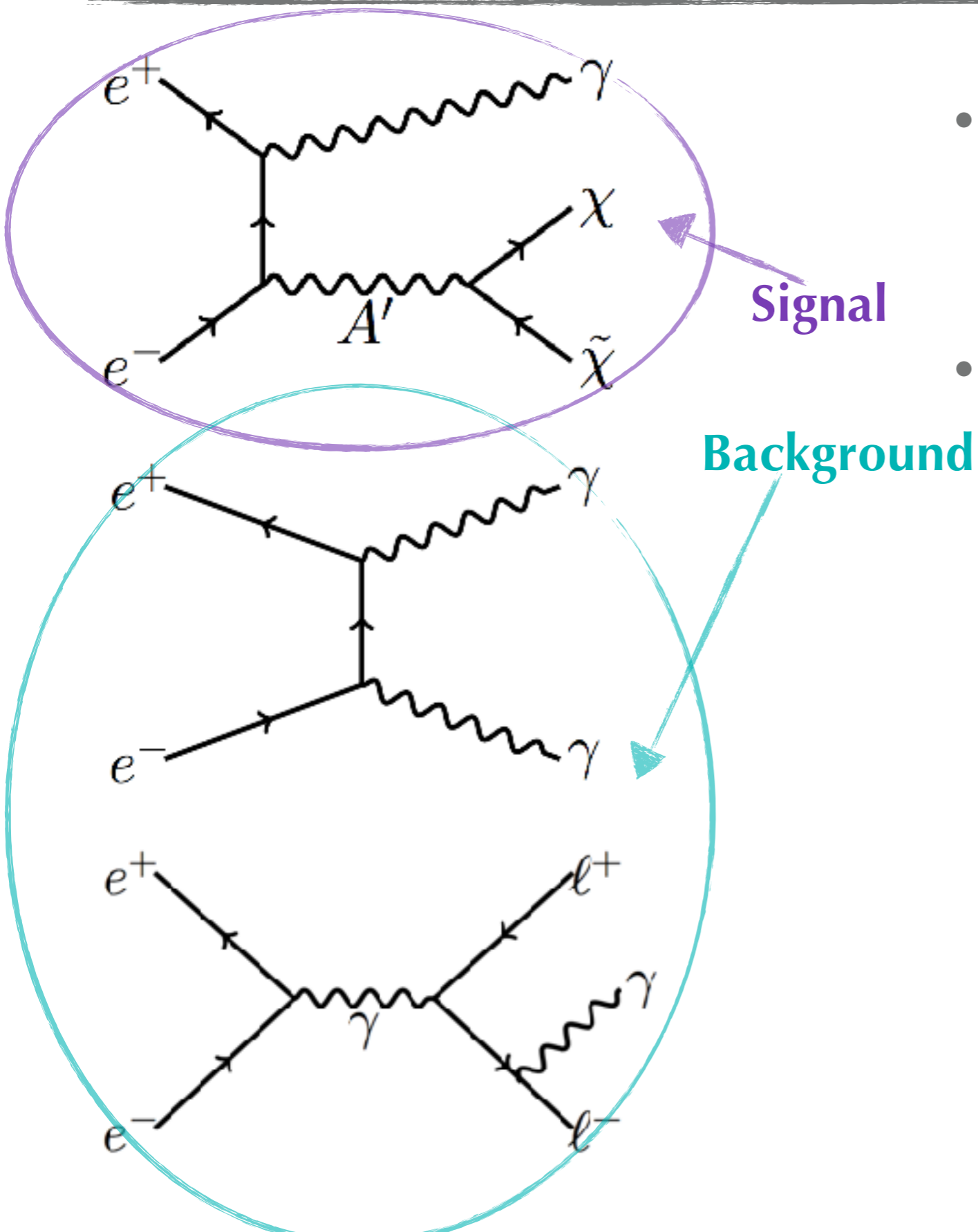


(BABAR): J.P. LEES ET AL., PHYS REV LETT 113(20):201801 (2014)

(BESIII): V. PRASAD ET AL., CHARM PHYS WORKSHOP (2015)

