

Dark Matter Search with Belle II

Enrico Graziani

INFN – Roma 3

on behalf of the Belle II Collaboration



OUTLINE OF THE TALK

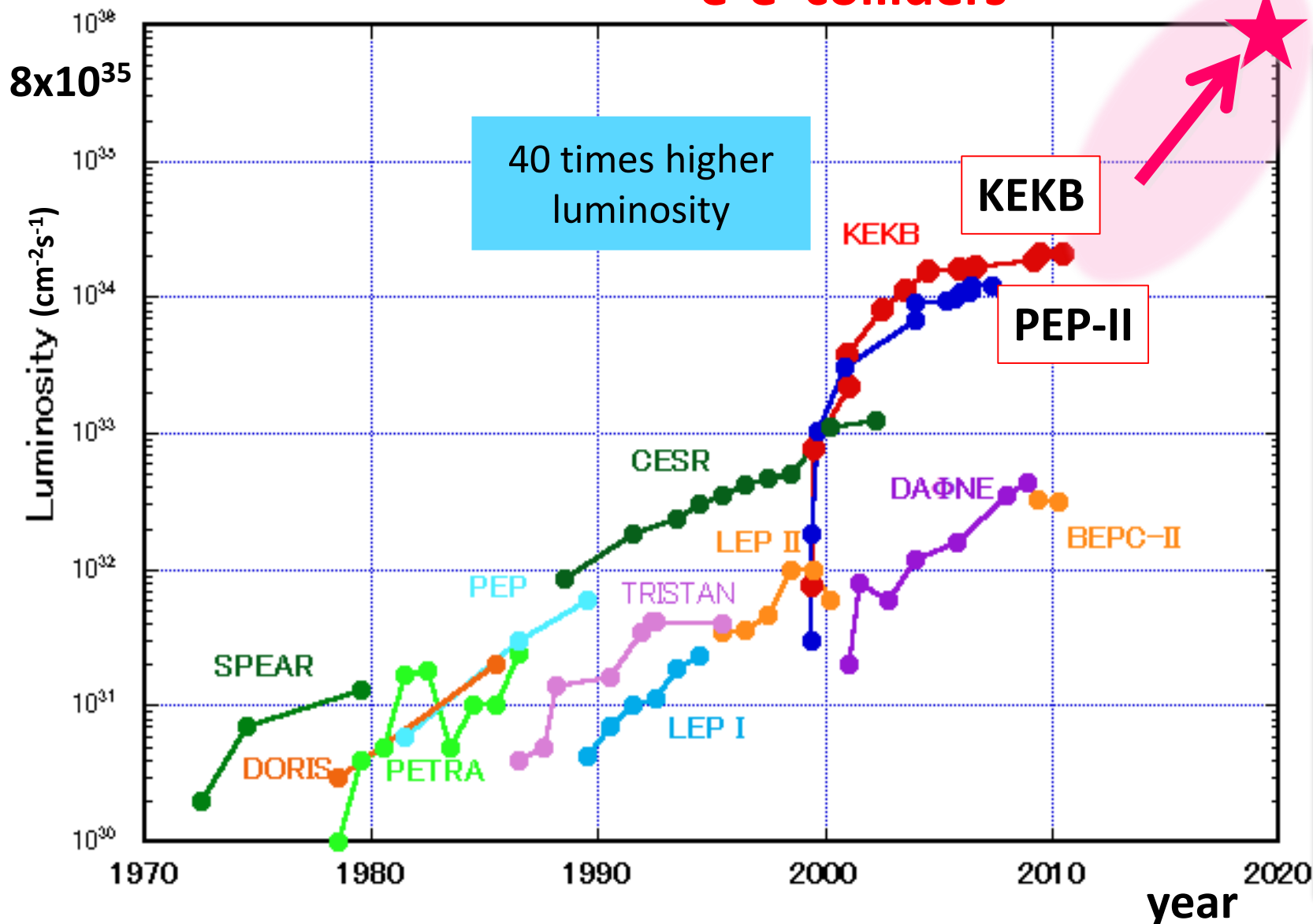
- Belle II and SuperKEKB
- Search of the invisible dark photon
- Search of ALP
- Search of Z' (invisible)
- Perspectives & Summary

A poster for the 'II South American Dark Matter Workshop'. The top left features the ICTP SAIFR logo and the text 'International Centre for Theoretical Physics South American Institute for Fundamental Research'. The main title is 'II SOUTH AMERICAN DARK MATTER WORKSHOP' in large white letters. The bottom text reads 'November 21-23, 2018 at Instituto de Física Teórica - UNESP, São Paulo, Brazil'. The background is dark with a circular inset on the right showing a colorful astronomical image of a star cluster or galaxy with various colors like blue, red, and yellow.

Peak luminosity trend

e^+e^- colliders

SuperKEKB



Very rich physics program

Flavour physics

- CKM matrix
- CPV in B decays

BSM physics

- Rare decays
- NP in loops in $b \rightarrow s\gamma$, $b \rightarrow sll$
- $B \rightarrow D^{(*)}\tau\nu$
- LFV in τ decays

New particles (quarkonium)

Dark sector

From KEKB to SuperKEKB



Beam-beam parameter

$$\xi_{y\pm} = \frac{r_e}{2\pi} \frac{N_{\mp} \beta_y^*}{\sigma_y^* (\sigma_x^* + \sigma_y^*)} R_{\xi_{y\pm}} \propto \frac{N_{\mp}}{\sigma_x^*} \sqrt{\frac{\beta_y^*}{\epsilon_y}}$$

Beam current

$$L = \frac{\gamma_{e\pm}}{2e r_e} \left(1 + \frac{\sigma_y^*}{\sigma_x^*} \right) \left(\frac{I_{e\pm} \xi_y^{e\pm}}{\beta_y^*} \right) \left(\frac{R_L}{R_{\xi_y}} \right)$$

Lorentz factor

Classical electron radius

Beam size ratio@IP
1 ~ 2 % (flat beam)

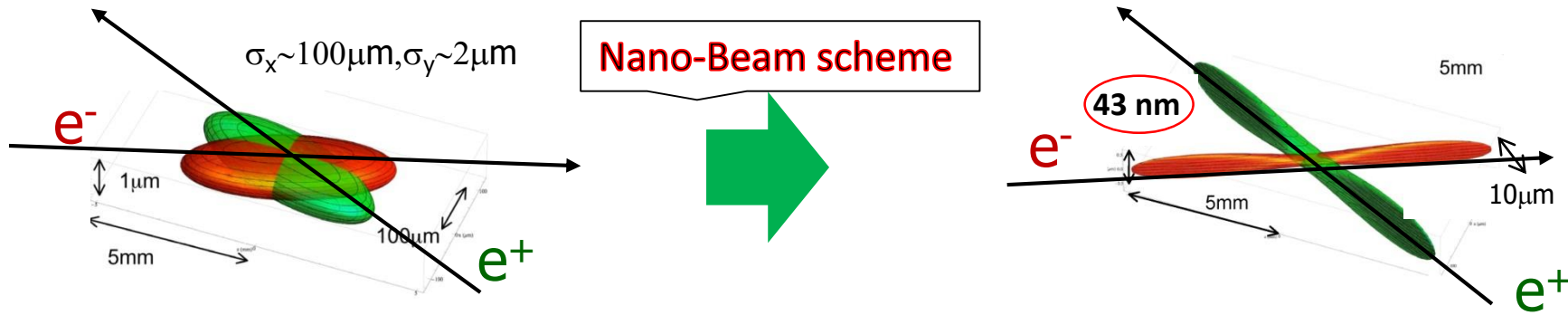
Lumi. reduction factor (crossing angle) & Tune shift reduction factor (hour glass effect)
0.8 ~ 1 (short bunch)

Vertical beta function@IP

- (1) Smaller β_y^* x20
- (2) Increase beam currents x2
- (3) Increase ξ_y

$\beta_y^* = 0.27/0.30$ mm
 $I_{+/-} = 3.6/2.6$ A

- New e⁺ Damping Ring
- New Superconducting Final Focus (QCS)



... For a 40x increase in intensity you have to make the beam as thin as a few x100 atomic layers

Belle II detector

Electromagnetic calorimeter (ECL):

CsI(Tl) crystals, waveform sampling to measure time and energy (possible upgrade: pulse-shape)
Non-projective gaps between crystals

K_L and muon detector (KLM):

Resistive Plate Counters (RPC) (outer barrel)
Scintillator + WLSF + MPPC (endcaps, inner barrel)

Magnet:

1.5 T superconducting

Trigger:

L1: < 30 kHz
HLT: < 10 kHz

Vertex detectors (VXD):

2 layer DEPFET pixel detectors (PXD)
4 layer double-sided silicon strip detectors (SVD)

Central drift chamber (CDC):

He(50%):C₂H₆ (50%), small cells,
fast electronics

Particle Identification (PID):

Time-Of-Propagation counter (TOP) (barrel)
Aerogel Ring-Imaging Cerenkov Counter (ARICH)

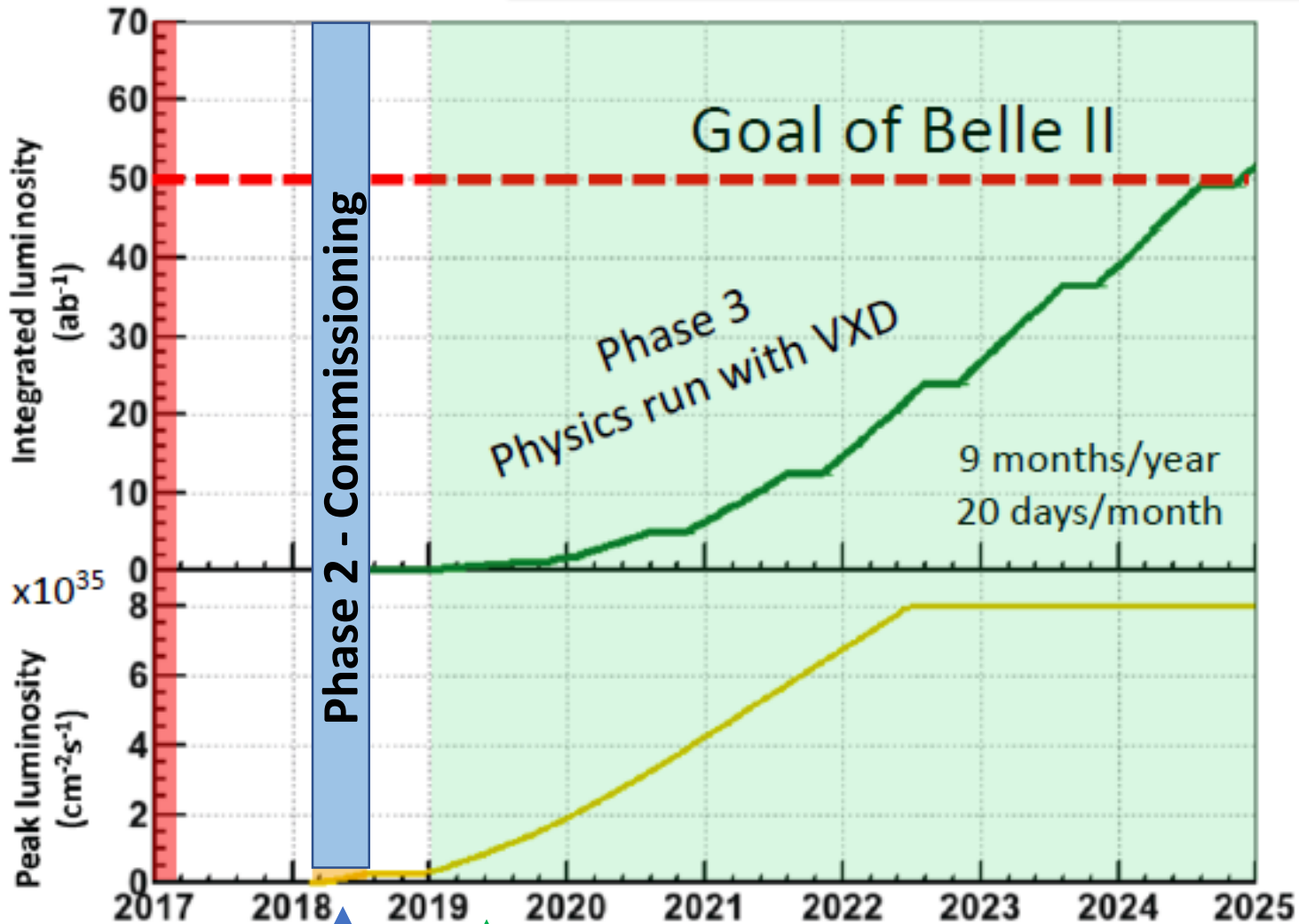
electrons (7GeV)

positrons (4GeV)

Belle II vs Belle

better resolution, PID and capability
to cope with higher background

Belle II data taking plan



Phase 2

- 1/8 of vertex detector
- Low backgrounds
- Pass-through HLT (software) trigger

Good conditions for dark searches

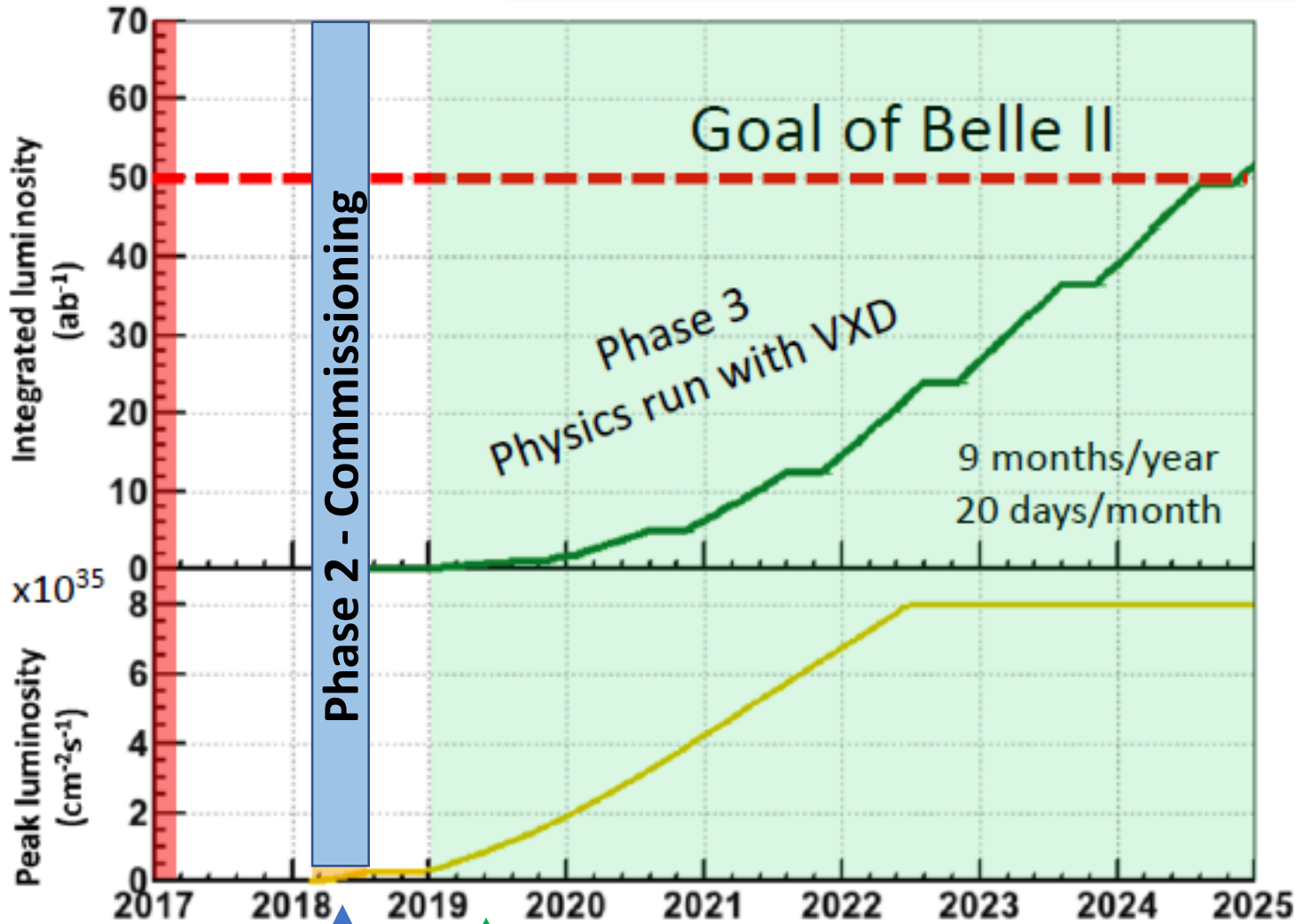
Phase 3

- $L \approx 50 \text{ ab}^{-1}$ with the full detector

Phase 2

Phase 3

Belle II data taking plan: today



Phase 2

Phase 2 finished July 17th 9 am

- Nano-beam scheme works!
- $L=5.5 \times 10^{33} \text{ cm}^{-2}\text{s}^{-1}$ achieved
- $L_{\text{int}} \approx 0.5 \text{ fb}^{-1}$ collected

