#### Belle II readiness for Phase II

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University of Victoria, BC on behalf of the Belle II collaboration

Lake Louise Winter Institute 18-24 February 2018

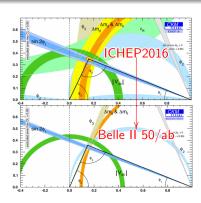
#### Belle II physics program



- B physics at the intensity frontier
  - precision measurements of CKM elements: is  $\phi_1+\phi_2+\phi_3=180^\circ?$  currently PDG gives  $(175\pm9)^\circ.$  expecting at Belle II with 50/ab  $\delta\phi_1\sim0.3^\circ,\,\delta\phi_2\sim1^\circ,\,\delta\phi_3\sim1.5^\circ.$
  - rare B decays
  - other B decay physics
  - . .
- Charm physics (Mixing, CPV in charm, rare charm decays, . . . )
- τ physics (LFV, CPV, ...)
- others
  - bottomonium spectrum
  - exotics state (tetraquark, . . . )
  - dark photon searches
  - other new physics searching (Higgs BSM, ALP, leptoquark, ...)

Hermetic detector  $\Rightarrow$  advantage for missing energy studies.

*B* meson can be identified by the full decay chain reconstruction.

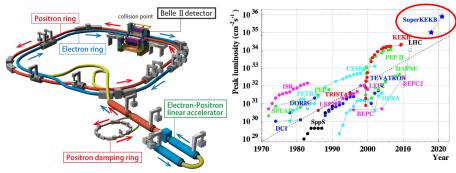


#### Phase II physics plans

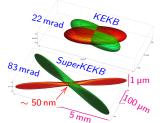
- Bottomonium spectroscopy
- Dark photon searches

#### Liminosity of KEKB and SuperKEKB





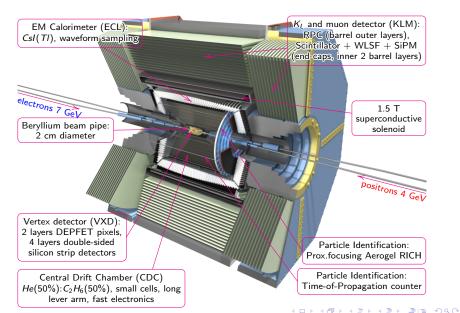
Nano-beam scheme by P.Raimondi:



	KEKB achieved	SuperKEKB nano-beam	
	LER/HER	LER/HER	
E <sub>beam</sub> (GeV)	3.5/7.5	4.0/7.0	$\beta\gamma\sim 2/3$
I <sub>beam</sub> (A)	1.6/1.2	3.6/2.6	×2
$eta_{\gamma}$ (mm)	5.9/5.9	0.27/0.30	×20
$\mathcal{L}$ (Hz/cm <sup>2</sup> )	$2.1 \times 10^{34}$	$8 \times 10^{35}$	×40

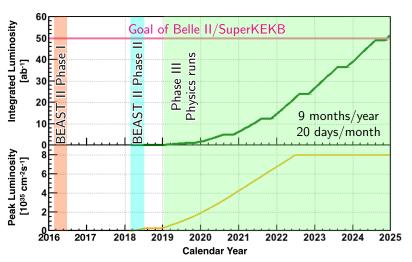
### Belle II detector (TDR arXiv:1011.0352)





# SuperKEKB luminosity projection



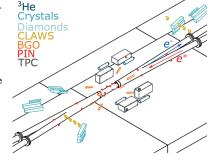


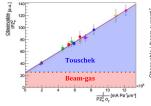
Phase II expected physics data  $-0...20 \text{ fb}^{-1}$ 

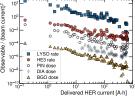
### Phase 1 February-June 2016

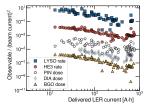


- SuperKEKB accelerator commissioning.
- Circulate both beams, no collisions, no Belle II.
- Vacuum scrubbing:
  - LER current of 1 A, dose of 780 A·h,  $10^{-6}$  Pa; HER current of 0.87 A, dose of 660 A·h,  $10^{-7}$  Pa;
    - Stable high current beams were achieved!
- Special set of detectors (BEAST II) to measure beam backgrounds important in Physics Phase (Touschek intra-beam scattering, beam-gas Coulomb, beam-gas bremsstrahlung, injection background, beam dust)
- Results submitted to NIMA: "First Measurements of Beam Backgrounds at SuperKEKB", arXiv:1802.01366.