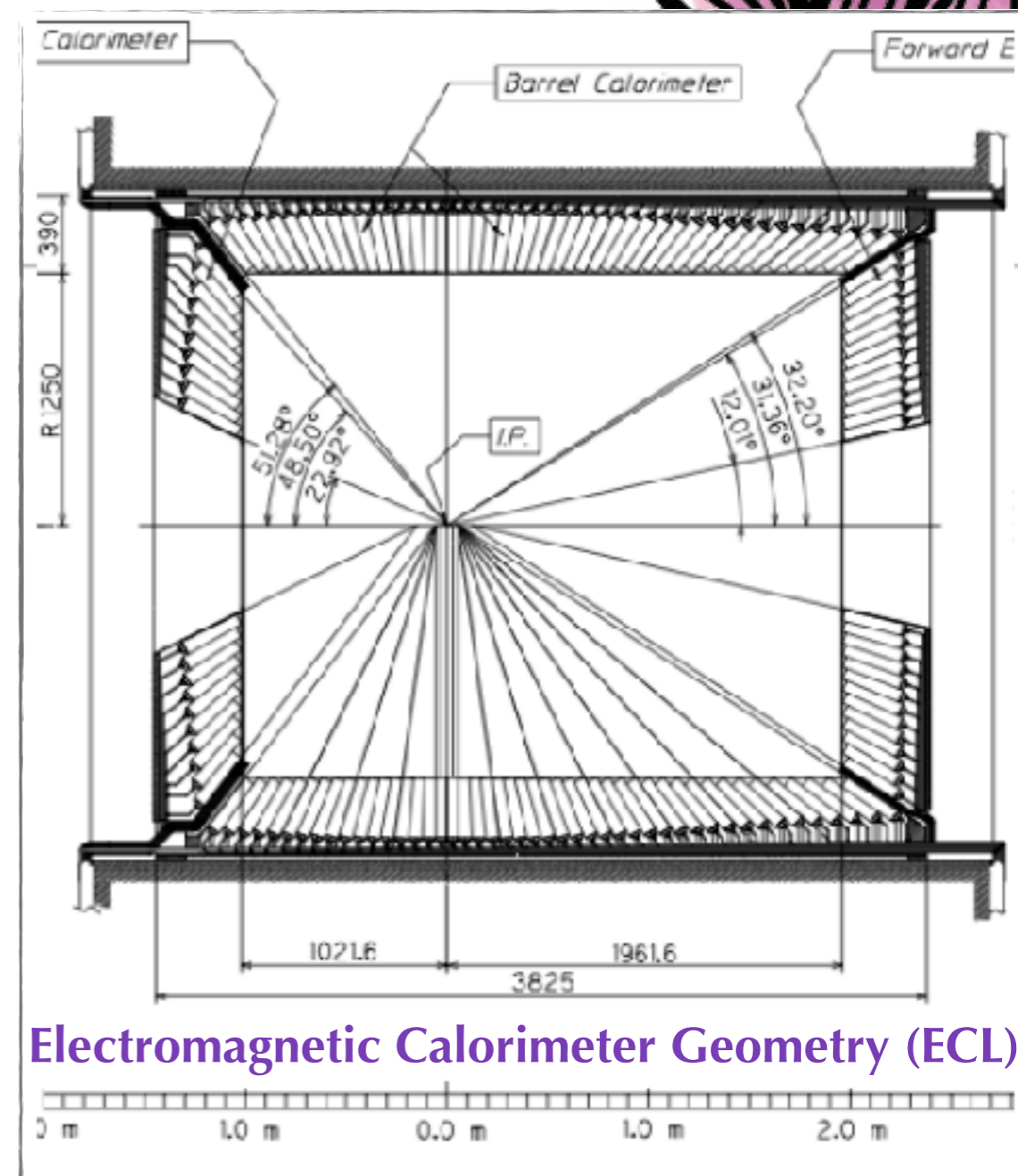
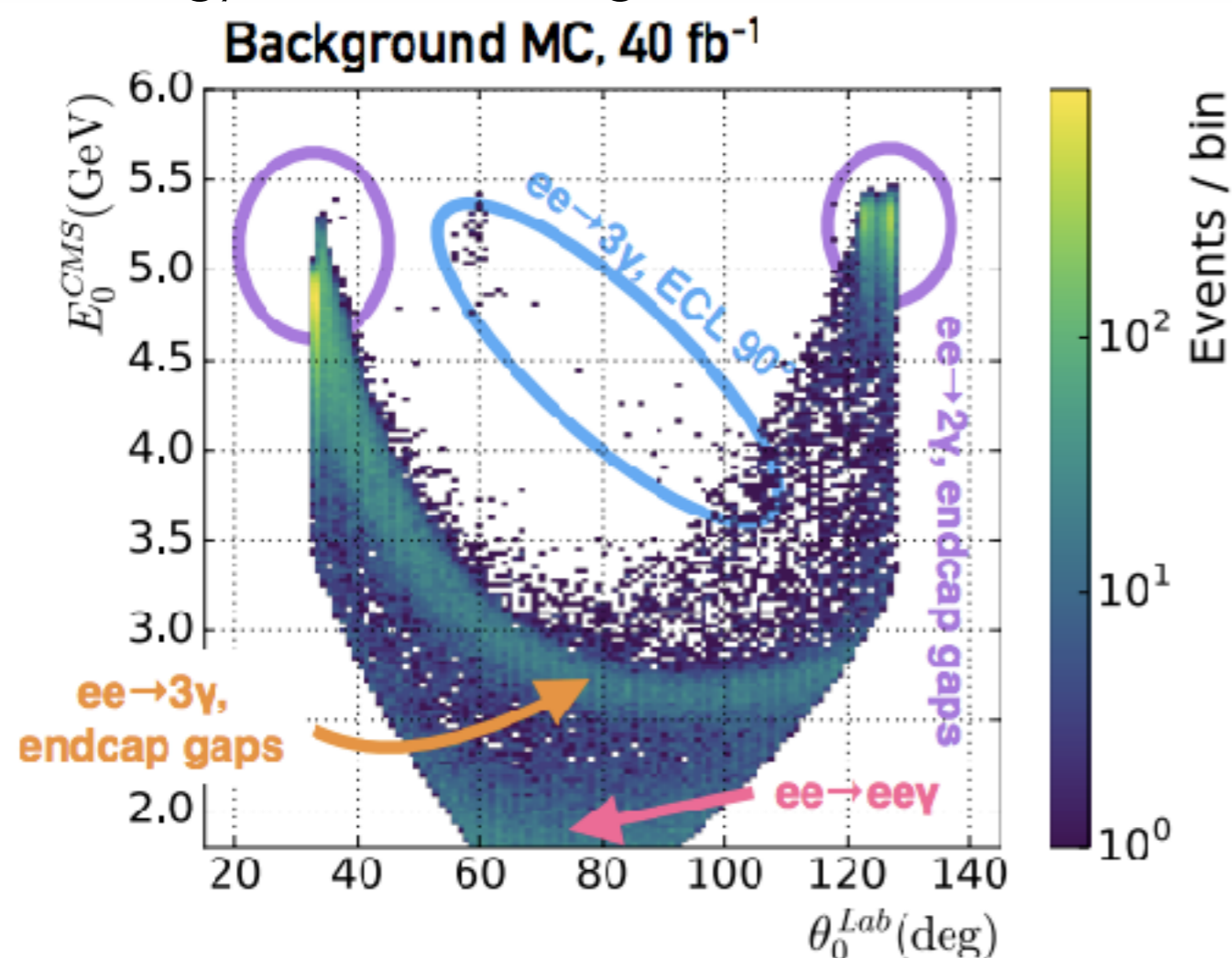
\*Better mass resolution and trigger efficiencies were assumed in making extrapolations