- 1/8 of vertex detector
- Low backgrounds
- Pass-through HLT (software) trigger
- Tracking and clustering L1 trigger
- Bhabha veto L1 trigger
- Some single photon L1 trigger

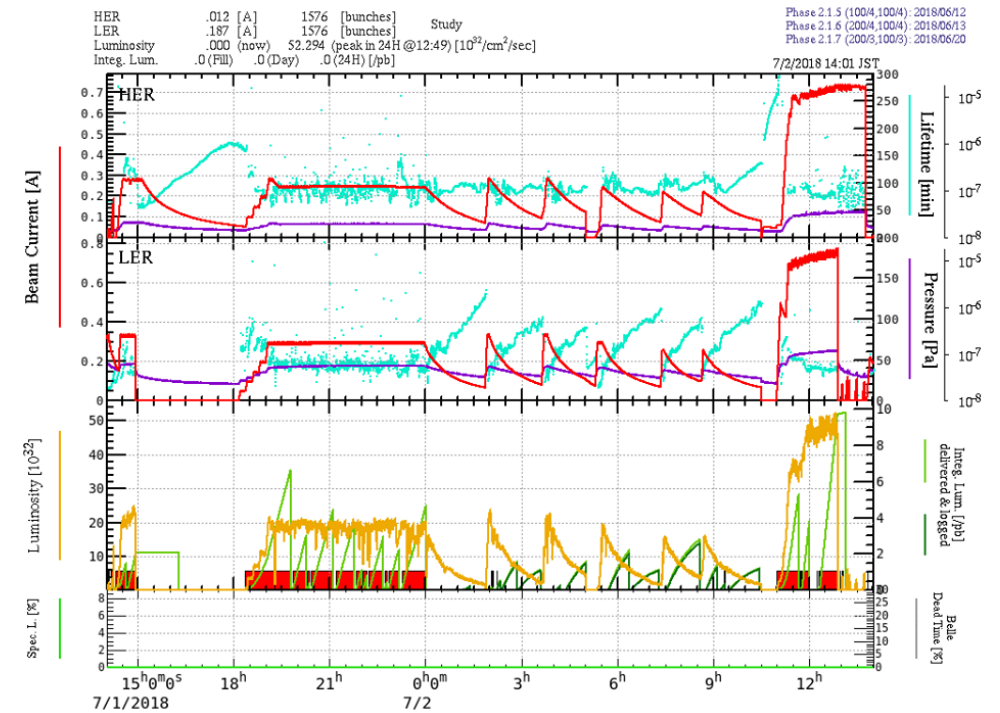
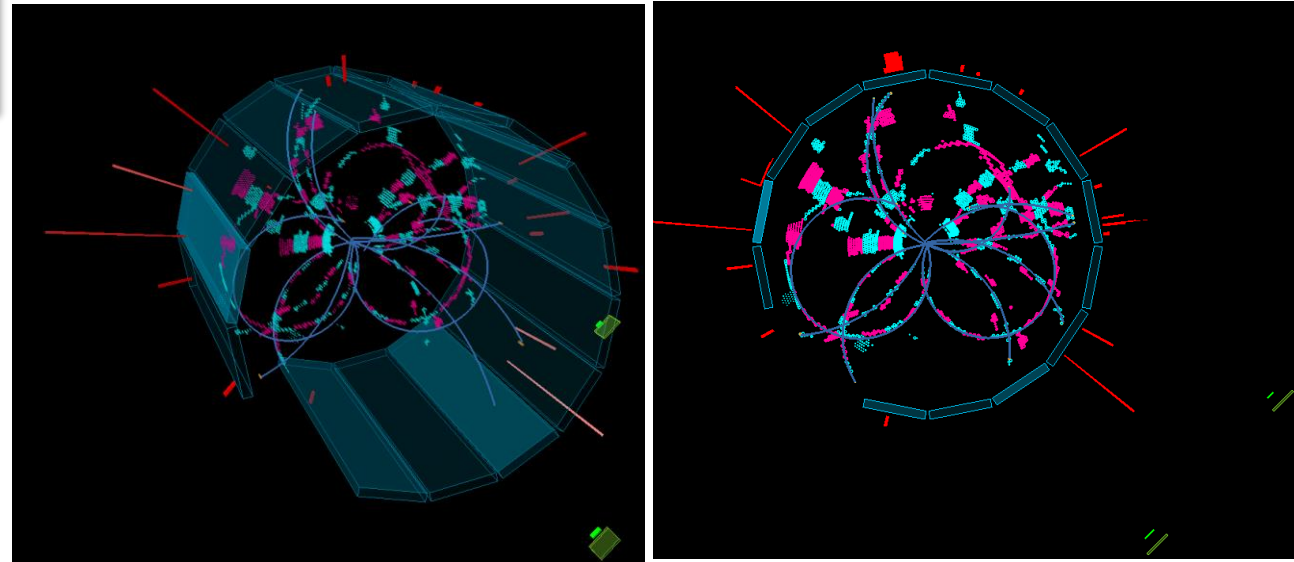
Phase 2

Phase 3

Good conditions for dark searches

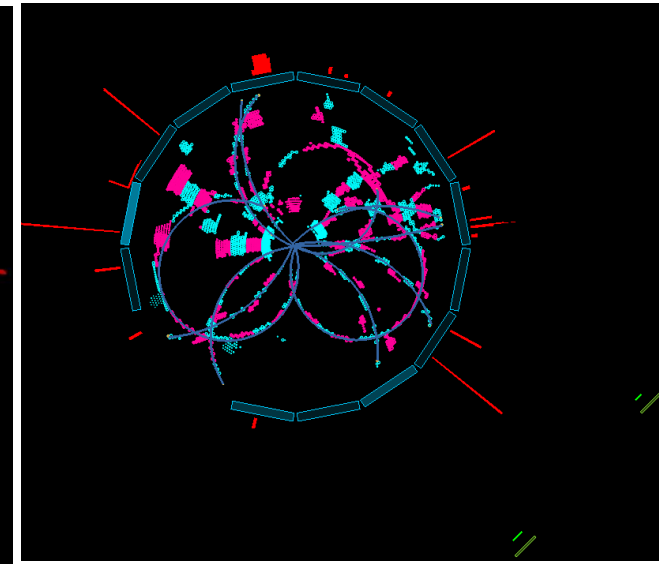
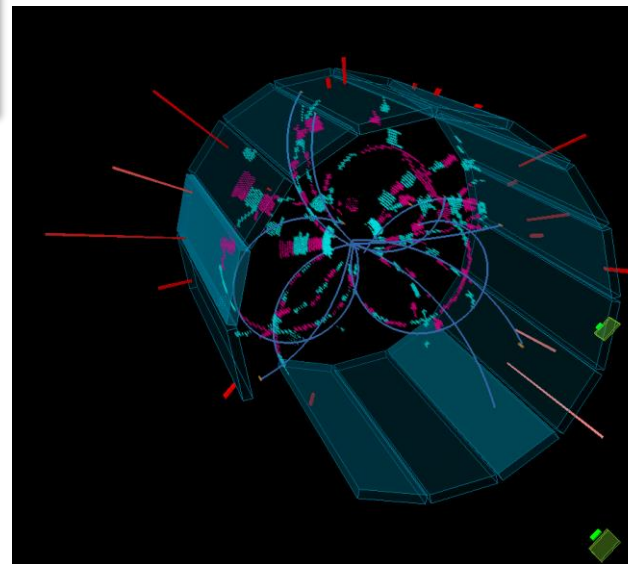
Belle II & SuperKEKB Phase 2

Start of collisions: April 25th



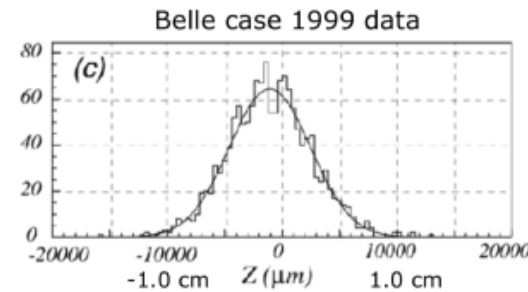
Belle II & SuperKEKB Phase 2

Start of collisions: April 25th



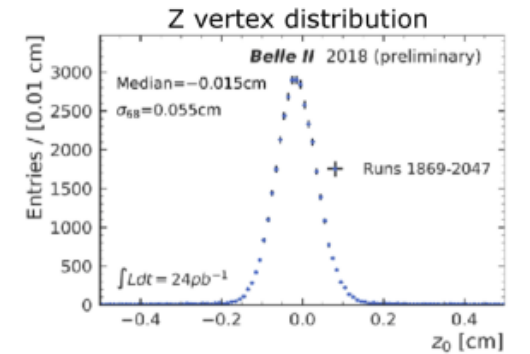
Effective bunch length: from KEKB to SuperKEKB Phase 2

Ordinary collision (KEKB)



$\sigma = 4.5 \text{ mm}$

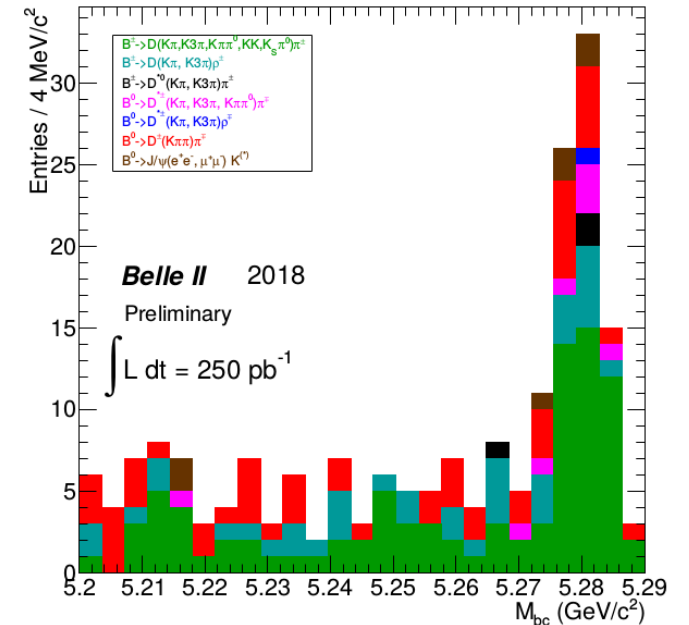
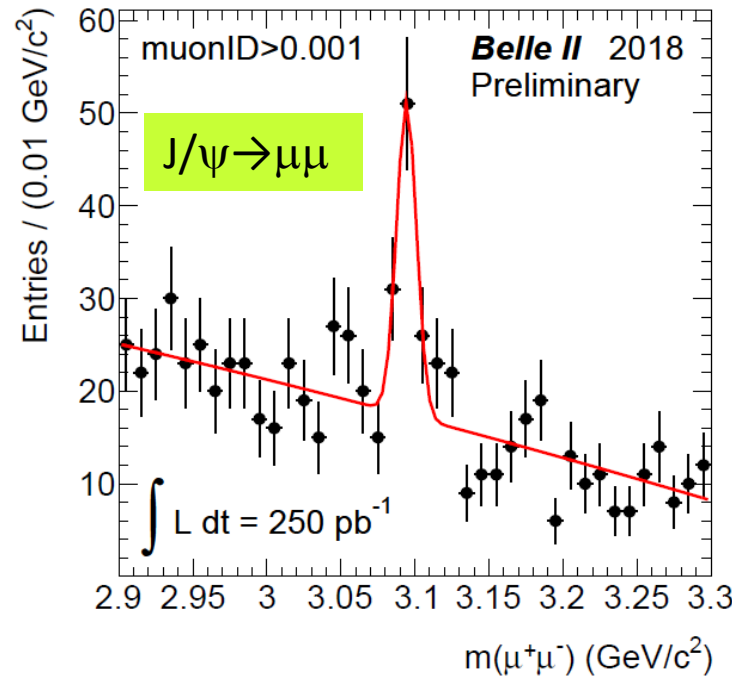
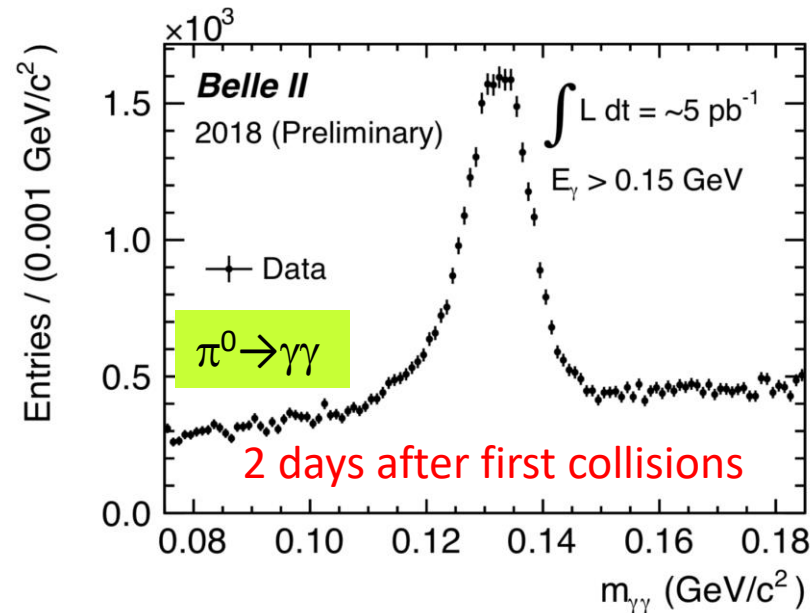
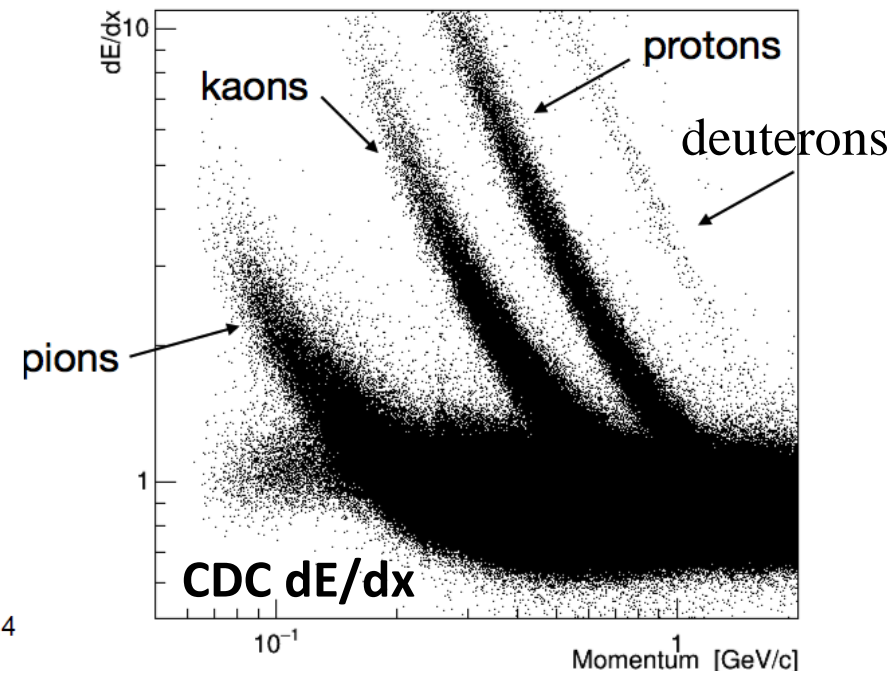
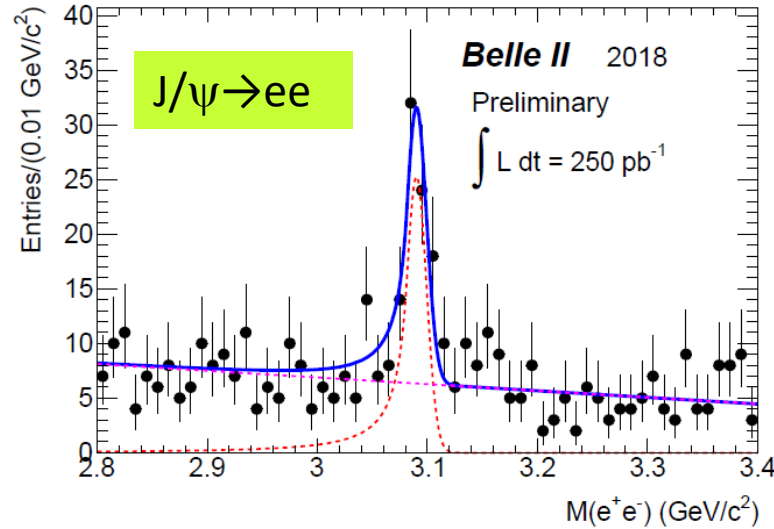
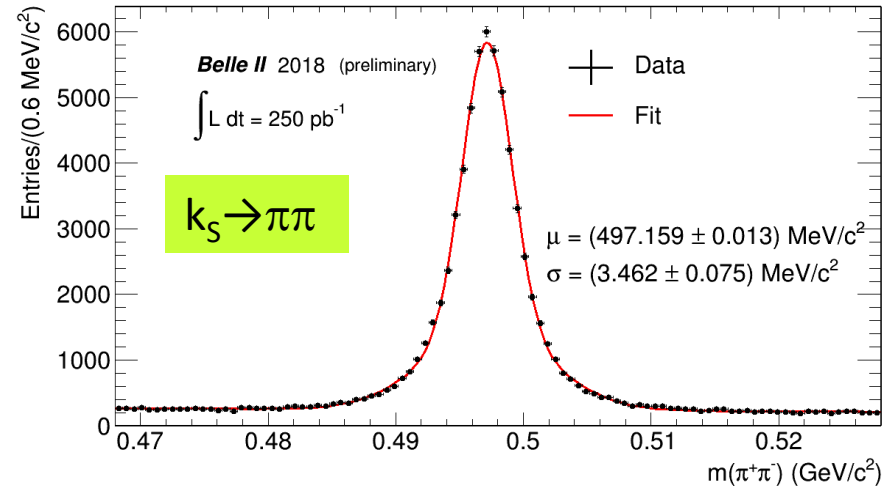
Nano-Beam (SuperKEKB Phase2)