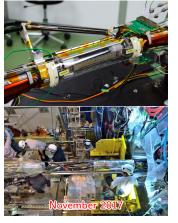




### Phase 2 February-July 2018







#### Phase II goals

- Almost final setup: Belle II is at IP with 1.5 T magnetic field + final focusing solendoids
- Accelerator commissioning
- Radiation safe environment for the full VXD
- First collisions at mid-April
- $\bullet$  Aim  $10^{34} \, \mathrm{Hz/cm^2}$  and 20 fb<sup>-1</sup>
- Further study of collision dependent beam backgrounds with radiation monitors in the VXD volume:
  - One sector of 2 PXD and 4 SVD layers where the highest backgrounds are expected.
  - FANGS, FE-I4 based hybrid pixel to investigate the Synchrotron Radiation (SR) background.
  - CLAWS, scintillators with SiPM to measure the time evolution of the injection background.
  - PLUME, double-sided high granularity MIMOSA pixels
- Outer detectors commissioning and performance study
- Possible first physics results in the bottomonium spectroscopy and the dark sector

### Vertex Detector (VXD)



#### Challenges for vertex reconstruction

Higher backgrounds (luminosity increase, nano-beam)  $\Rightarrow$  higher occupancy Boost reduced from  $\beta\gamma=0.42$  to  $0.28\Rightarrow B$ -meson flight length of 125 µm

#### Pixel Detector (PXD)

- 40 DEPFET modules into 2 layers;
- Pixel size: 50 × [55, 60, 70, 85] μm<sup>2</sup>;
- Occupancy: 0.4 hits/μm²/s (3% max);
- Integration time: 20 μs (rolling shutter);
- Thickness: 75 μm, 0.21% X<sub>0</sub>/layer;

#### Silicon Vertex Detector (SVD)

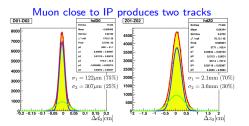
- 172 double-sided silicon strip detectors (DSSDs) in 4 layers;
- Slant shapes in FWD region for the material budget reduction;
- Thickness: 0.7% X<sub>0</sub>/layer;

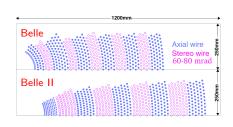


### Central Drift Chamber (CDC)



- 14336 sense wires in 56 stereo-layers in He(50%):  $C_2H_6(50\%)$  gas mixture.
- Larger compared to Belle.
- Smaller drift with sense wires and more layers allow better charged track reconstruction and dE/dx measurement compared to Belle.
- Faster readout electronics.
- CDC was installed on Oct 2016.
- Cosmic runs to test readout electronics and to validate performance.





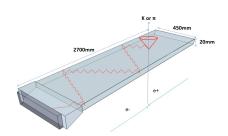


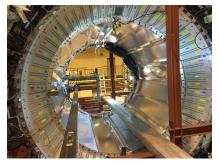
Belle2 CDC

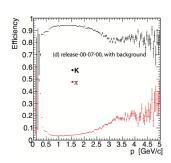
Belle CDC

# PID: TOP (Time of Propagation)





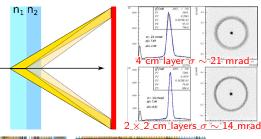




- $K/\pi$  velocity difference  $\Rightarrow$  different Cherenkov angle  $\theta_c$ .  $\theta_c$  is reconstructed from hit position (x,y) in the photo detector plane and time of propagation.
- $\bullet \ \ 16 \ quartz \ bars: \ 2 \times 125 \ cm \times 45 \ cm \times 2 \ cm$
- 32 (segmented anode 4 × 4)
  Micro-channel plate PMTs Hamamatsu
  SL-10 MCP PMT
- May 2016 TOP is integrated into the Belle II detector.

#### PID: ARICH (Aerogel Ring Imaging Cherenkov detector)





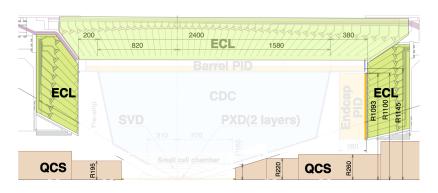




- Non-homogeneous radiator of two 2 cm aerogel tiles ( $n_1 = 1.046$  and  $n_2 = 1.056$ ) to reduce the emission point uncertainty.
- ullet 4 cm aerogel is enough to register  $\sim 10$  photons.
- 420 Hybrid Avalanche Photo Detectors, 144 channels each, 5 mm pixelated.
- Oct. 2017 ARICH is fully integrated into the Belle II detector.