# BACKGROUND PROCESSES FOR THE MONOPHOTON MODE



- Search performed by fitting the **recoil mass distribution**
- Background arises from events in which all final state particles except one photon are outside the detector acceptance

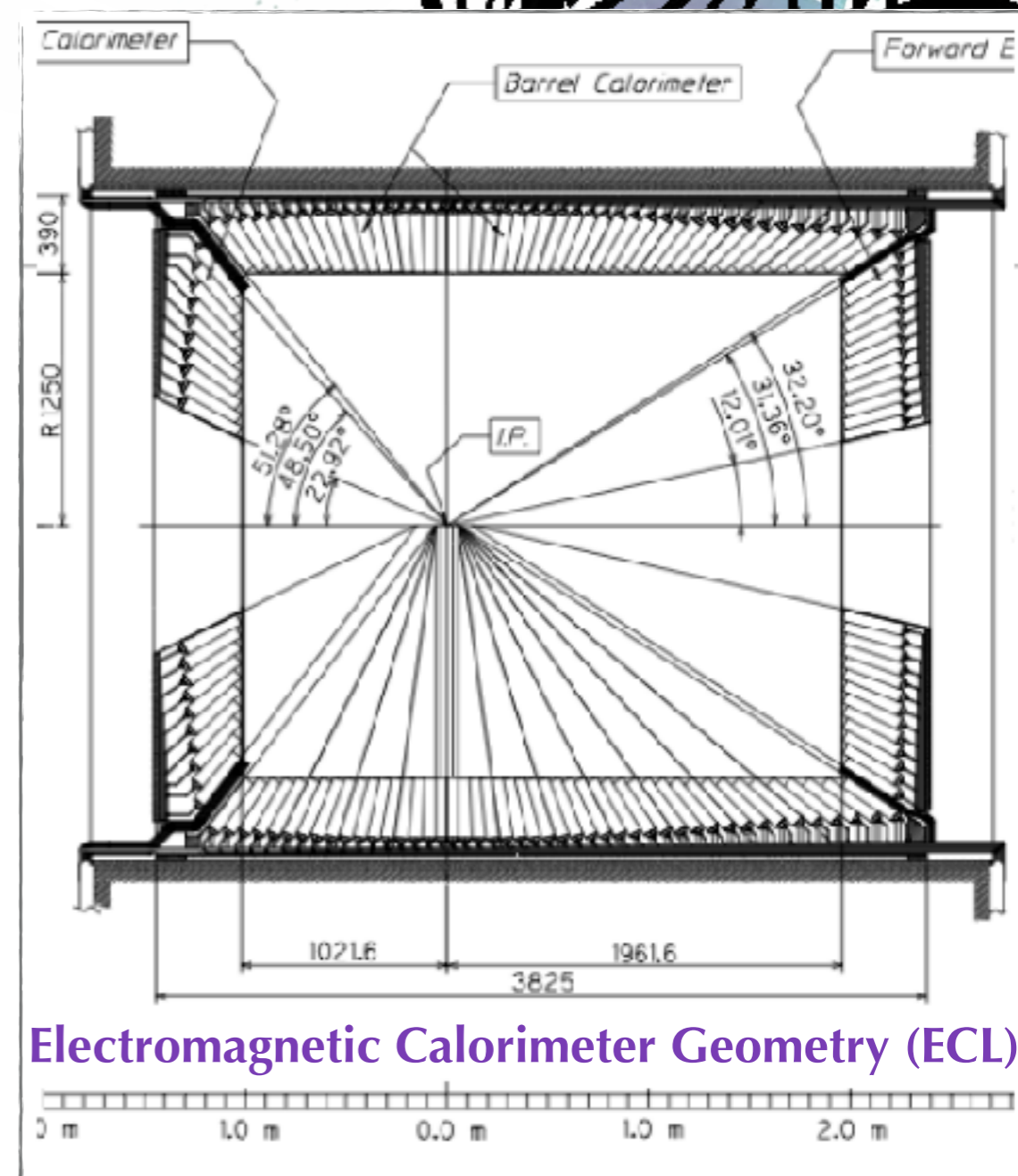
- $e^+e^- \rightarrow \gamma\gamma$  with one photon going through the ECL gaps.
- $e^+e^- \rightarrow \gamma\gamma\gamma$  with one photon going through the backward ECL gap and one photon near  $\theta=0^\circ$ .
- Irreducible backgrounds — events at low energy and wide angles.