$\sigma = 550 \text{ }\mu\text{m}$

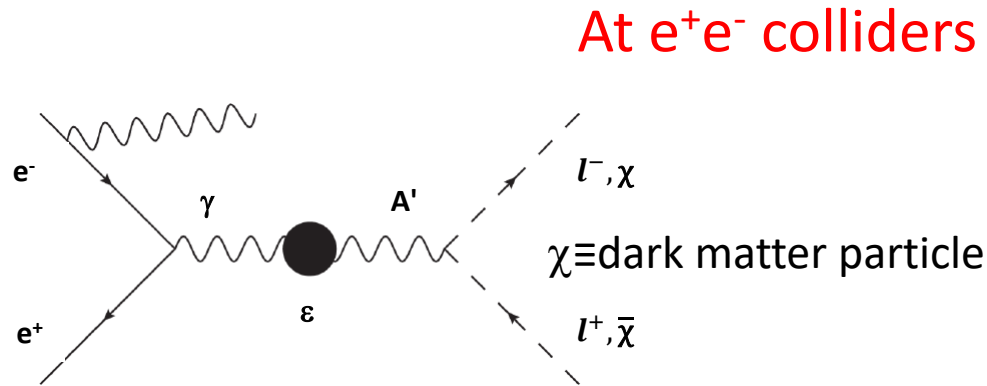
Nano-beam scheme works!

Belle II performance snapshots



Dark photon: introduction

Some astrophysical observations suggest the possibility of the existence of a new light (GeV scale) hidden dark sector with a mediator A' (dark photon), weakly coupled to the Standard Model via kinetic mixing, and light dark matter.



two basic scenarios depending on A' vs matter mass relationship

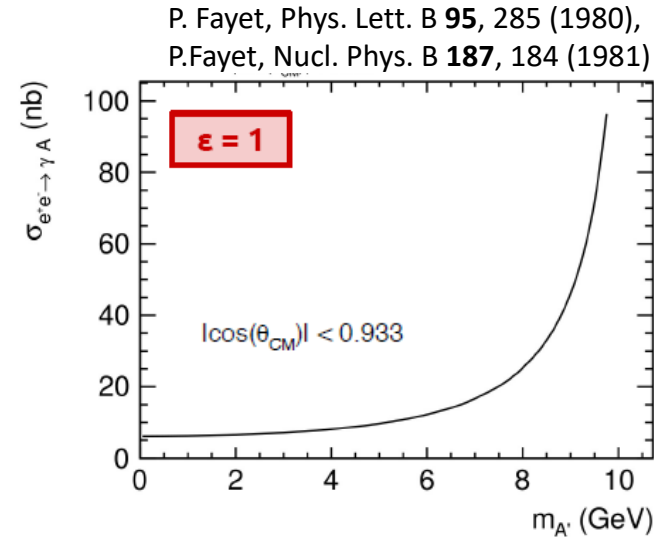
$m_\chi > 1/2 m_{A'} \Rightarrow A'$ visible decays (SM particles)

- $A' \rightarrow l^+l^-$
- $A' \rightarrow \pi^+\pi^-, h^+h^-$
- $h' A'$ dark higgstrahlung
 - $h' \rightarrow A'A', A'A'A' \rightarrow 6 l^\pm + \pi^\pm$
 - $A' + \text{missing}$

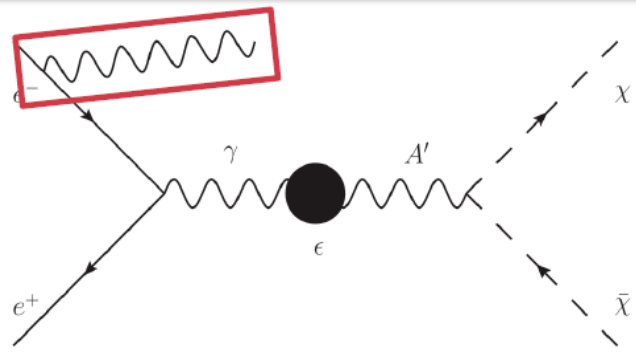
$m_\chi < 1/2 m_{A'} \Rightarrow A'$ invisible decays to LDMA

$A' \rightarrow \chi\bar{\chi}$

access to light dark matter particles



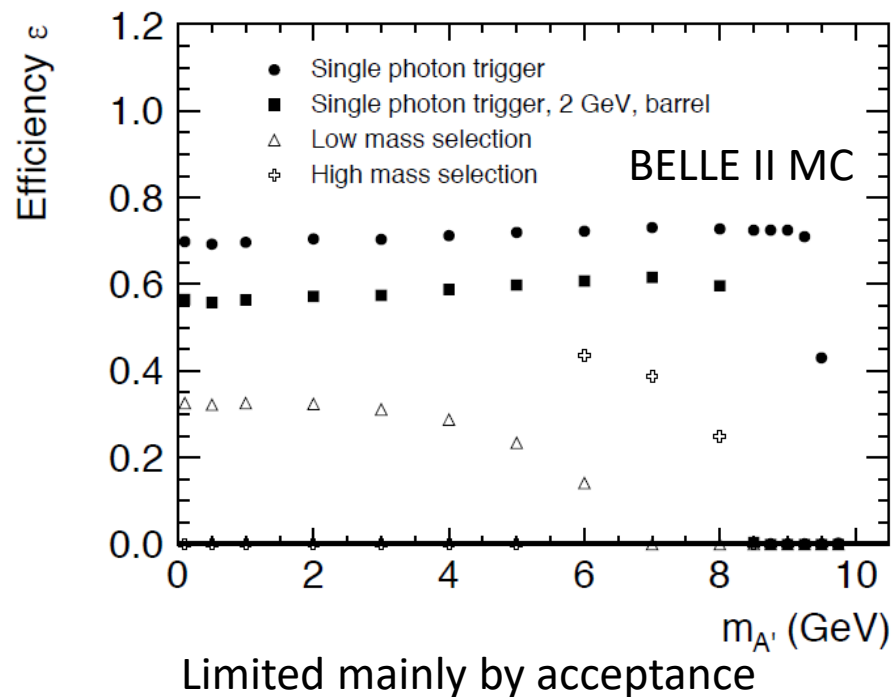
Invisible dark photon: experimental signature



Only **one photon** in the detector.
Needs a **single photon trigger**
(not available in Belle, $\approx 10\%$ of data in BaBar)

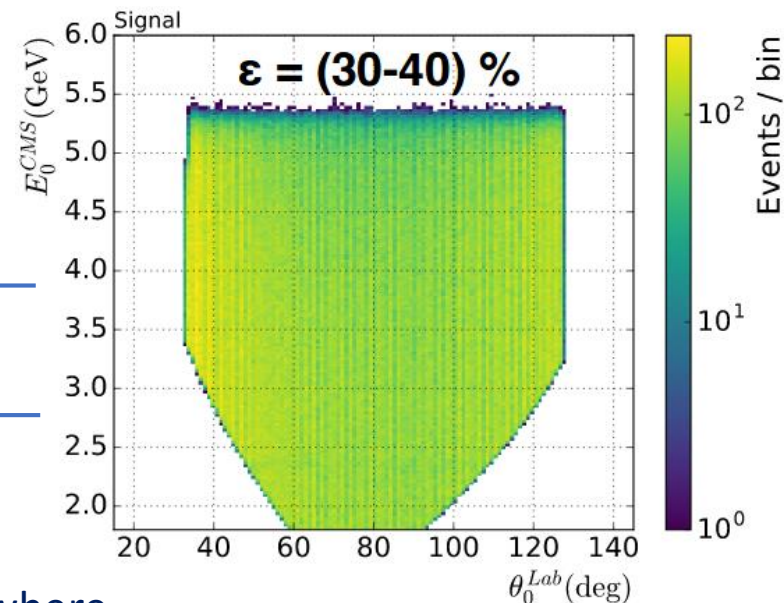
$$E_\gamma = \frac{s - M_{A'}^2}{2\sqrt{s}}$$

Bump in recoil mass or photon energy



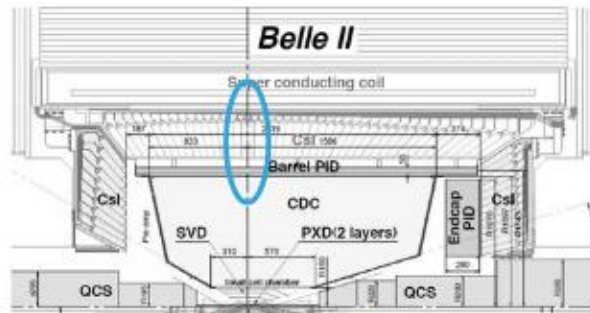
Backgrounds
 $e^+e^- \rightarrow e^+e^-\gamma(\gamma), e^+e^- \rightarrow \gamma\gamma(\gamma)$

Trigger logic	L1 rate at full luminosity
$E > 1 \text{ GeV}$	4 kHz (barrel)
+ 2 nd cluster $E < 300 \text{ MeV}$	7 kHz (endcaps)
$E > 2 \text{ GeV}$	5 kHz (barrel)
+ Bhabba & $\gamma\gamma$ vetoes	

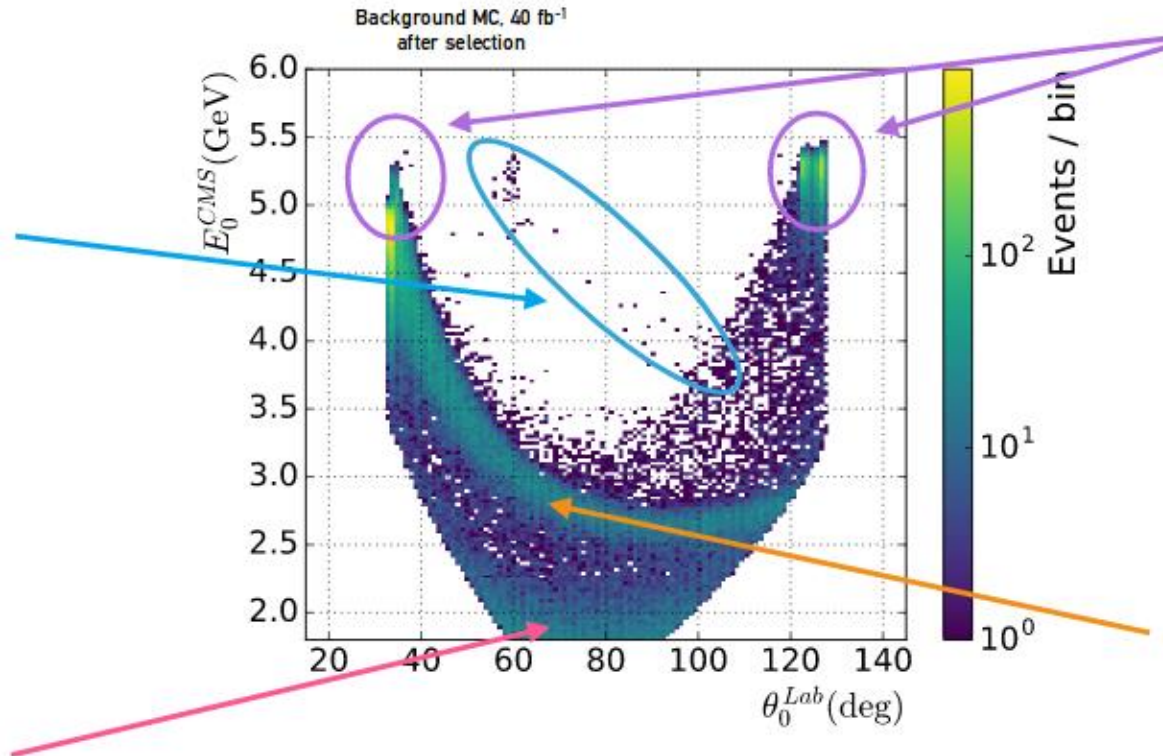


Probably not sustainable in deep Phase 3, where some prescaling or threshold adjustment will be needed