### Electromagnetic calorimeter (ECL)





# ECL re-uses Belle's CsI(TI) crystals and mechanical structure

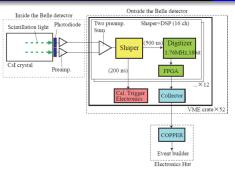
- $6 \times 6 \times 30 \text{ cm}^3 (16.1 X_0)$  in length
  - 1152 in forward endcap
- 6624 in barrel
- 960 in backward endcap

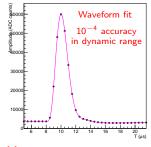
R&D to replace in future endcap crystals with pure  $\mathit{CsI}$  with faster light emission but smaller yield.

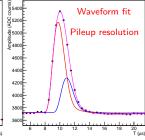
### Electromagnetic calorimeter (ECL)

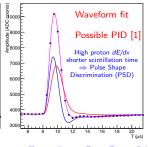


- New readout electronics with waveform sampling (18 bit @ 1.76 MHz) to cope higher rate (physics + background).
- Store ADC samples in a FPGA internal buffer. Perform waveform fit depend on trigger signal.
- Extract amplitude and timing information.
- ECL readout electronics was installed and DAQ integration tests are going on.
   Record waveforms above 50 MeV threshold in Phase II for the offline waveform fit.



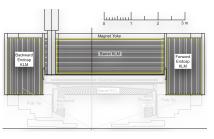


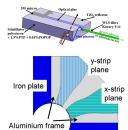




# $K_L$ and muon system (KLM)







- Sensitive layers are interleaved with the iron plates (47 mm thick) of the flux return yoke.
- 14 endcap sensitive layers (scintillator strips + WLS + SiPM)
- 2 innermost barrel sensitive layers (scintillator strips + WLS + SiPM) to resist higher rates. 13 barrel sensitive layers (double glass RPC + 5 cm orthogonal  $\phi$ , z strips)
- Installed on 2014 ongoing cosmic tests of readout electronics.



#### L1 trigger, DAQ and near term schedule



#### L1 trigger information

- Belle II Level 1 trigger (CDC + ECL + TOP + KLM)
- Beam bunch crossing rate 254 MHz (max.)
- ${\color{red} \bullet \ \, \text{Nominal beam background rate} } \\ {\color{gray} \sim 10 \, \text{MHz} }$
- $\bullet \ \, \text{Maximal L1 trigger rate} \sim 30 \, \text{kHz}$
- L1 maximal latency 5 μs
- L1 Z-vertex trigger
- L1 Global Reconstruction Logic
- $\bullet \ \, \text{Logic for single} \, \, \gamma \, \, \text{trigger for the dark} \\ \text{photon searches}$

#### DAQ

- Event size: ~35 kB with SVD + CDC + ECL (PXD ~5-20 kB in addition)
- $\begin{tabular}{ll} \hline \bullet & Tested readout from detector to \\ storage: $\sim$200 MB/sec/unit $\times$ 5 \\ units = 1 GB/sec in total (3 GB/sec max.) \\ \hline \end{tabular}$
- SVD + CDC + ECL + ARICH(partial) + TRG is confirmed to work at the rate of 30 kHz.
- PXD, TOP and KLM: work in progress (bottlenecks are identified, new firmware in preparation).
- The high-rate operation for Phase II is confirmed!

#### Short term plan

February	March	April
Global Cosmic	Beam circulation	Tuning ⇒ beam collision
with Belle Solenoid ON	$HER \Rightarrow LER$	

#### Conclusion



- Belle II is going to cover a wide range of physics goals.
- SuperKEKB and Belle II are preparing for the first collisions in April 2018.
- First 20 fb<sup>-1</sup> will allow to study detector performance, calibrate detector subsystems and reliably estimate future background conditions.
- First physics results are expected in the bottomonium spectorscopy and the dark photon search.
- The most critical backgrounds for Belle II are luminosity dependent, to be measured in Phase 2.
- Physics data taking with the fully equiped detector is planed in the begining of 2019 as well as further maching tuning to reach designed limunosity.

# Backup slides

### First cosmic events in Phase II SVD