T. FERBER, "EPS CONFERENCE: DS AT BELLE II" (JULY 2017)



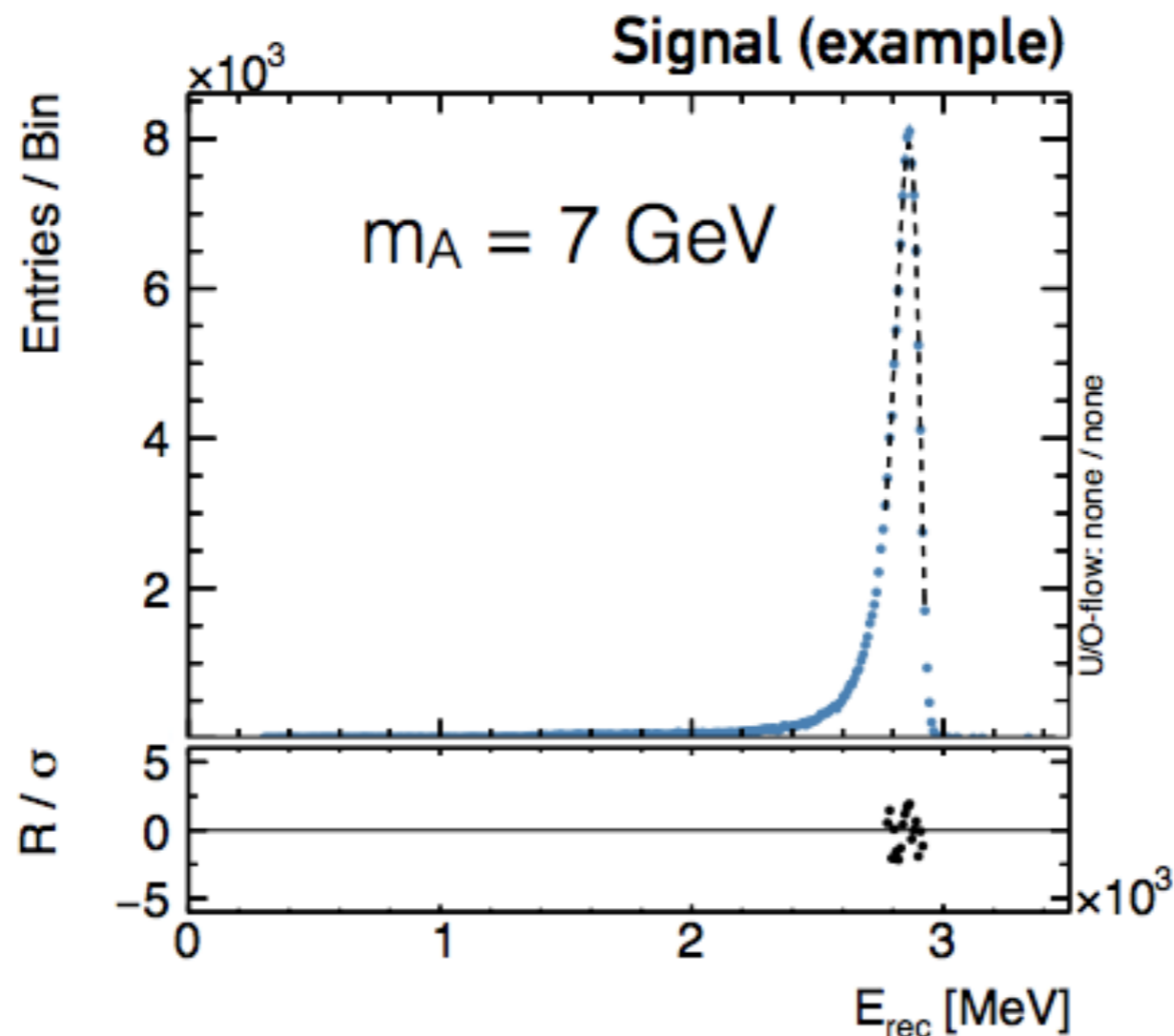
- Dedicated **2 GeV and 1 GeV (CMS Energy) threshold triggers** are developed
- The clusters in the ECL are **divided into barrel and end cap clusters** with different trigger logics applied to each
- Tighter trigger logic is applied in the end cap to suppress high background in the region



Electromagnetic Calorimeter Geometry (ECL)