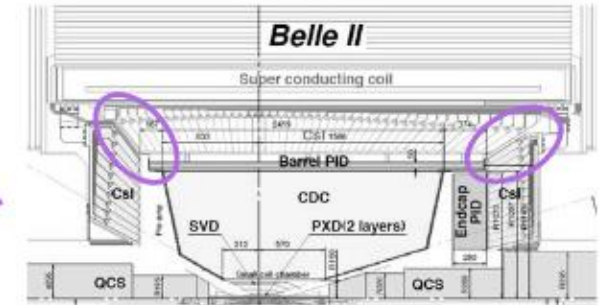
Invisible dark photon: backgrounds



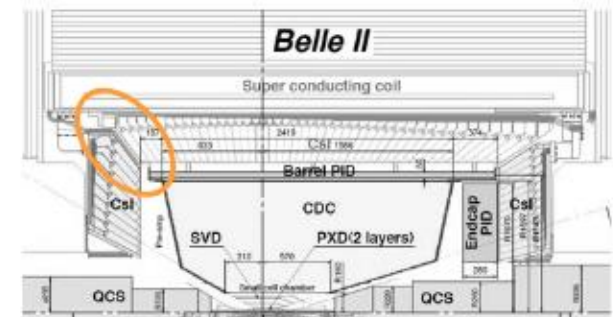
$ee \rightarrow 2\gamma$ and 3γ
 1 γ in ECL 90° gap
 1 γ out of ECL acceptance



$ee \rightarrow eey$
 both electrons
 out of tracking acceptance



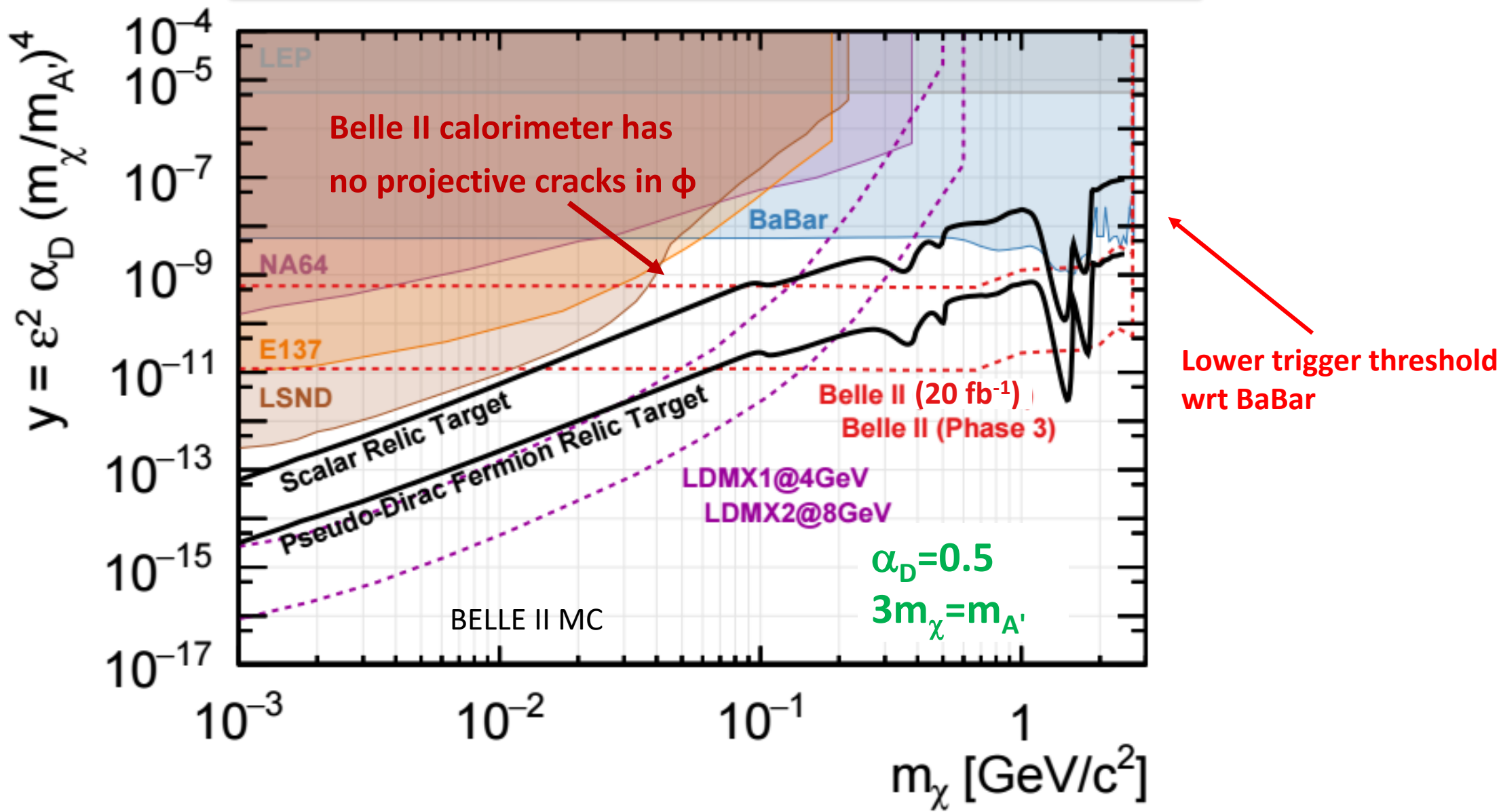
$ee \rightarrow 2\gamma$
 1 γ in ECL BWD or FWD gap



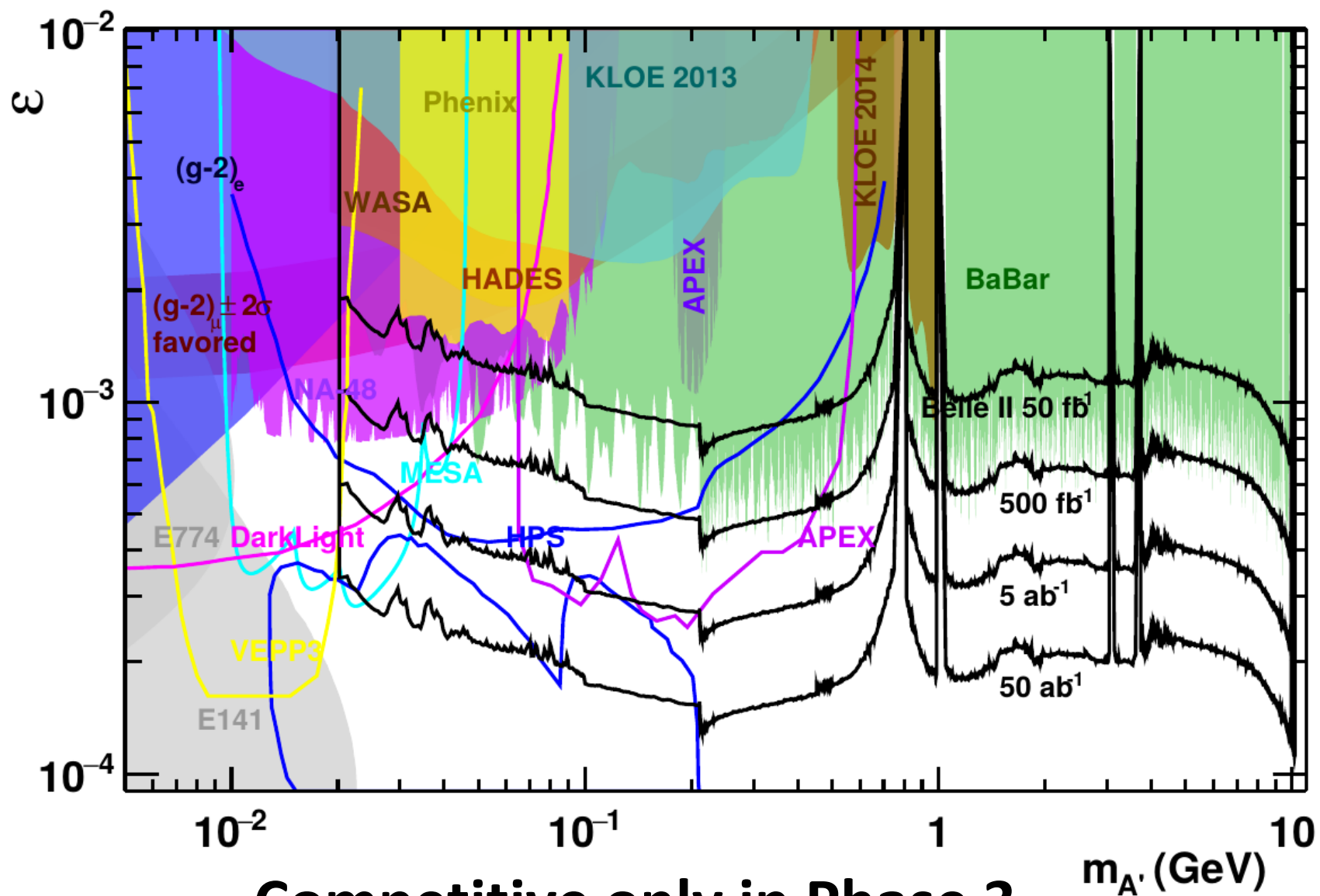
$ee \rightarrow 3\gamma$
 1 γ in ECL BWD gap
 1 γ out of ECL acceptance

Crucial usage of KLM to veto photons in ECL gaps

Invisible dark photon: sensitivity



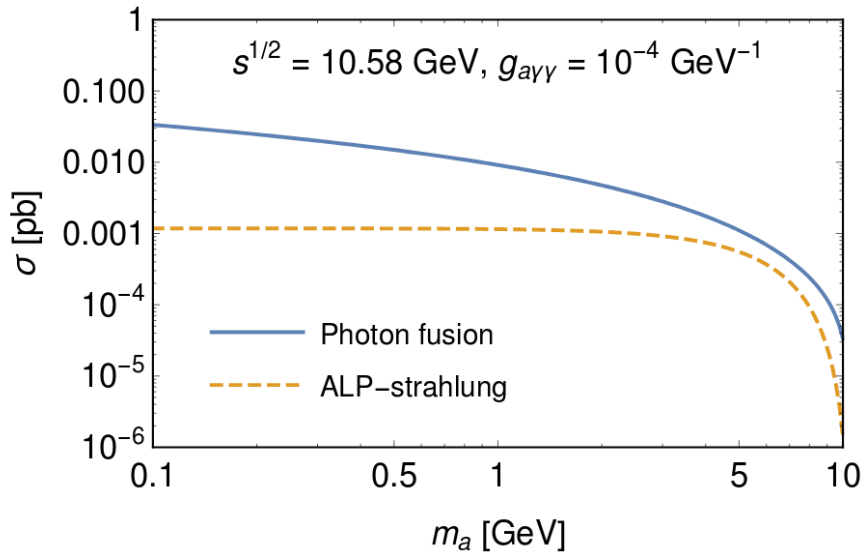
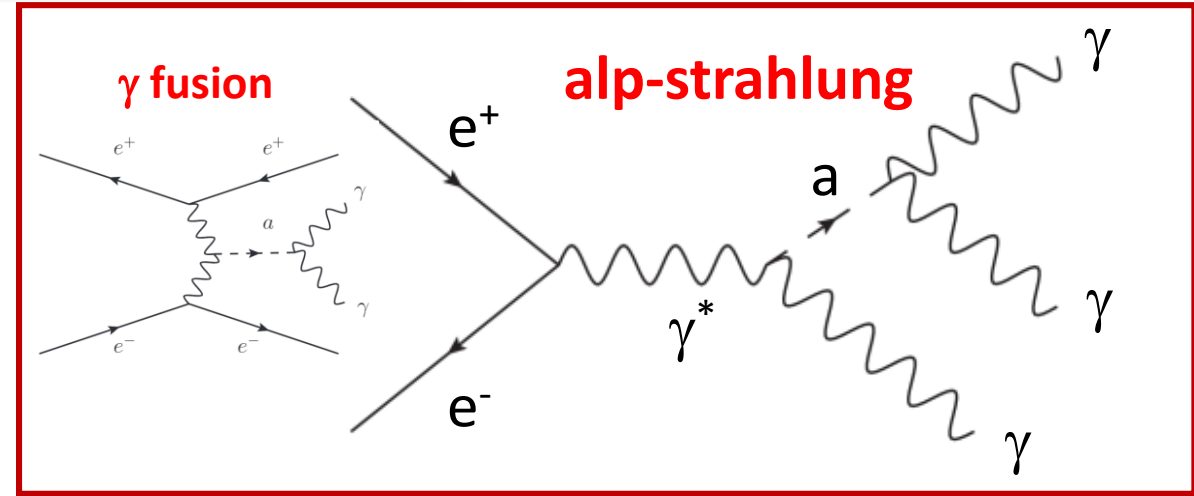
Visible dark photon: sensitivity



Competitive only in Phase 3

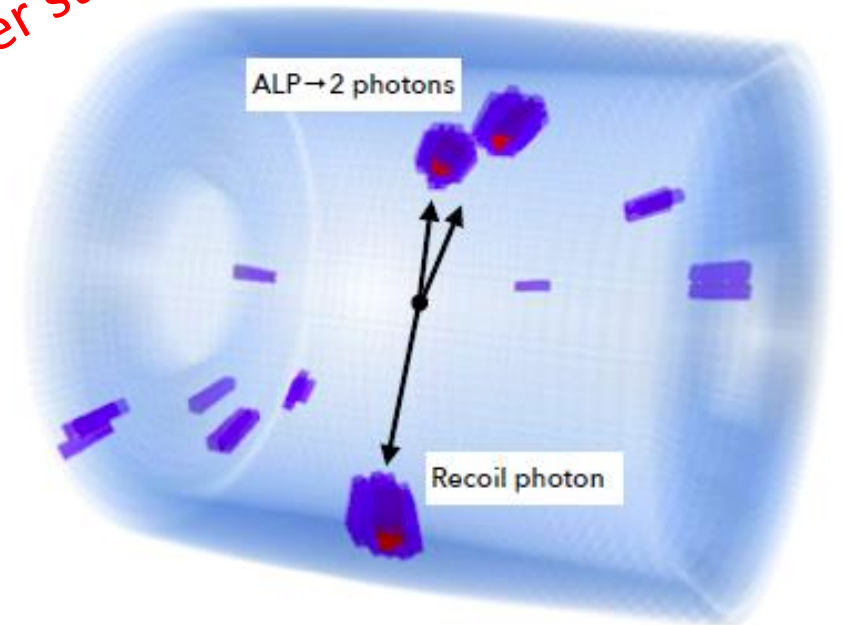
Axion Like Particles (ALPs)

- Pseudo-scalars particles which couple to bosons.
- Differently from QCD axions, no relation between mass and coupling
- Focus on coupling to photons: $g_{a\gamma\gamma}$
- Alp-strahlung + photon fusion production mechanisms
- $\tau \sim 1 / g_{a\gamma\gamma}^2 m_a^3$
- No results at B factories yet