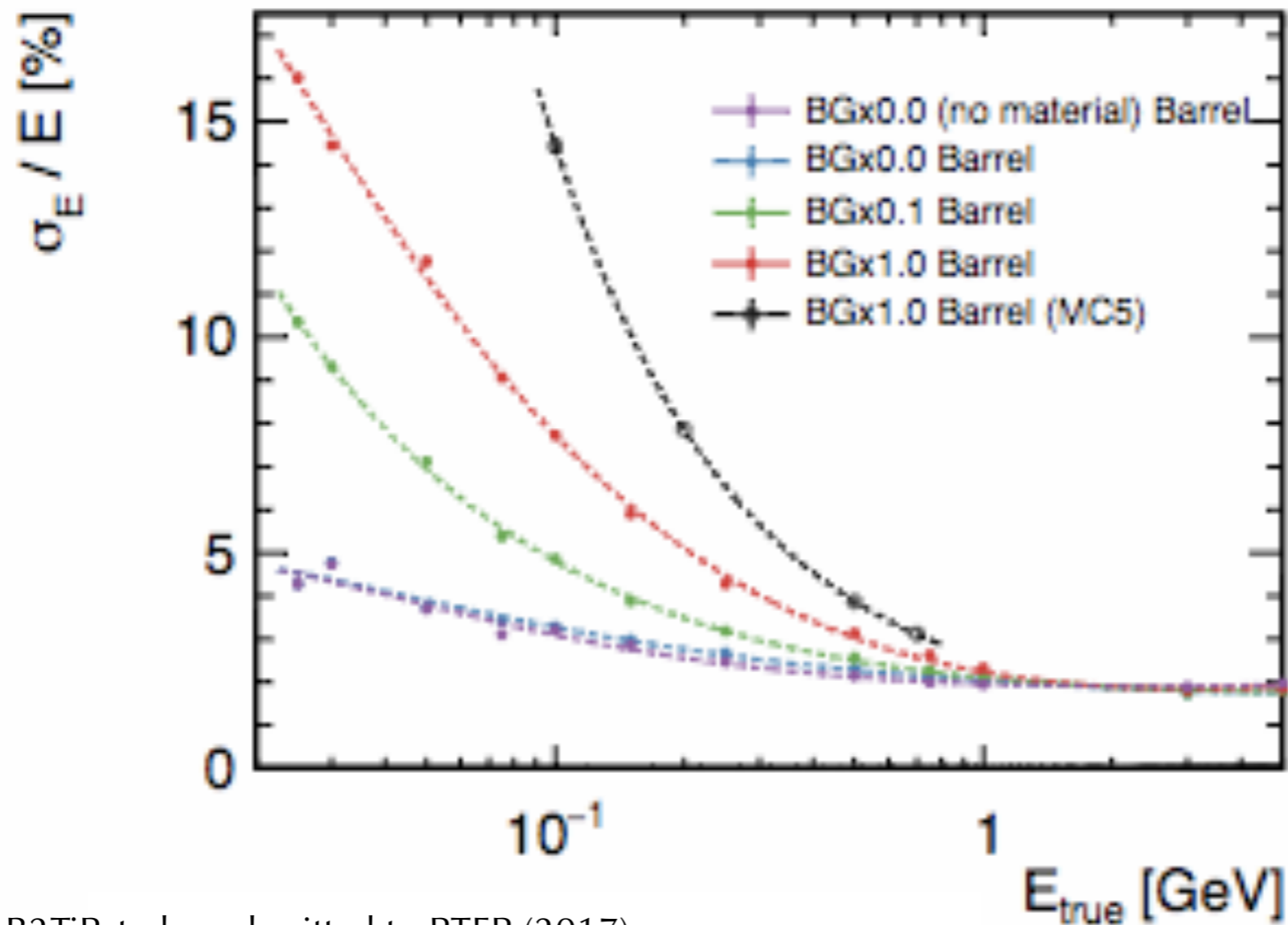
# MONOPHOTON SELECTION CRITERIA

- One photon with  $E_\gamma \geq 1.8$  GeV (selection not yet optimized)
  - Additional ECL clusters are permitted so long as  $E \leq 0.1$  GeV
- Charged particle tracks are permitted so long as  $p_T \leq 0.2$  GeV/c
- The analysis method fits the recoil mass squared distribution



T. FERBER, "EPS  
CONFERENCE: DS AT  
BELLE II" (JULY 2017)

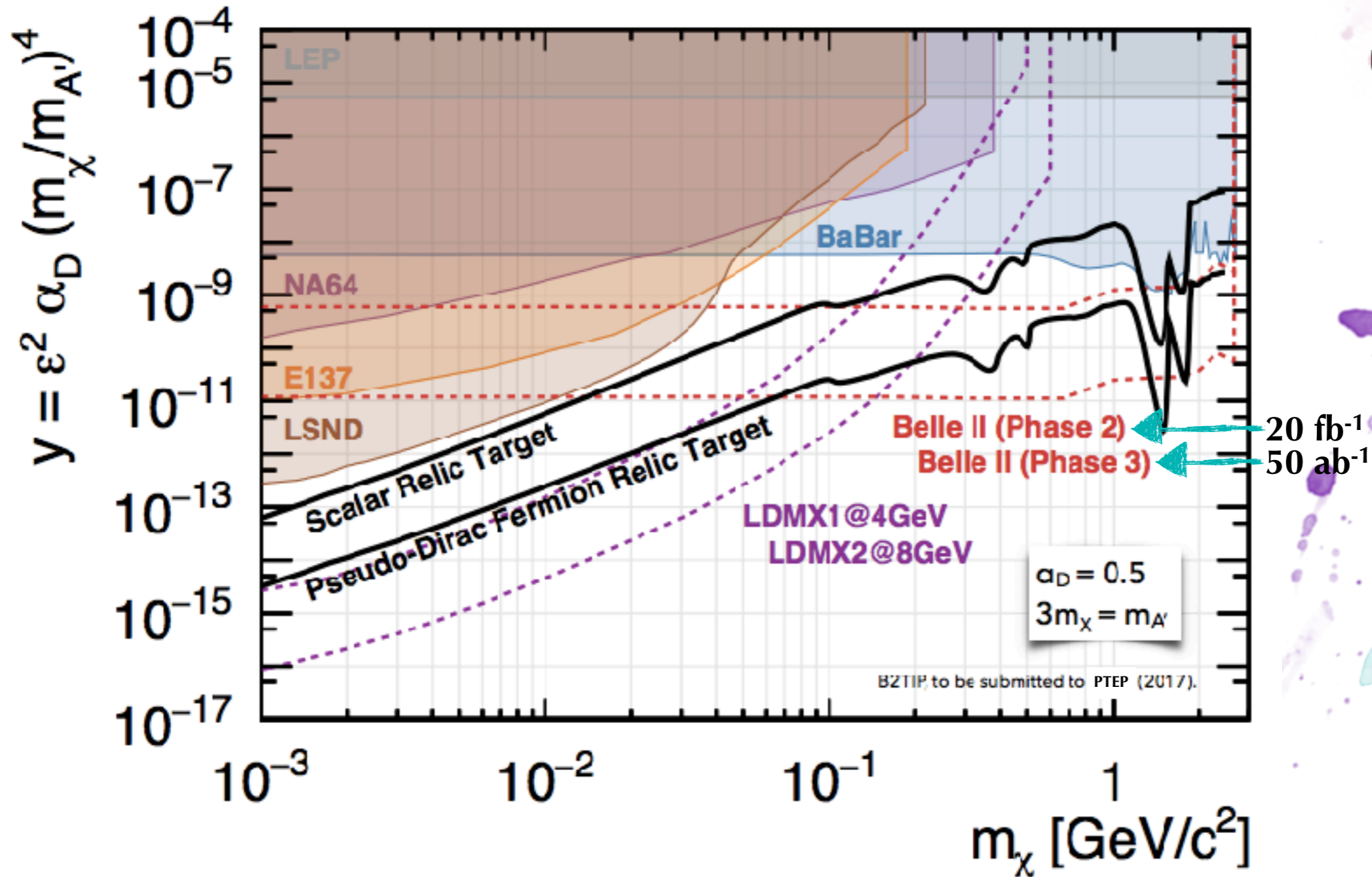
## Barrel Region Efficiency



B2TiP, to be submitted to PTEP (2017)

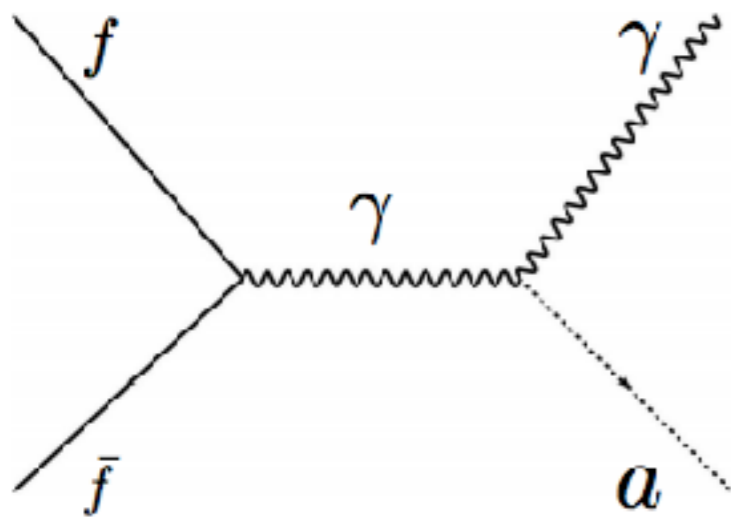
- Good photon efficiency is crucial for the monophoton signal