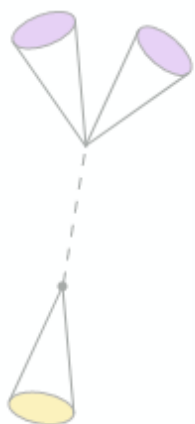
photon fusion sensitivity under study

3 γ topology

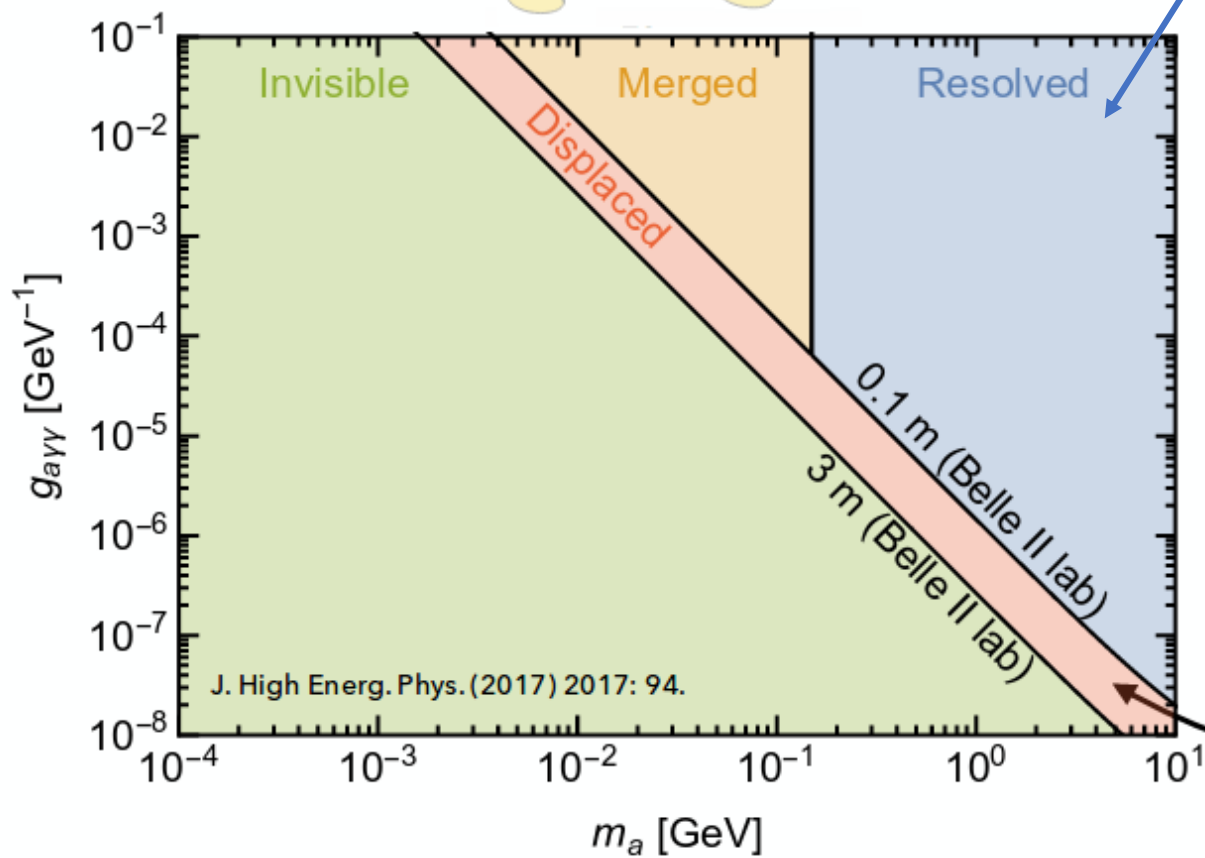
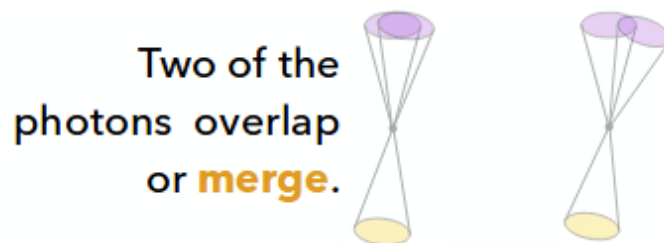


Axion Like Particles (ALPs): signal

3 γ topology, but...



ALP decays outside of the detector or decays into **invisible** particles: Single photon final state.



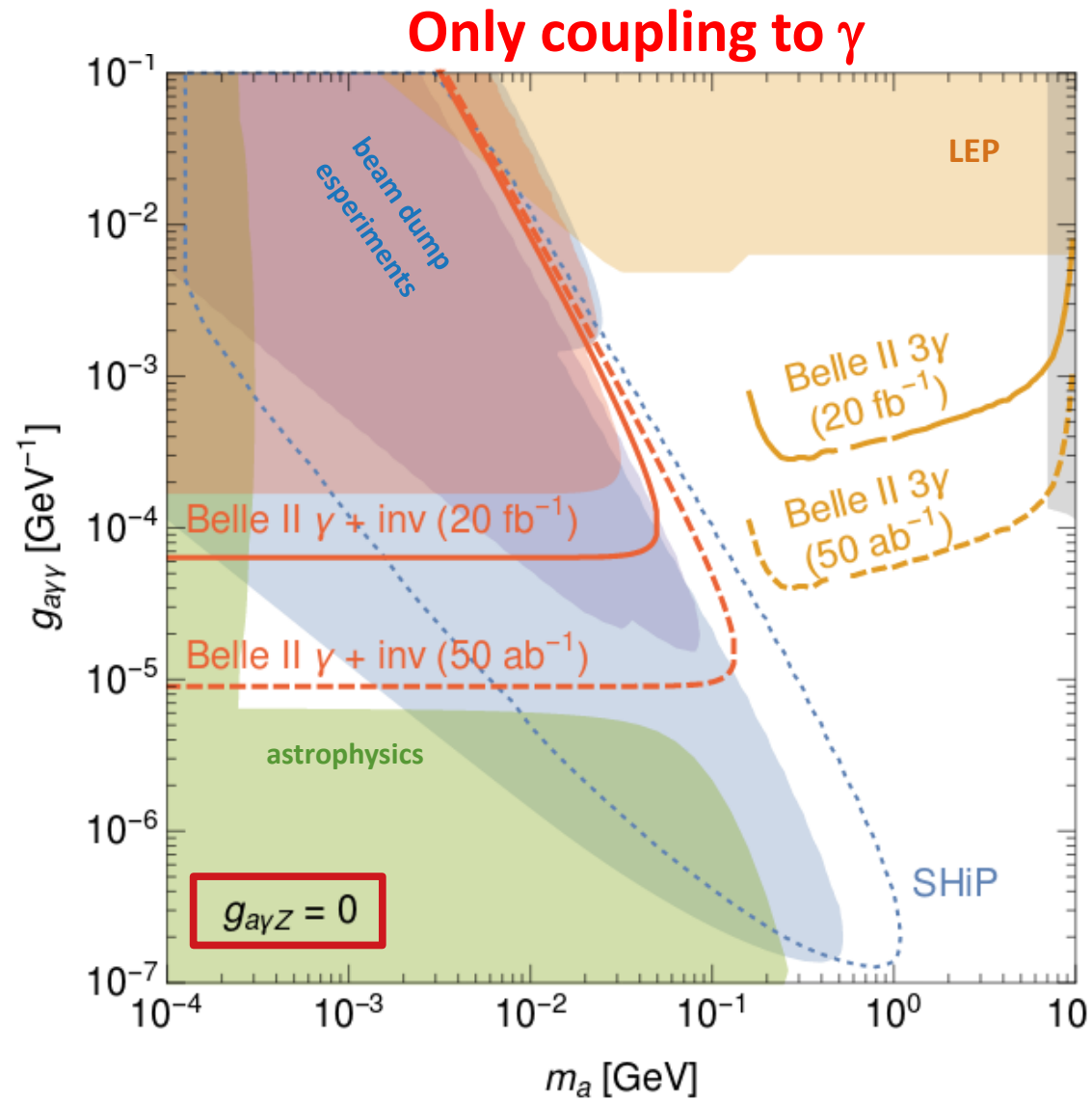
Three **resolved**, high energetic photons.



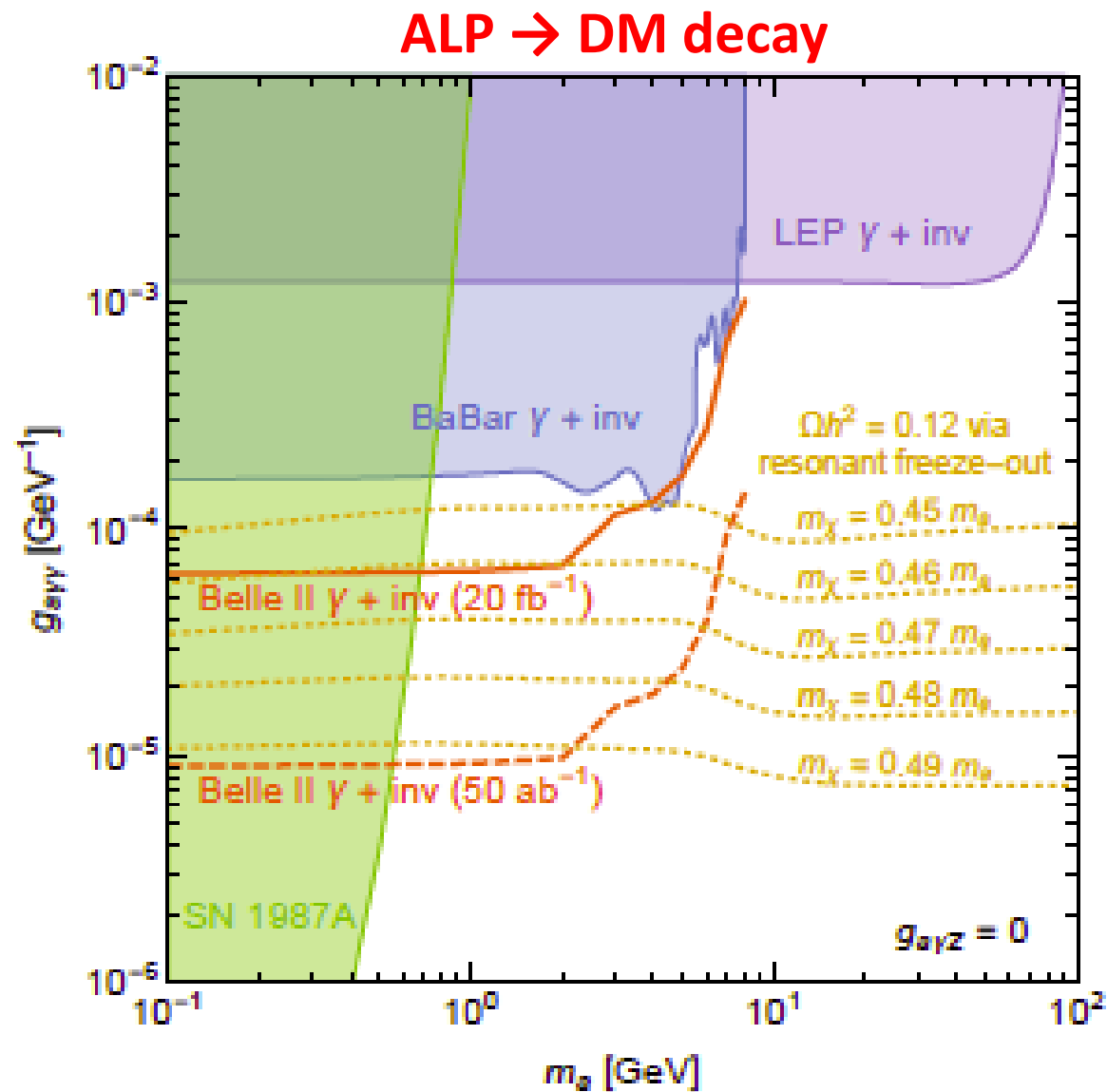
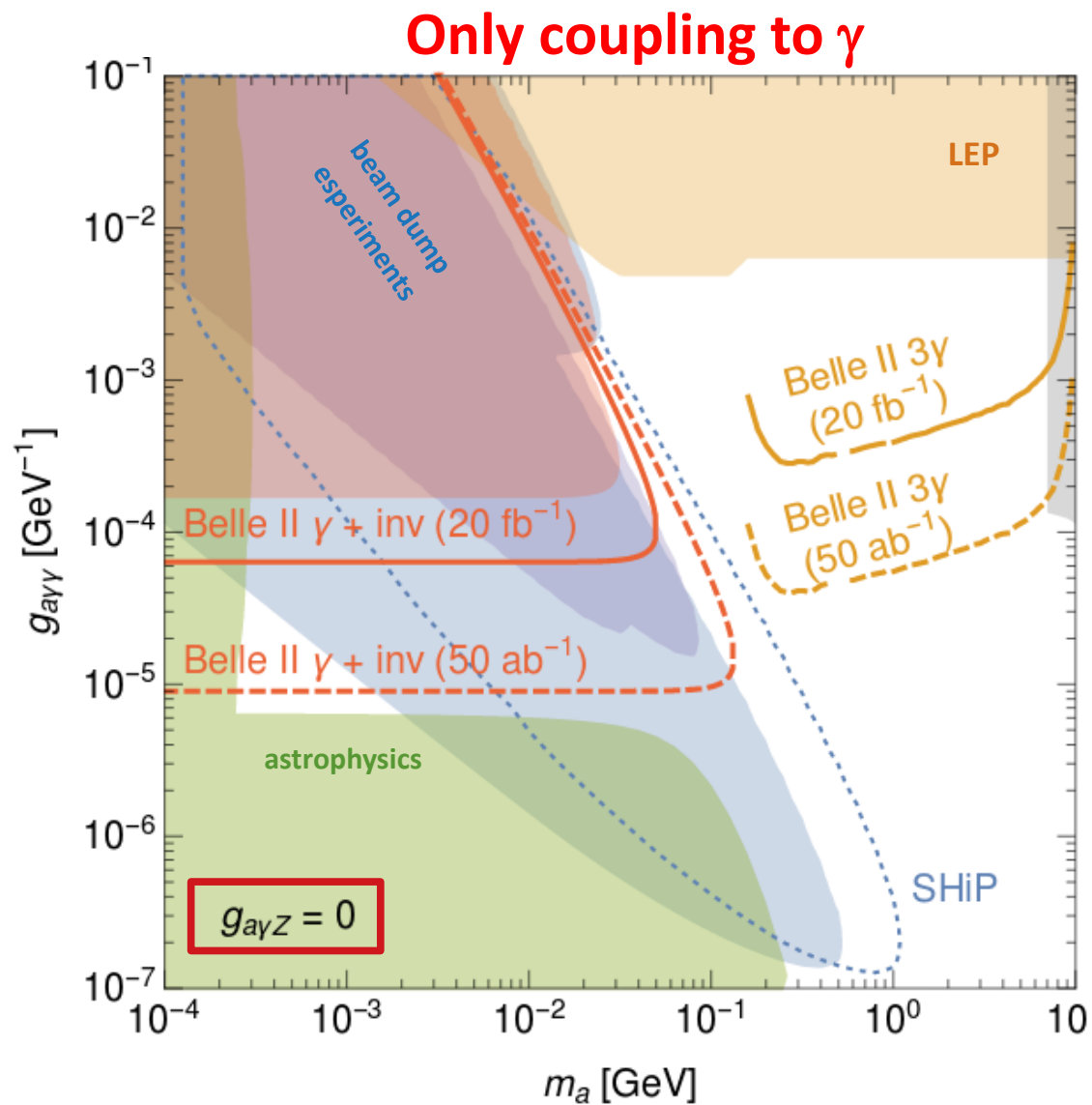
The searches for invisible and visible ALP decays veto this region.