# PROPOSED REACH OF BELLE II: $A' \rightarrow$ INVISIBLE



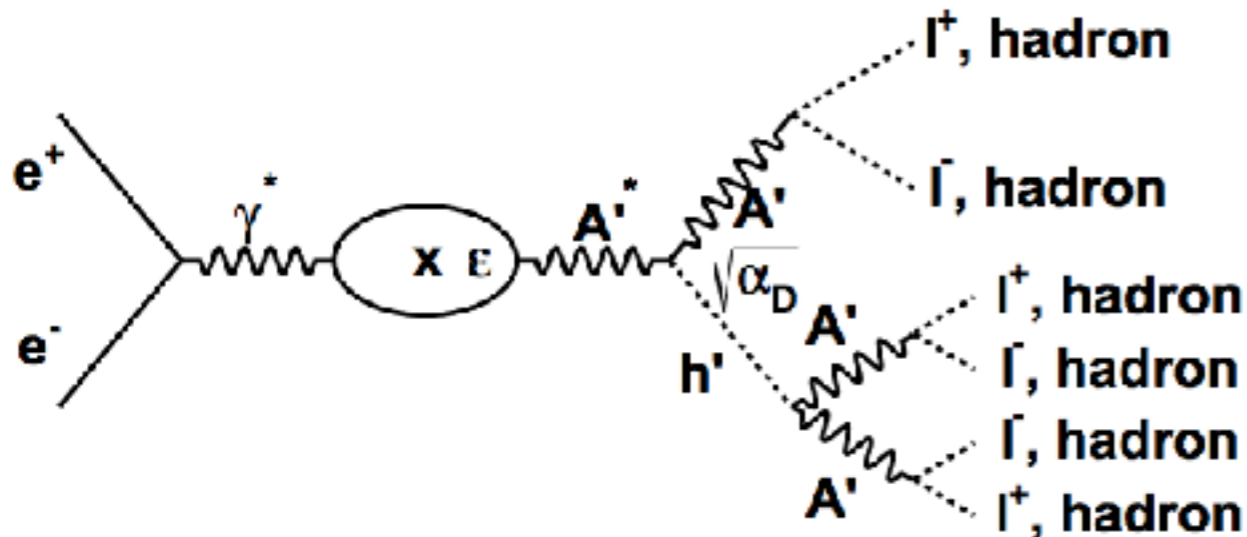
Results indicate we can perform dark photon searches for the monophoton mode in Phase 2 (starting in February 2018)

T. FERBER, "EPS CONFERENCE: DS AT BELLE II" (JULY 2017)



- Axion-like particle searches following the theory laid out in the paper by K. Mimasu and V. Sanz.

K. MIMASU, V. SANZ, ALPS AT COLLIDERS, JHEP 1506 (2015)  
 C. HEARTY, T. FERBER, B2TIP REPORT: SUBMITTED TO PTEP (2017)



- Dark Higgs searches following the work performed at Belle and BaBar.
- The dark Higgs is produced via Higgsstrahlung (replacing the SM Higgs) in these searches.

J.P. LEES ET AL., BABAR COLLABORATION, ARXIV:1202:1313 (2012)  
 I. JAEGLE ET AL., BELLE COLLABORATION, ARXIV:1502:00084 (2015)

# FINAL REMARKS

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- Belle II is preparing many dark sector searches — **Dark Photon, ALPs, Dark Higgs**
- **Phase 2 physics** data taking (February 2018,  $20 \text{ fb}^{-1}$ ) provides an opportunity to examine dark sector searches
- Dark sector searches at Belle II rely on implementation of **unique trigger logic**

# CONTACT INFORMATION

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**Caitlin MacQueen**



[cmq.centaurus@gmail.com](mailto:cmq.centaurus@gmail.com)



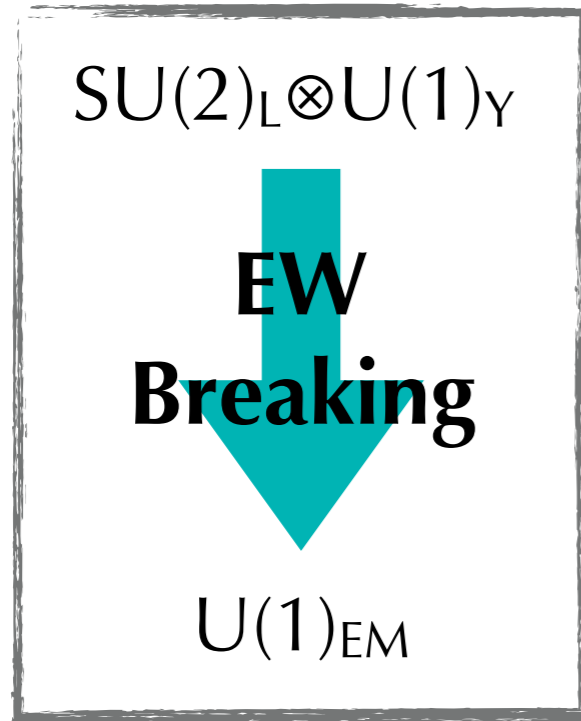
[@cmqcentaurus](https://twitter.com/cmqcentaurus)