ALPs can also decay to DM \rightarrow single photon topology

Axion Like Particles (ALPs): sensitivity

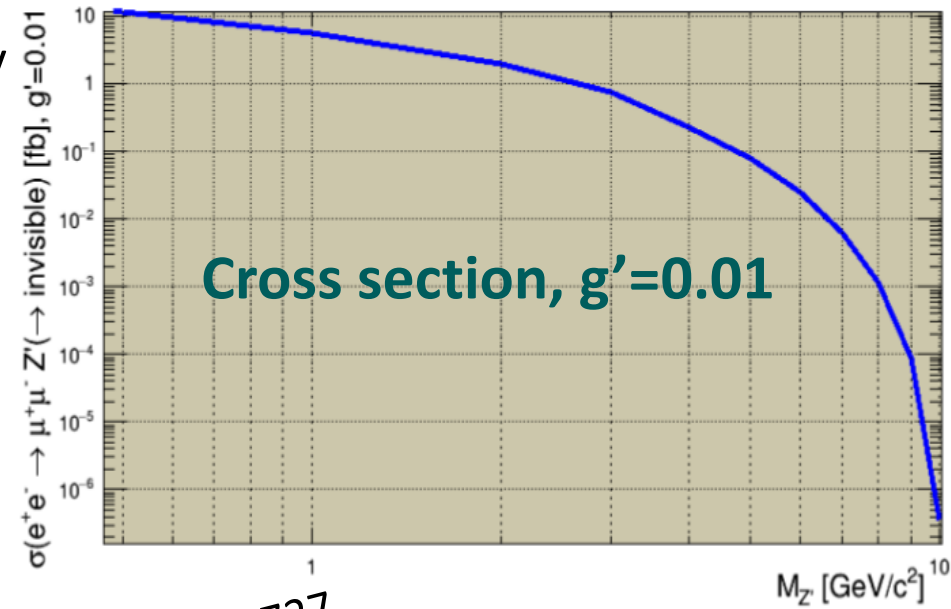
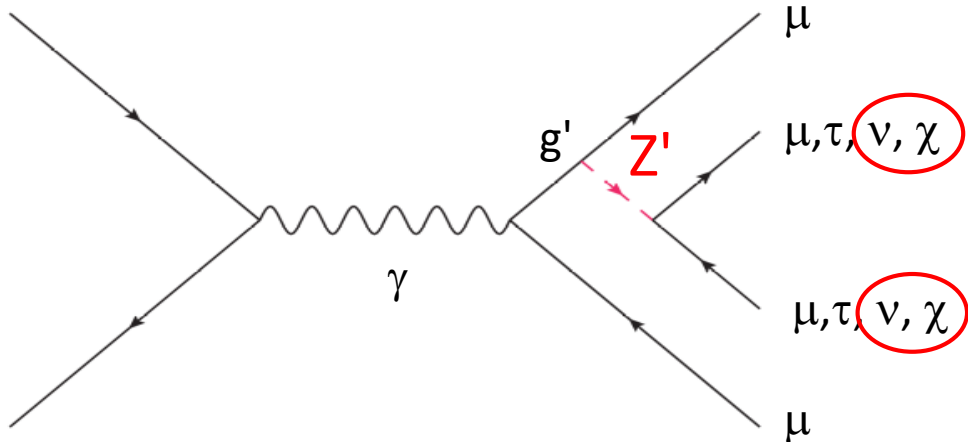


Axion Like Particles (ALPs): sensitivity



$L_\mu - L_\tau$: Z' invisible decay

- A new gauge boson Z' which couples only to the 2^o and 3^o lepton family
- May explain $(g-2)_\mu$
- Invisible decay channel to be explored for the first time
- Invisible decay channel BR possibly enhanced by the presence of kinematically accessible dark matter (e.g. sterile neutrinos)
- Might solve $B \rightarrow K^{(*)} \mu \mu$, R_K , R_{K^*} anomalies
- Sometimes invoked to explain EDGES results



Shuve et al. (2014), arXiv 1408.2727

Invisible Branching Ratios

Branching ratios to SM ν 's:

$$M_{Z'} < 2 M_\mu \rightarrow \Gamma(Z' \rightarrow \text{inv.}) = 1$$

$$2 M_\mu < M_{Z'} < 2 M_\tau \rightarrow \Gamma(Z' \rightarrow \text{inv.}) \sim 1/2$$

$$M_{Z'} > 2 M_\tau \rightarrow \Gamma(Z' \rightarrow \text{inv.}) \sim 1/3$$

If LDMA kinematically available $\rightarrow \approx 1$

$L_\mu - L_\tau, Z'$ invisible decay sensitivity

Look for bumps in recoil mass against a $\mu^+\mu^-$ pair

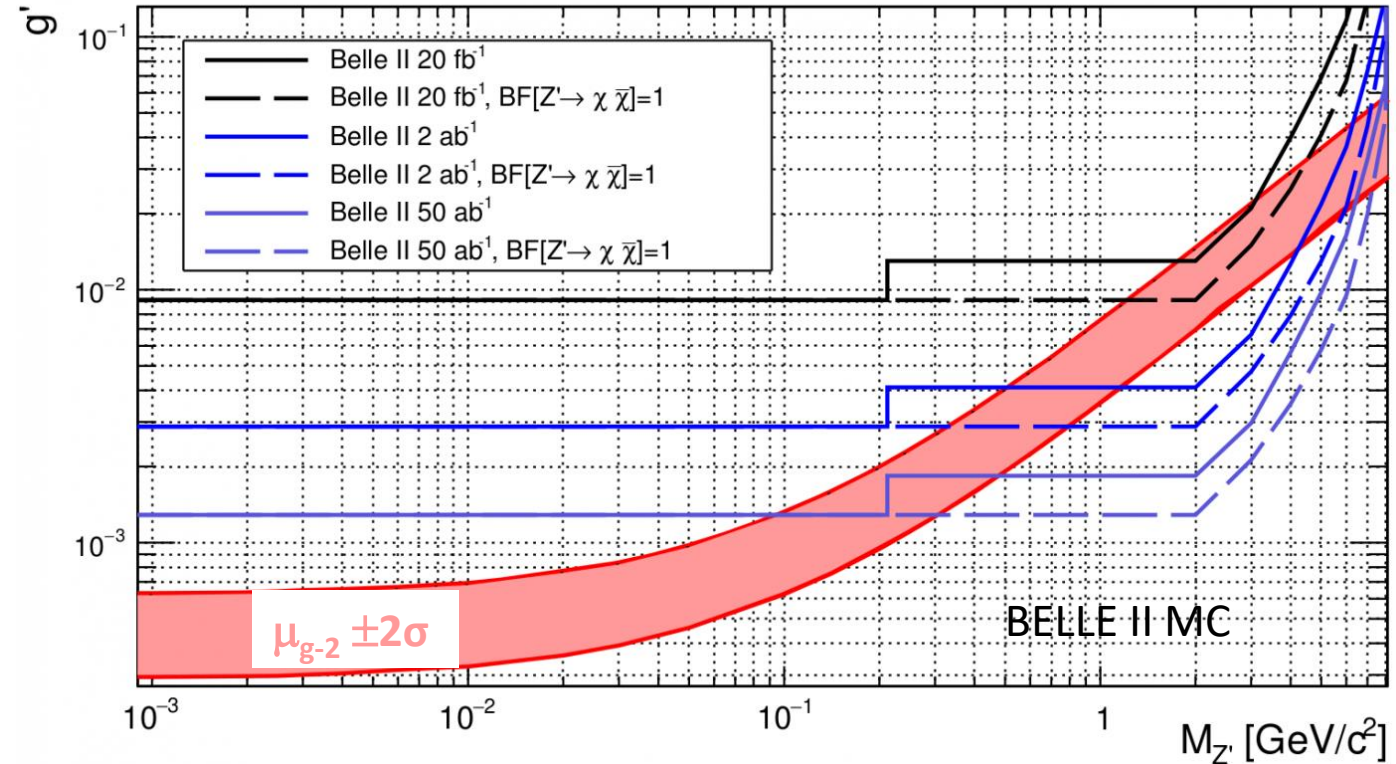
Main backgrounds:

$$e^+e^- \rightarrow \mu^+\mu^- (\gamma)$$

$$e^+e^- \rightarrow \tau^+\tau^- (\gamma), \tau^\pm \rightarrow \mu^\pm\nu\nu$$

$$e^+e^- \rightarrow e^+e^- \mu^+\mu^-$$

Belle II expected sensitivity for $Z' \rightarrow$ invisible



$L_\mu - L_\tau, Z'$ invisible decay sensitivity

Look for bumps in recoil mass against a $\mu^+\mu^-$ pair

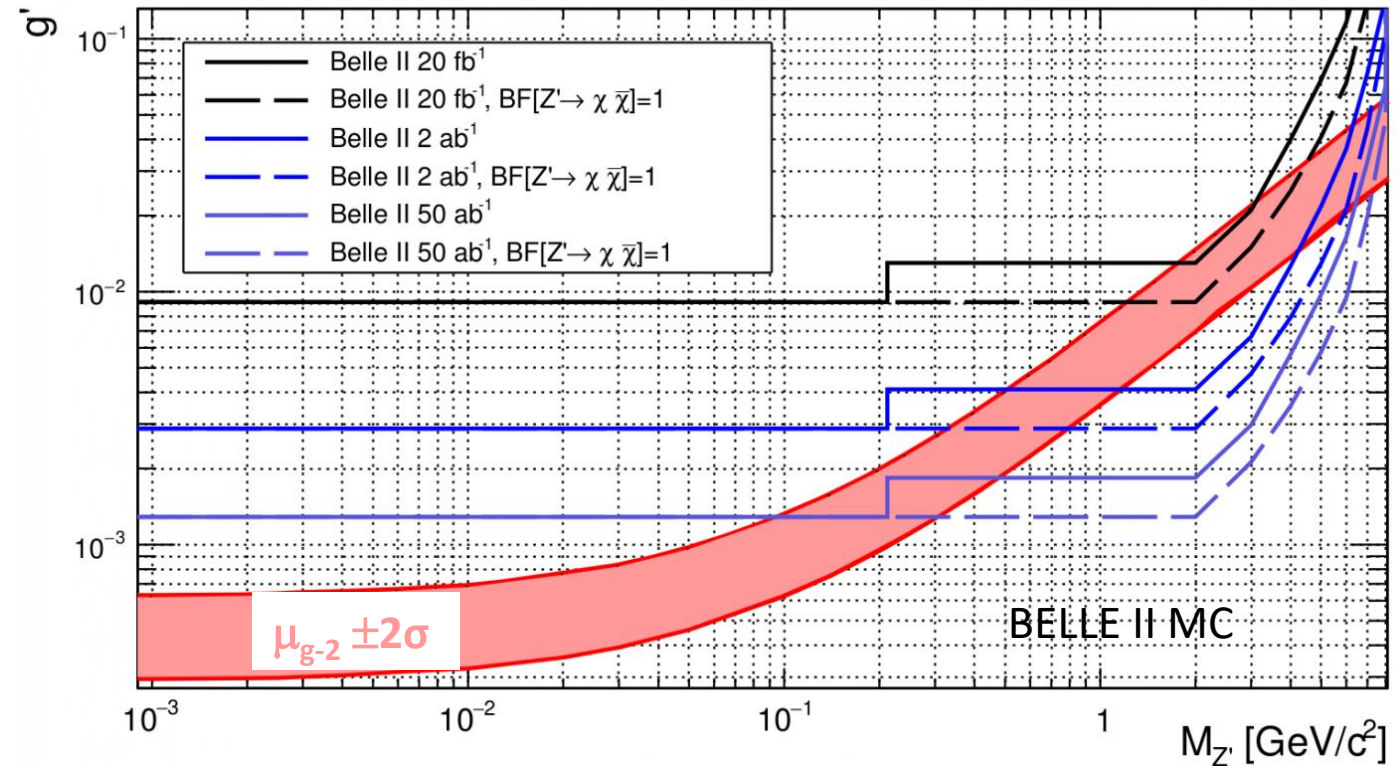
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Belle II expected sensitivity for $Z' \rightarrow$ invisible



$Z' \rightarrow$ visible decay (muonic dark force)

$e^+e^- \rightarrow \mu^+\mu^- Z'$; $Z' \rightarrow \mu^+\mu^-$ will be competitive in Phase 3
(due to BaBar result)

$L_\mu - L_\tau, Z'$ invisible decay sensitivity

Look for bumps in recoil mass against a $\mu^+\mu^-$ pair

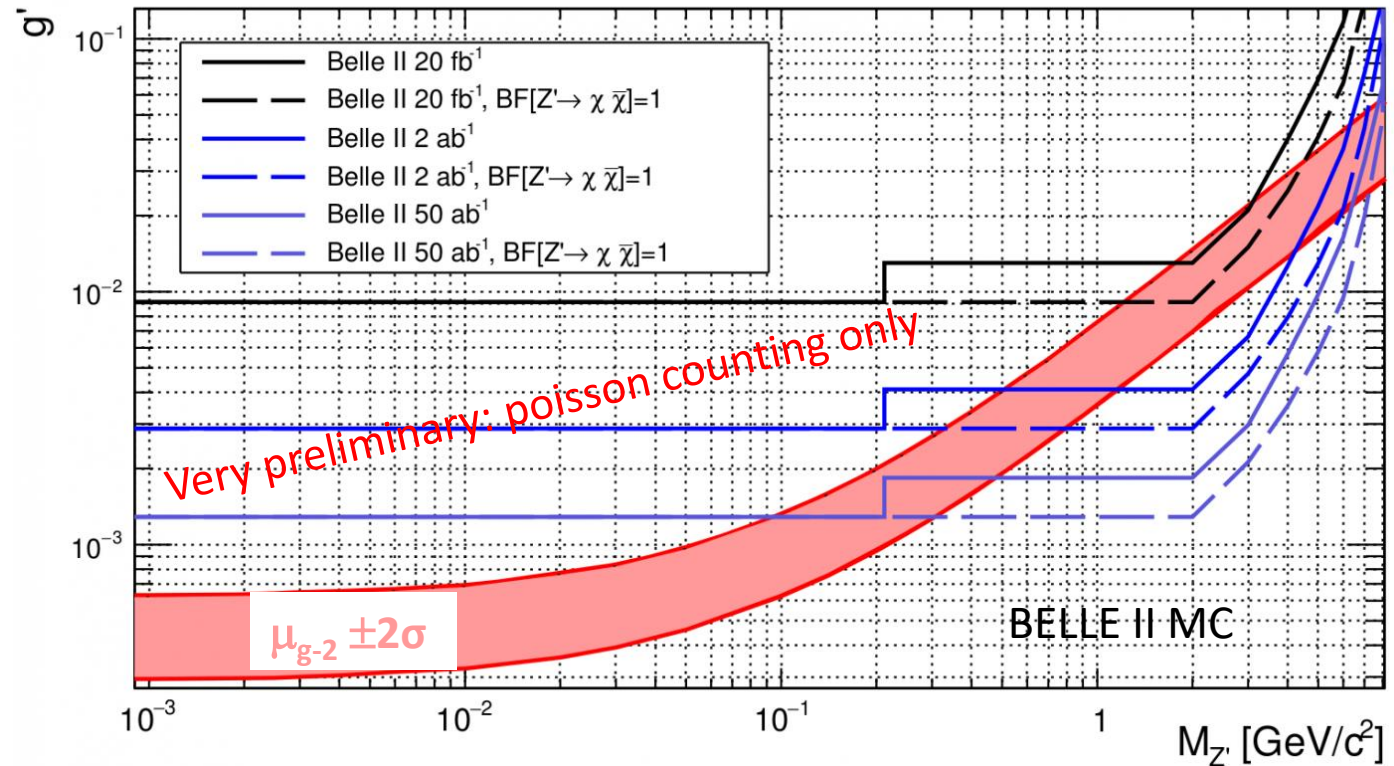
Main backgrounds:

$$e^+e^- \rightarrow \mu^+\mu^- (\gamma)$$

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$$e^+e^- \rightarrow e^+e^- \mu^+\mu^-$$