# BACKUP SLIDES

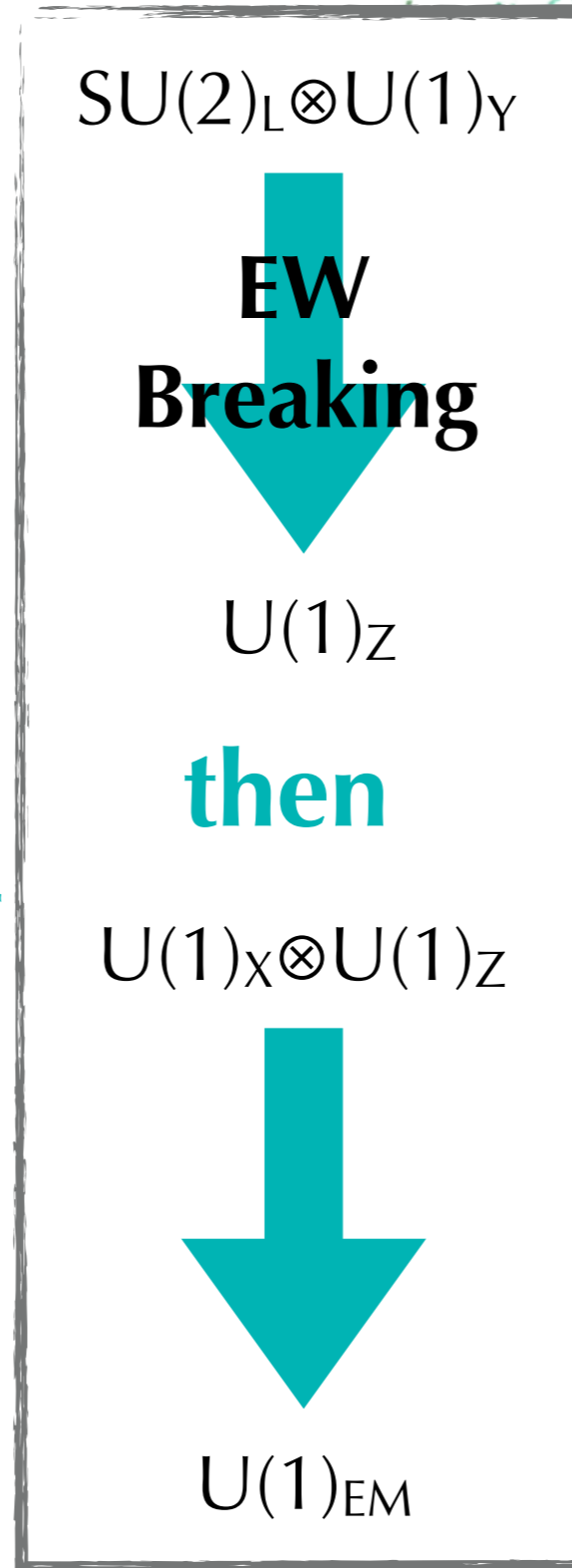


Standard Model

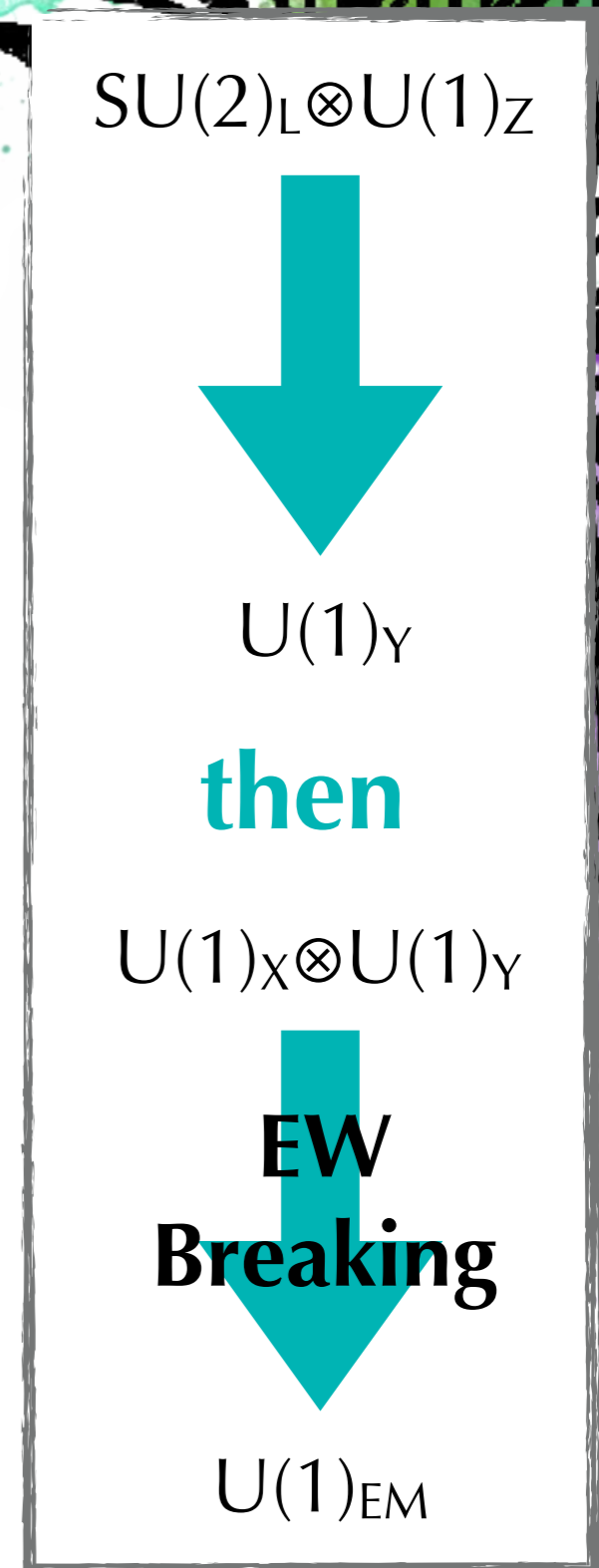


- ▶  $U(1)_X \otimes U(1)_Z$  undergoes SSB similar to that of  $SU(2)_L \otimes U(1)_Y$  in the SM.
- ▶ This gives rise to mass eigenstates  $A$  and  $A'$ , where  $m_A = 0$  and  $m_{A'} > 0$ .

My Model



Alternative Model



B. BRAHMACHARI, A. RAYCHAUDHURI, NUCL PHYS B:887 (2014)