Belle II expected sensitivity for $Z' \rightarrow$ invisible



Alternative model under search

LFV Z' ($e\mu$ coupling)

$$e^+e^- \rightarrow e^+\mu^- Z' ; Z' \rightarrow \text{invisible}$$

$$e^+e^- \rightarrow e^+\mu^- Z' ; Z' \rightarrow e^+\mu^- \text{ (no SM background expected)}$$

$Z' \rightarrow$ visible decay (muonic dark force)

$e^+e^- \rightarrow \mu^+\mu^- Z' ; Z' \rightarrow \mu^+\mu^-$ will be competitive in Phase 3 (due to BaBar result)

Summary

- Belle II Phase2 finished in July
- Early data taking mostly devoted to commissioning
- $L_{\text{int}} \approx 0.5 \text{ fb}^{-1}$, with $L_{\text{MAX}} = 5.5 \times 10^{33} \text{ cm}^{-2} \text{ s}^{-1}$
- Hardware L1 trigger extensively studied (both tracks and neutrals)
- Resonances, b-physics and charm physics «rediscovered»

Some dark sector searches may lead to interesting new limits even with small data sets

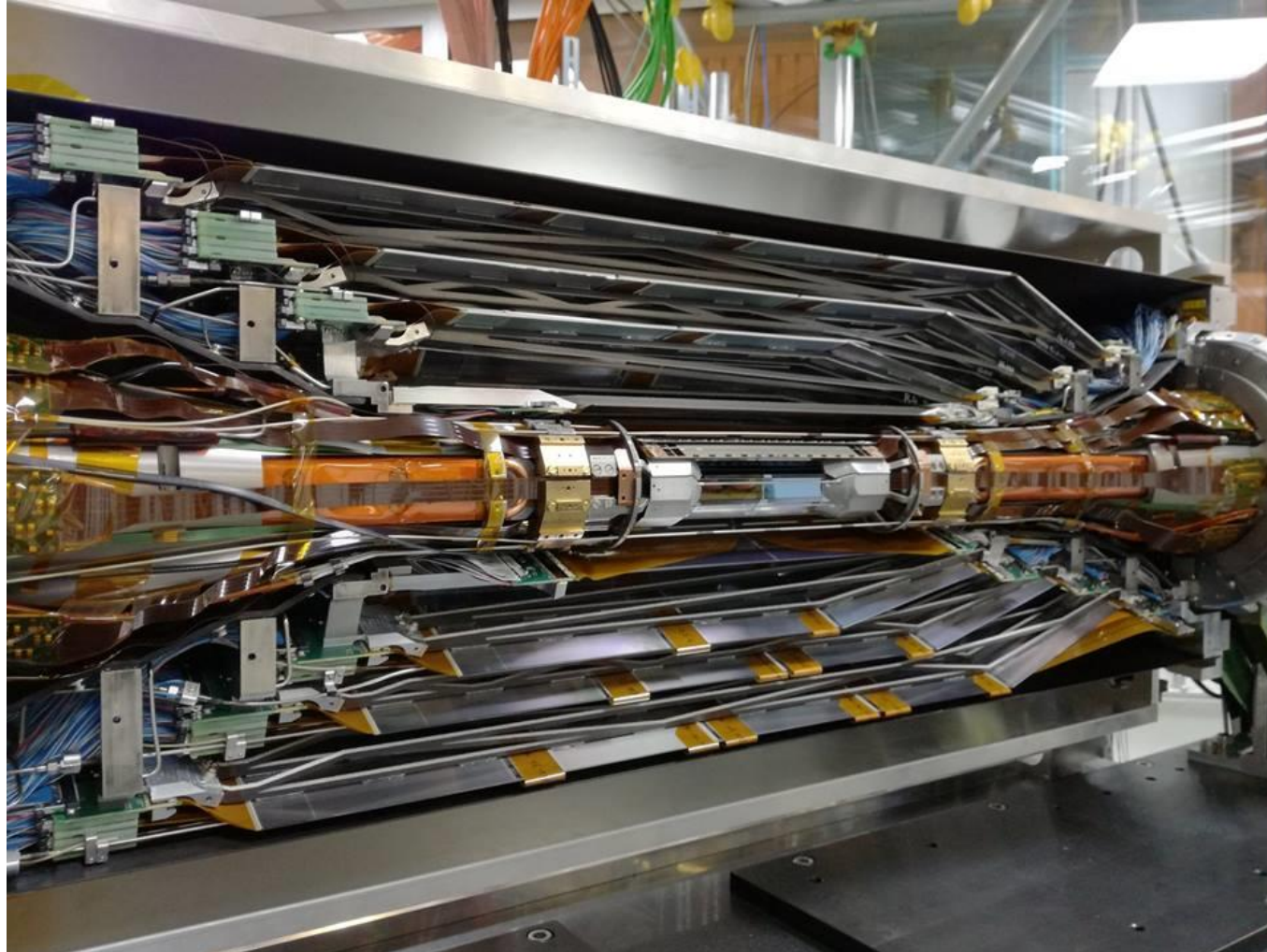
- **Invisible dark photon search**
- **ALP search**
- **Z' to invisible search**
- **Z' LFV search**

Not even mentioned

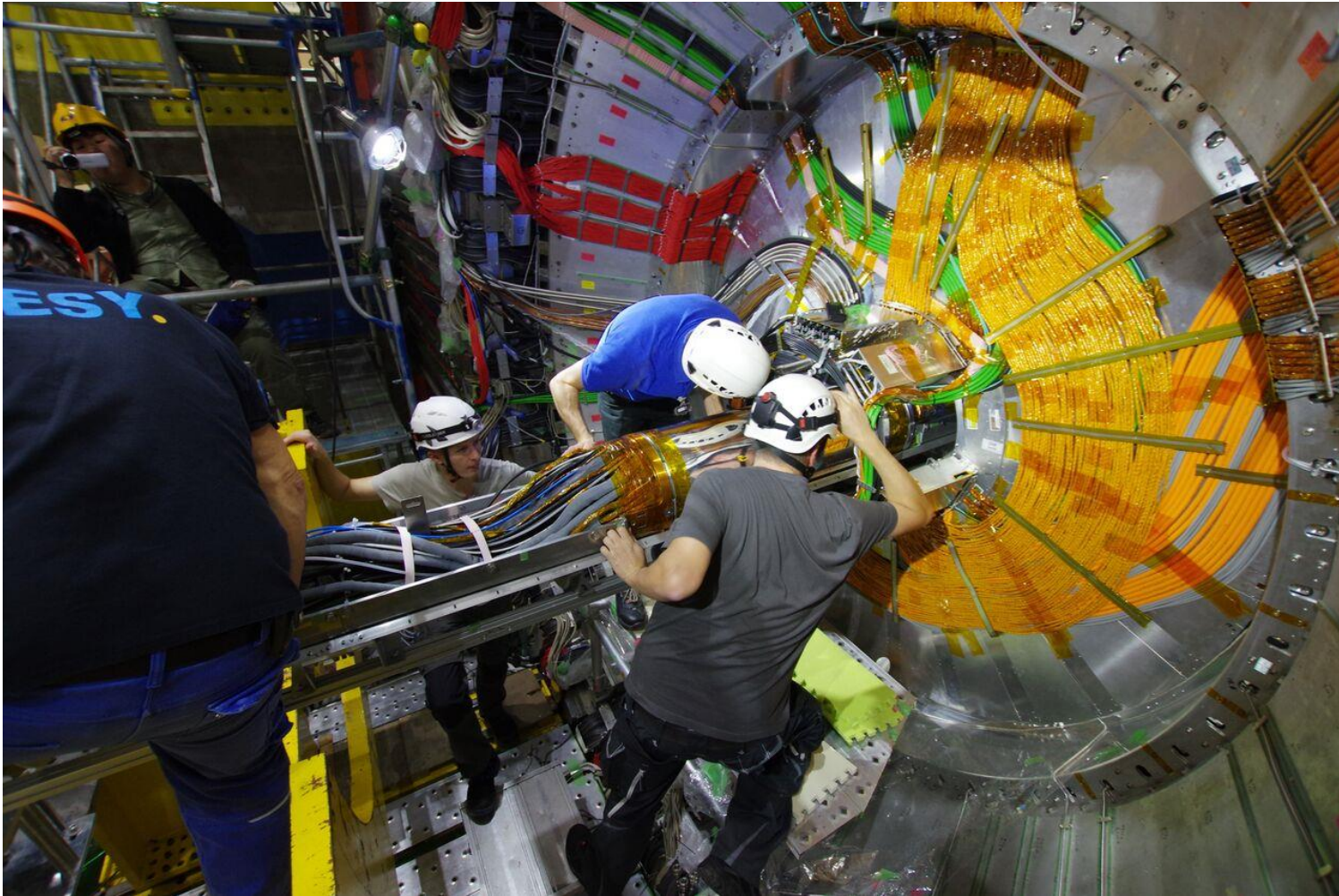
- Magnetic monopoles
- Y(1S) to invisible
- muonic dark force
- dark Higgs
- dark Higgstrahlung
- dark scalars
- inelastic dark matter
- long-lived particles
- ...

All searches in progress, to be finalized soon, aiming at more sensitive results in (the beginning of) Phase 3

Phase 3 (full detector, higher luminosity) will start in Spring 2019



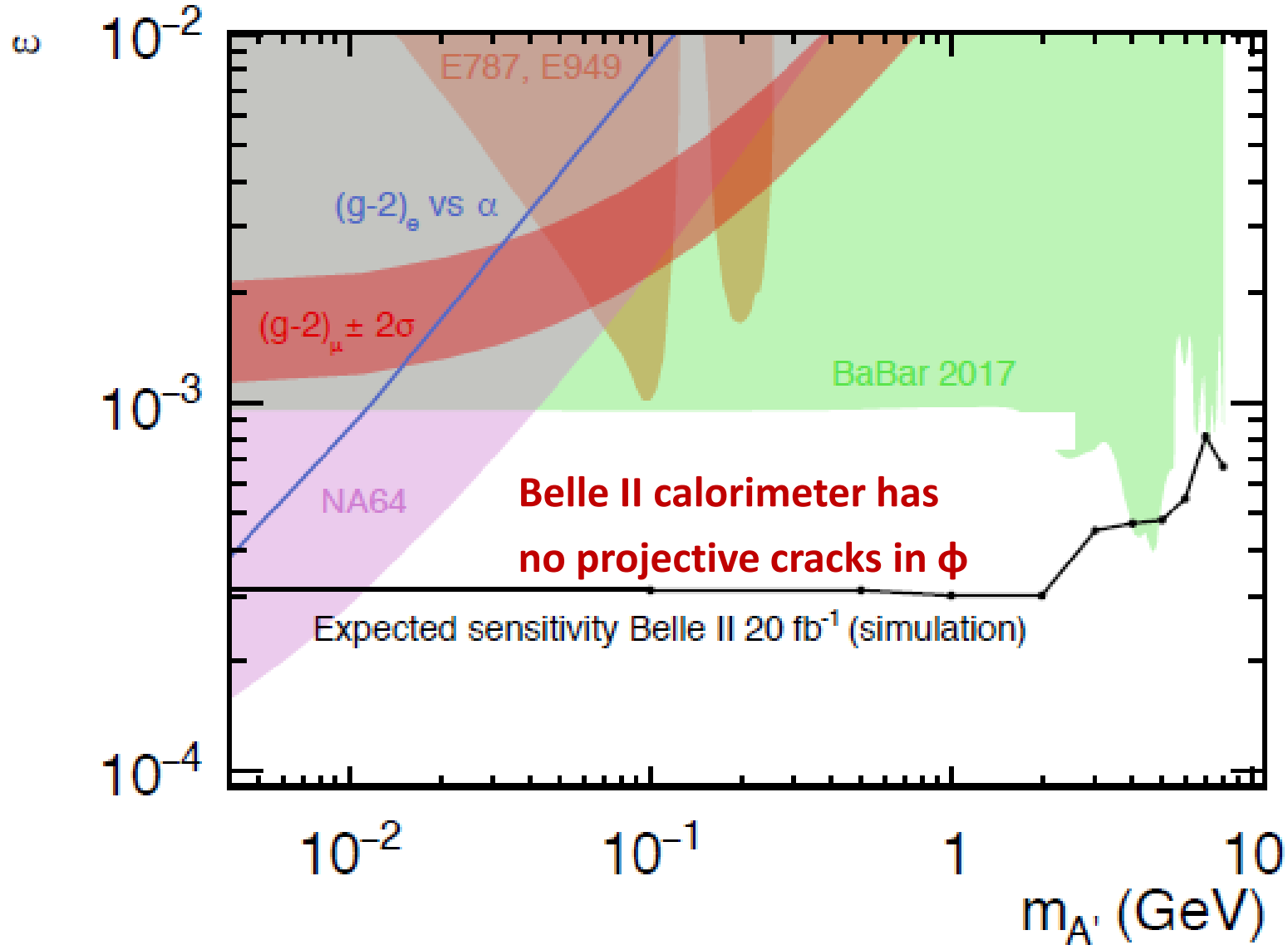
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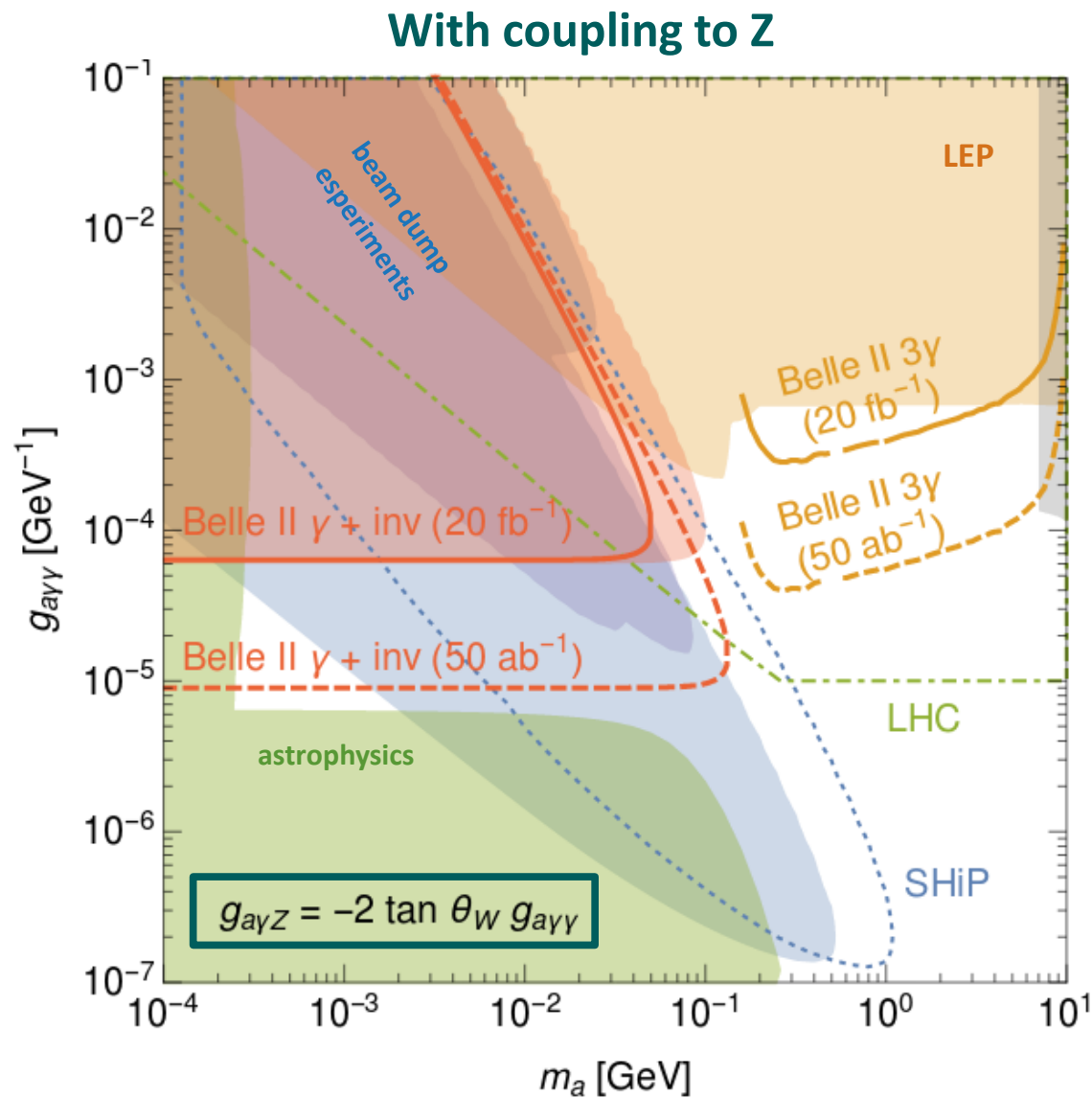
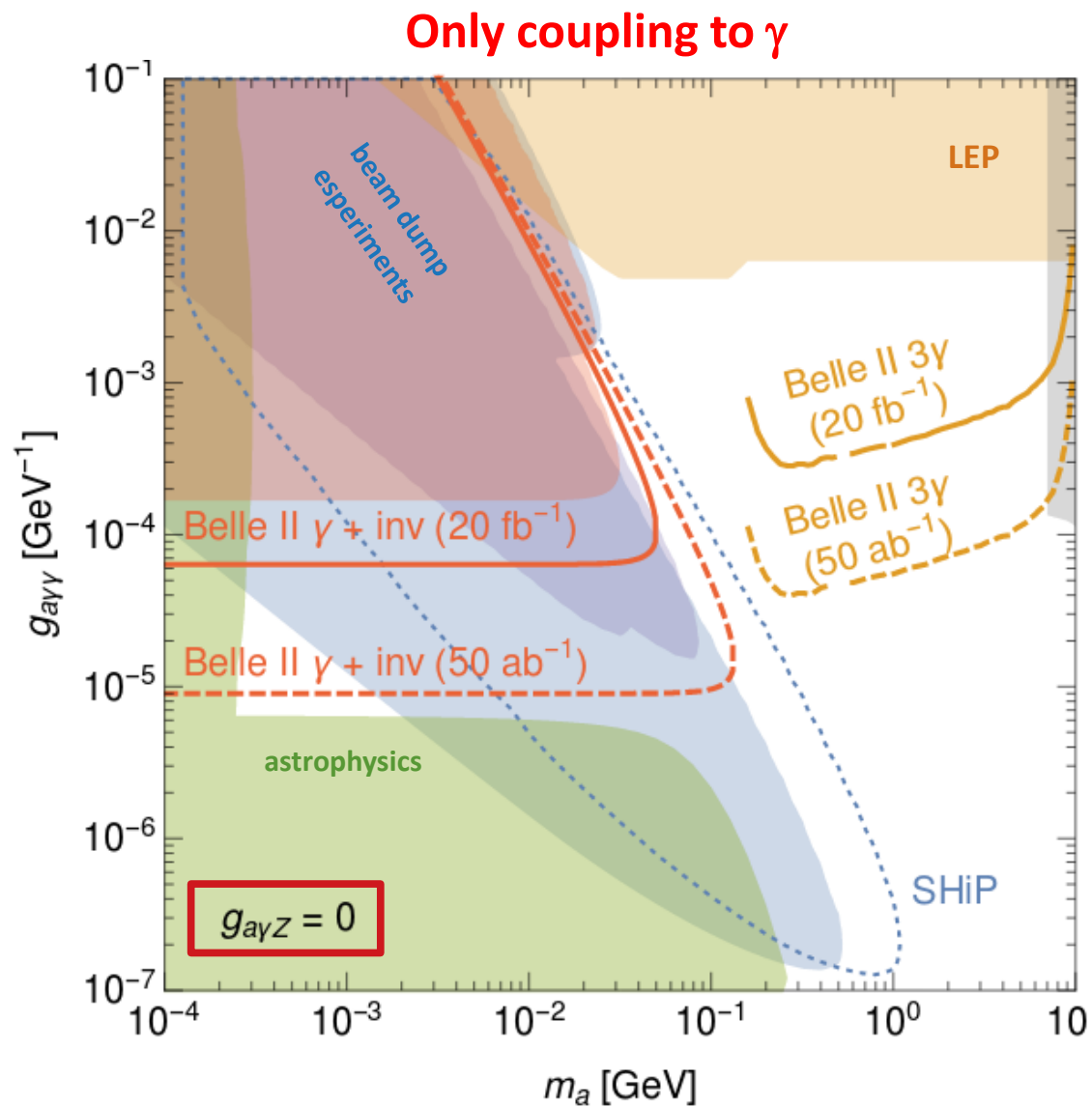
Vertex detector inserted in Belle II two days ago

SPARE SLIDES

Invisible dark photon: sensitivity



Axion Like Particles (ALPs): sensitivity



Z' LFV: invisible + visible

What if symmetries of SM are not kept in the Dark Sector?

What if DM violates Lepton Flavour?

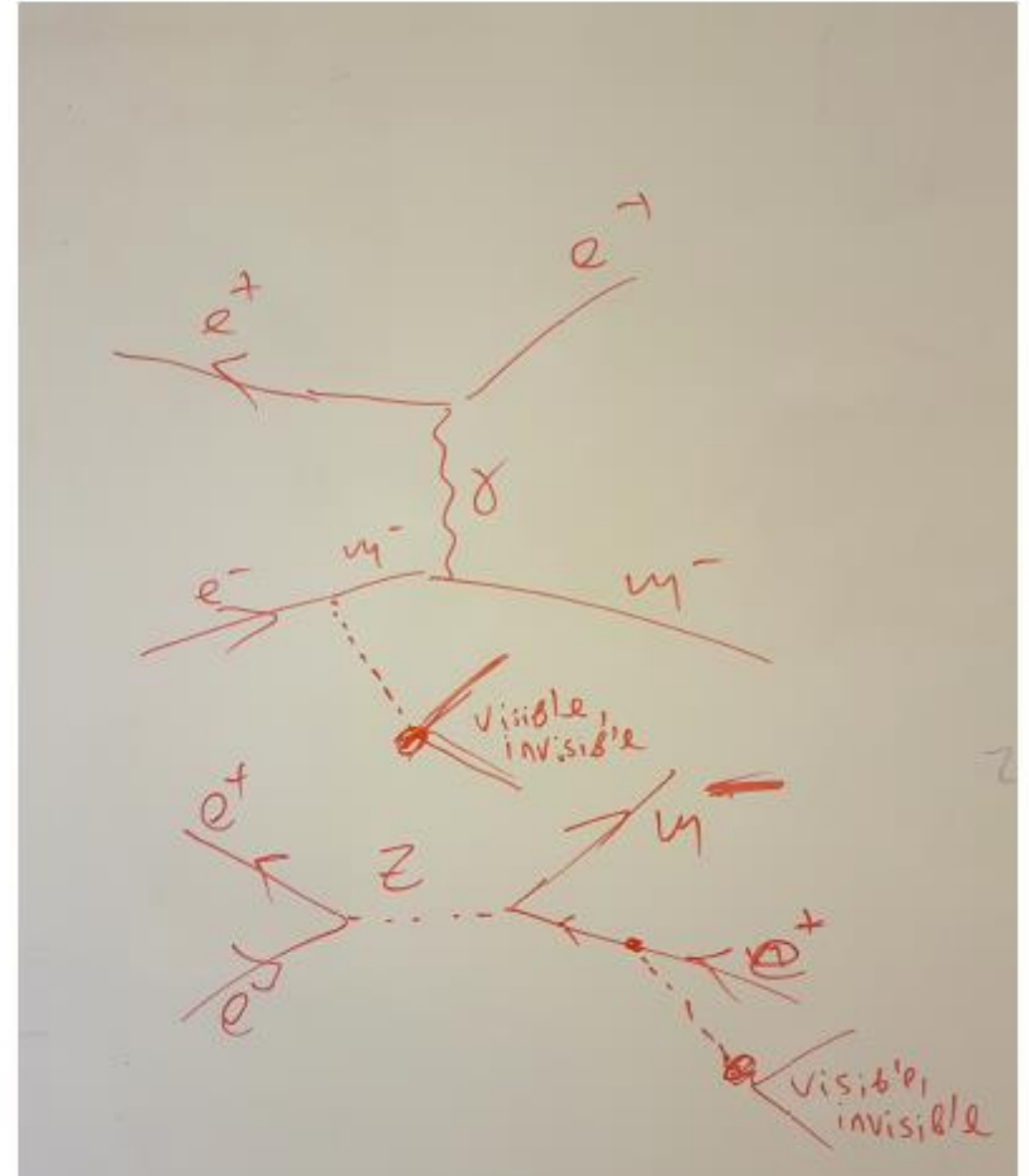
One can imagine, for example, $e\mu$ coupling

$$e^+ e^- \rightarrow e^+ \mu^- Z' ; Z' \rightarrow \text{invisible}$$

Dominant background: $e^+ e^- \rightarrow \tau^+ \tau^- (\gamma)$, $\tau^\pm \rightarrow \mu^\pm, e^\pm \nu\nu$

$$e^+ e^- \rightarrow e^+ \mu^- Z' ; Z' \rightarrow e^+ \mu^- + \text{c.c.}$$

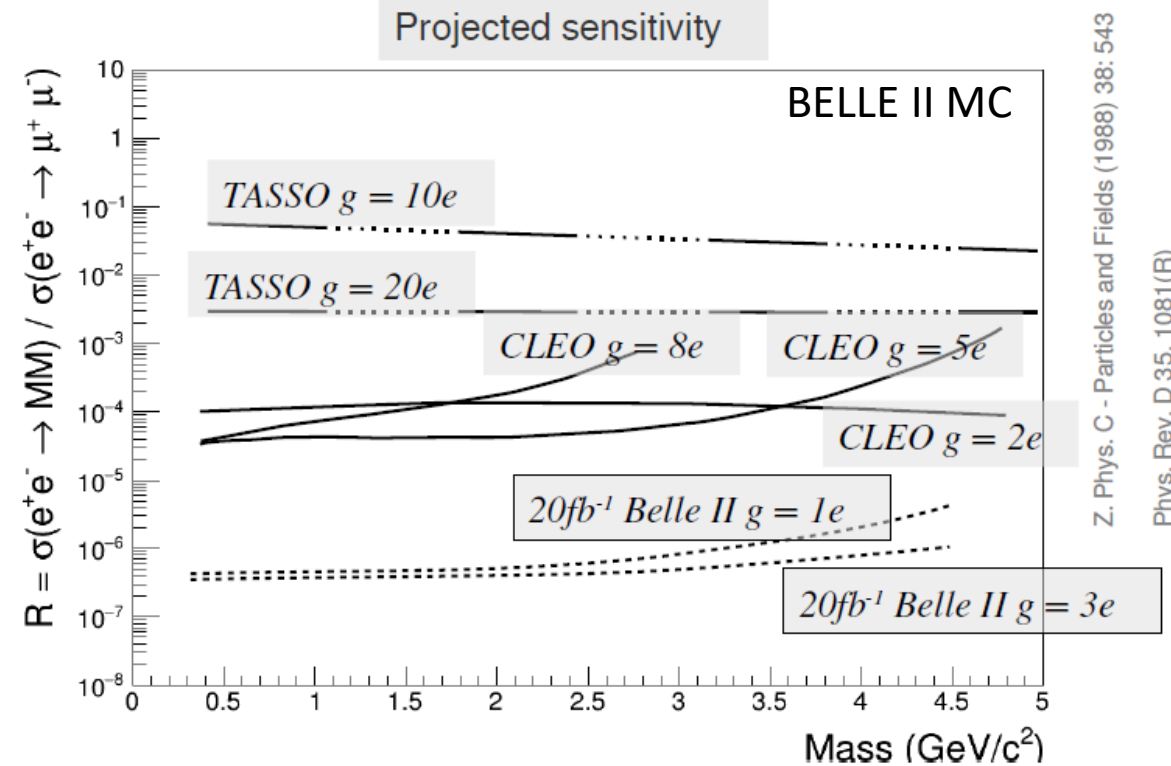
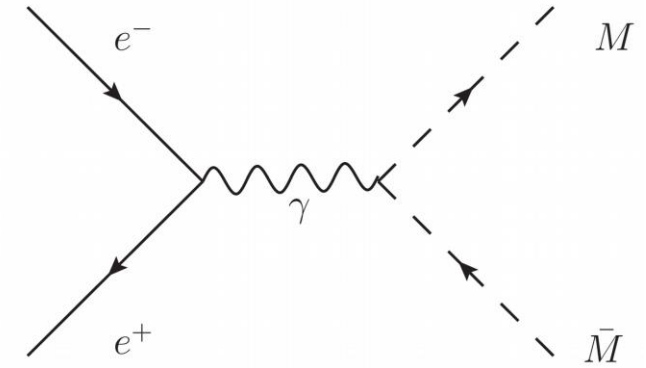
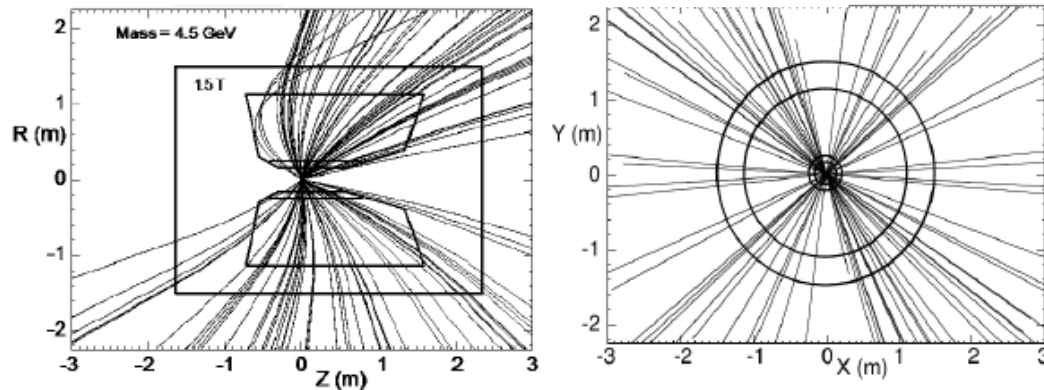
no SM background



Magnetic monopoles

- Particle carrying magnetic charge
- Recent searches for magnetic charges $g > 68.5e$
- Small charges $g < 10e$ are not excluded
- Weaker ionisation due to absence of $1/\beta^2$ factor for magnetic charges
- Tracks are straight in XY and curved in RZ
- They need a dedicated tracking (parabolas rather than helices)

$$z(s) = z_0 + \frac{p_z}{p_T} s + \frac{gBm}{2p_T^2} s^2$$



Z. Phys. C - Particles and Fields (1988) 38: 543
Phys. Rev. D 35, 1081(R)